GEOCHEMICAL AND GEOCHRONOLOGICAL CONSTRAINTS ON THE TECTONOTHERMAL HISTORY OF THE CENTRAL AND EASTERN NEPAL HIMALAYA: IMPLICATIONS FOR THERMAL-MECHANICAL MODELS

by

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The following individuals read and discussed the dissertation submitted by student Stephanie Stacey Starr, and they also evaluated her presentation and response to questions during the final oral examination. They found that the student passed the final oral examination, and that the dissertation was satisfactory for a doctoral degree and ready for any final modifications that they explicitly required.

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ABSTRACT

Syntheses of Himalayan tectonics imply an average shortening rate of ~2 cm/yr across the Himalaya over the last 20-25 Myr. However, despite decades of research, basic questions still remain about how strain was partitioned. Fluctuations in the timing and rate of thrusting could have significant implications for strain partitioning across the Himalaya, and for partitioning during collisional orogenesis as a whole.

Due to the observed chronology and thermobarometry of Greater Himalayan Sequence (GHS) rock units in the Annapurna region of central Nepal, we revise the original interpretation of the Bhanuwa Fault from a normal fault to a thrust, and identify a new thrust called the Sinuwa Fault structurally above the Bhanuwa thrust. The calculated slip rates for the Sinuwa (ST), Bhanuwa (BT), and Main Central (MCT) Thrusts show P-T-t consistency with a 2 cm/yr convergence rate component across the Himalaya from ~25 to ~15 Ma. These results support models that presume constant rates since at least 25 Ma. The similarity in inferred peak ages for the ST, BT, and MCT could in principle indicate simultaneous initial cooling of all three sheets at ~22 Ma. A simple explanation would involve transport and cooling of each rock package along the same basal thrust (the MCT), and later juxtaposition of already-cooled rocks along the ST and BT. That is, this model implicitly requires as much as 100 km of out-of-sequence thrusting. Additional measurements from other parts of the orogen in India and Bhutan are needed to verify whether these estimates could be applied to the orogen as a whole.

Lu-Hf dates from garnet separates in one relict eclogite from the Arun River valley in eastern Nepal indicate an age of 20.7±0.4 Ma, the first direct measurement of the timing of eclogitization in the central Himalaya. Four proximal garnet amphibolites from structurally lower horizons are 14-15 Ma, similar to post-eclogitization ages published for rocks along strike in southern Tibet. P-T calculations indicate three metamorphic episodes for the eclogite: eclogite-facies metamorphism at 23-16 Ma, a peak-T granulite metamorphism, and a late-stage amphibolite-facies metamorphism at \sim 14 Ma. Three models are considered to explain the observed P-T-t evolution. The first assumes that the Main Himalayan Thrust (basal thrust of the Himalayan thrust system) cuts deeper at the Arun River valley than elsewhere. While conceptually the most simple, this model has difficulty explaining both the granulite-facies overprint and the pulse of exhumation between ~25 and 14 Ma. A second model assumes that (aborted) subduction, slab breakoff, and ascent of India's leading edge occurred diachronously: ~50 Ma in the western Himalaya, ~25 Ma in the central Himalaya of Nepal, and presumably later in the eastern Himalaya. This model explains the P-T-t path, particularly heating during initial exhumation, but implies significant along-strike diachroneity, which is generally lacking in other features of the Himalaya. A third model assumes repeated loss of mantle lithosphere, first by slab breakoff at \sim 50 Ma, and again by delamination at \sim 25 Ma; this model explains the P-T-t path, but requires geographically restricted tectonic behavior at Arun. The P-T-t history of the Arun eclogites may imply a change in the physical state of the Himalayan metamorphic wedge at 25-16 Ma, ultimately giving rise to the Main Central Thrust by 16-15 Ma.

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CHAPTER 1. INTRODUCTION

1.1 Introduction to Himalayan Orogen

Stretching over 2400 km from Nanga Parbat in Pakistan eastward across parts of India, Nepal, and Bhutan, to Namche Barwa in Tibet (Figure 1.1), the Himalayan mountain belt is one of the most prominent and striking examples of continent-continent collisional orogenesis. Approximately 400-700 km of shortening have been taken up across the Himalaya since 20-25 Ma (e.g., Yin and Harrison, 2000; DeCelles et al., 2001; Guillot et al., 2003; Robinson et al., 2006), which corresponds well with the modern Himalayan accommodation rate of ~2 cm/yr extrapolated over 20-25 Myr (Bilham et al., 1997; Larson et al., 1999). During this time, sediments initially deposited on the northern continental margin of India were buried, deformed, metamorphosed, and exhumed, with these same processes continuing through to the present. The Himalaya therefore present scientists with an ideal setting in which to test models of collisional orogenesis.

1.2 Description of the Research

Metamorphic rocks are commonly exposed in the Himalaya, and geologists recognized that the relationships between the metamorphic rocks and the fault zones that cut them are critical to understanding the processes that created the orogen (e.g., Gansser, 1964; Le Fort, 1975). Numerous studies have focused on Himalayan metamorphic rocks, particularly the highest-grade rocks exposed in the thrust belt, the Greater Himalayan Sequence (GHS). GHS rocks are exposed along a strip in the geographic interior of the orogen, where they rest structurally above the Lesser Himalayan Sequence (LHS) along the Main Central Thrust (MCT). Most researchers find increasingly high-grade rocks exposed at progressively higher structural levels as one progresses from the Main Frontal Thrust northward toward the South Tibetan Detachment. This pattern of metamorphism is referred to as inverted metamorphism, and explanations for this feature are critical components of tectonic evolution models for Himalayan orogenesis (e.g., Le Fort, 1975; England and Molnar, 1993; Jamieson et al., 1996; Harrison et al., 1999a).

Understanding any evolutionary process is directly related to how well we understand the timing of events. The goal of this study is to contribute to our understanding of the timing and processes involved in the formation of the Himalayan orogen through an examination of the tectono-thermal history of portions of the exhumed crystalline core of the Himalaya (the GHS). Combining metamorphic petrology with geochronology and geochemistry permits the construction of pressure-temperature-time (P-T-t) histories and an estimation of the rates and mechanisms of crustal shortening. This information allows examination of the thermal evolution of the crystalline core of the Himalaya in the context of recent evolutionary models. Key questions addressed by this research include:

1) Has the Indo-Asian convergence rate remained constant at ~ 2 cm/yr, or has it fluctuated significantly over time?

2) Can strain measurements and shortening rate estimates from one location be extrapolated across the orogen, or are they only valid for that region?

3) Was movement along the major faults (specifically, the MCT) in different regions contemporaneous, or was movement stalled in some areas while active in others?

4) Was the Greater Himalayan Sequence in the central Himalaya partially subducted in the Cenozoic?

5) Is the behavior of the crust during Himalayan orogenesis sufficiently explained by the thermal-mechanical model of channel flow (Beaumont et al., 2001, 2004; Jamieson et al., 2002, 2004), which calls for a combination of focused erosion and an extremely weak middle and lower crust, or instead the critical taper model (e.g., Henry et al., 1997), which is based on maintaining a certain wedge geometry?

Addressing these issues is not only important to the development of the Himalaya, but also is relevant to the overall behavior of continental crust during collisional orogenesis.

1.3 Organization of the Dissertation

This dissertation is organized into six chapters. The first provides the reader with an introduction to the Himalayan orogen, and sets out the questions and goals of this dissertation.

Chapter 2 describes the geology of the Himalaya, emphasizing the major tectonostratigraphic units and major fault systems. This chapter also reviews the two major end-member thermal-mechanical models of Himalayan development and how each model can be tested by examining the age and P-T history of the crystalline core of the Himalaya (i.e., the Greater Himalayan Sequence, GHS).

Chapter 3 reviews the methodology used to obtain geochronologic and geochemical data from rock samples collected along two transects through the crystalline Himalaya of central and eastern Nepal. It details operating conditions for *in situ* monazite Th-Pb ion microprobe dating as well as the chemical and instrumental protocols for Lu-Hf garnet dating via multi-collector inductively coupled plasma mass spectrometry (MC-ICP-MS). Additionally, operating procedures for gathering chemical information via the electron microprobe and laser ablation inductively coupled plasma mass spectrometer (LA-ICP-MS) are given.

Chapter 4 presents thermobarometry and monazite geochronology from the first of two transects studied over the course of this project – the Annapurna transect along the Modi Khola in central Nepal. The data are used to suggest 1) that distinct chemical zones in monazite are petrologically linked to geochronology; 2) the presence of three distinct groups of rocks within GHS Formation I; 3) the presence of an unrecognized fault within Formation I that we refer to as the Sinuwa Thrust; and 4) that strain measurements in one part of an orogen can be realistically extrapolated to another part within a few hundred kilometers.

Chapter 5 presents thermobarometry and garnet geochronology from the second transect – the Arun River valley in eastern Nepal. These data represent the first attempts to directly date high-pressure eclogitization in the central Himalaya, although several attempts have been made to date the over-printing granulite metamorphism in the region. The age and evolution of eclogites in this study contrast markedly with the ultra-high pressure eclogites from the northwest Himalaya in a number of ways. The Arun rocks were metamorphosed at much lower pressures, show significant heating during initial exhumation, and were metamorphosed 20-30 Myr later, potentially signaling different tectonic processes between the two regions.

Chapter 6 is the conclusion of the dissertation. The major contributions of this research are summarized and the potential for new work is discussed.



Purple squares are major cities, green triangles are the major peaks representing the syntaxes, yellow Figure 1.1. Satellite image of the Himalayan orogen, showing the expanse of the Himalayan Range. boxes represent field areas for this study.

CHAPTER 2. GEOLOGY OF THE HIMALAYA

Prior to Indo-Asian collision, the northern Indian plate was probably composed of a thinned continental margin covered by Proterozoic clastic deposits and a Cambrian-Eocene shelf sequence (Le Fort, 1996). At present, the collisional mountain belt features roughly parallel, south-directed crustal-scale fault systems that bound distinctive lithologic units along strike.

2.1 Major Tectonic Subdivisions and Structures

2.1.1 Geology of the Major Tectonostratigraphic Units

In general, laterally extensive, crystalline thrust sheets of the Nepalese Himalaya can be grouped into two main lithotectonic units: the Greater Himalayan Sequence (GHS) and the Lesser Himalayan Sequence (LHS), which are bounded above by the South Tibetan Detachment System (STDS), and below by the Main Boundary Thrust (MBT) (Figure 2.1).

Greater Himalayan Sequence (GHS)

Also called the High Himalayan crystallines, the high-grade metasediments of the GHS form the metamorphic core of the Himalayan orogen. Throughout much of Nepal, the crystalline rocks of the GHS are commonly grouped into three formations: I, II, and III (LeFort, 1975; Colchen et al., 1986; Pêcher and Le Fort, 1986). In central Nepal, the

lowest structural unit of the GHS, Formation I, consists of migmatitic upper amphibolite-facies two-mica schist and gneiss of generally pelitic composition. Formation II is composed of alternating beds of pyroxene and amphibole-bearing calcsilicate, marble, and quartzite. Formation III is a thin unit of pelitic schist and an augen gneiss. In this study, we further divide GHS Formation I into three subgroups: 1a, 1b, and 1c based on petrology. Formation 1a is muscovite-rich and contains garnet that exhibits growth zoning and/or oscillations of the major elements. Formation 1b rocks are locally migmatic and the garnets have homogenous zoning. Formation 1c rocks are migmatitic with segregated leucosomes and zoning in garnet that shows thick resorption rims following diffusional homogenization.

In eastern Nepal, the convention of dividing the GHS into 3 units is maintained, though the names and mineralogies differ (Lombardo et al., 1993; Pognante and Benna, 1993; Carosi et al., 1999; Goscombe and Hand, 2000). Comparable to Formation I in central Nepal, the Barun Gneiss also consists of migmatized pelitic schist and gneiss, but is additionally interbedded with intermittent metabasites and calc-silicates. The middle unit, the Namche Migmatite orthogneiss, is composed of sillimanite-bearing granitic orthogneisses. The upper unit in the GHS of eastern Nepal is the Black Gneiss, a high-grade pelitic paragneiss with interbedded quartzite. The uppermost sections of the GHS typically contain abundant leucogranite intrusions that are the product of early Miocene crustal anatexis from dehydration melting of the GHS (Harris and Massey, 1994; Harrison et al., 1999a, 1999b; Kohn, 2008). The depositional age for the sedimentary protolith of the GHS is estimated at 600-500 Ma based on detrital zircon ages and c. 500 Ma felsic intrusions (DeCelles et al., 2000; Gehrels et al., 2003). Some researchers have

suggested a two-stage metamorphic history of the GHS, with Neogene metamorphism preceded by an Eocene ("Eohimalayan") metamorphic and anatectic(?) event in the GHS related to nappe burial and crustal thickening (Pognante and Benna, 1993; Vannay and Hodges, 1996; Hodges et al., 1996; Coleman, 1998; Godin et al., 2001). Chronologic evidence for an Eohimalayan event comes from monazite grains containing an inherited component of Oligocene-Eocene age (Hodges et al., 1996; Edwards and Harrison, 1997; Coleman, 1998; Coleman and Hodges, 1998). However, without the benefit of direct correlation between the ages and chemical zonation, it is uncertain whether these monazites formed due to early solid-state metamorphic reactions, as in the Langtang region of Nepal, ~200 km to the east (Kohn et al., 2004, 2005), or as a result of partial melting, as was inferred by Godin et al. (2001).

Lesser Himalayan Sequence (LHS)

The LHS is often split into two sub-units: lower Lesser Himalayan (LLHS) and upper Lesser Himalayan (ULHS). The LLHS consists of the weakly metamorphosed metasediments of the Kuncha Formation, typically chlorite schists and quartzites, sporadically interbedded with orthogneisses, felsic metavolcanics, and metabasalts. The ULHS comprises greenschist- to amphibolite-facies schists, marbles, and calc-silicates of the Dhading (dolomitic marble), Benighat (carbonaceous pelite), and Malekhu (marble and calcic schist) Formations. Older and younger LHS rocks are separated by a major unconformity contemporaneous with Pan-African tectonics (Valdiya, 1995). The lower, older formations are interpreted as proximal to a Paleoproterozoic (c. 1830 Ma) continental arc that bordered the northern margin of the Indian plate, the leading edge of which is now buried beneath the Himalaya and Tibet (Kohn et al., 2010). Permian to Paleocene sedimentary rocks of Gondwanan affinity overlie the Dhading-Benighat-Malekhu sequence (Sakai, 1983).

In general, the metamorphic grade within the LHS increases structurally upward towards the MCT, i.e., an inverted metamorphic field gradient (Pêcher, 1989). Metamorphic grade in LHS rocks ranges from chlorite- to garnet-grade toward the MCT, and in some areas, the upper part of the unit reached kyanite-grade conditions. Metamorphic grade in the LHS is seemingly correlated with structural position below the MCT. A number of studies have endeavored to explain this phenomenon, as it is a key feature in tectonic reconstructions of the Himalaya (more in Section 2.1.3).

2.1.2 Major Fault Systems

Within the GHS and LHS are several north-dipping faults with various displacement amounts, timing, and rates. Himalayan faults have typically progressed in sequence, so from highest structural position to lowest, the major faults of the crystalline Himalaya are:

South Tibetan Detachment System (STDS)

This fault is an extensional feature rather than a thrust, and its development is important to the tectonic evolution of the Himalayan metamorphic core. The STDS juxtaposes low-grade Tethyan metasediments against the high-grade migmatites of the GHS, and it is now generally believed that movement of the MCT and STDS was synchronous at \leq 16-17 Ma (see Edwards and Harrison, 1997; Hodges et al., 1998; Harrison et al., 1998 for age of STDS; Kohn et al., 2004 regarding timing of main movement on the MCT).

Main Central Thrust (MCT)

Generally speaking, the MCT is the fault along which the GHS was thrust southward over the LHS. Despite the fact that the MCT is a fundamental feature in tectonic reconstructions of the Himalaya (e.g., Yin, 2006 and references therein), some contention exists regarding the characteristics and location of the boundary, stemming from dispute over the various criteria used to identify the MCT. The crustal architecture of the orogen is made difficult in some regions by the presence of at least two major faults, one or both of which have been mapped as the MCT (or MCT-I, MCT-II, upper MCT, lower MCT, etc.) at several different structural and stratigraphic levels (see review of Yin, 2006). The MCT has been mapped: 1) as a metamorphic contact by following the kyanite isograd (e.g., Bordet, 1961; Le Fort, 1975; Colchen et al., 1986); 2) as a lithological contrast between a distinctive quartzite beneath an orthogneiss (Daniel et al., 2003); 3) by differences in Sr and/or Nd isotope compositions (Ahmad et al., 2000; Robinson et al., 2001; Martin et al., 2005; Richards et al., 2005, 2006); 4) by differences in U-Pb detrital zircon ages (Parrish and Hodges, 1996; DeCelles et al., 2000); 5) by discrete Th-Pb monazite ages (Harrison et al., 1997; Catlos et al., 2001, 2002; Kohn et al., 2004); 6) using strain indicators to find the maxima of a shear gradient (Martin et al., 2005; Searle et al., 2008); 7) as an entire zone of high ductile strain up to 10 km thick that is bounded above and below by related thrusts (e.g., Catlos et al., 2001; Grujic et al., 2002). There are difficulties in using any one of these methods independently to define

the MCT. General lithology, detrital zircons, and isotopic analysis provide information on stratigraphy, and mineral isograds and monazite ages provide information about metamorphic reactions. As a structural feature, the MCT should be defined to a first order by structural methods, i.e., strain analysis (Martin et al., 2005; Searle et al., 2008). However, due to variable rheological responses from different rock types, potential preexisting stratigraphic and tectonic structures, and possible migration over time of the main locus of thrusting on the MCT, strain analysis alone does not present a clear definition of what, or where, the MCT should be. Additionally, a lack of an obvious break in metamorphic grade between the LHS and GHS across the MCT makes it difficult to distinguish in the field. For this study, the MCT is assigned to the GHS-LHS contact, defined on a combination of lithologic, isotopic, and geochronologic criteria.

Chronologically speaking, initial movement on the thrust was generally viewed as being ~20-22 Ma based on hornblende 40 Ar/ 39 Ar and monazite U-Pb ages (Hubbard and Harrison, 1989; Hodges et al., 1996; Johnson et al., 2001). However, more recent work has suggested initial movement on the MCT in central Nepal and Bhutan at 16±1 Ma and crystallization of *in situ* melts by 13 Ma (Daniel et al., 2003; Kohn et al., 2004, 2005). The displacement on the thrust is estimated to be ~125 km (DeCelles et al., 2001, 2002; Robinson et al., 2003, 2006; Kohn et al., 2004), with a displacement rate of about 2 cm/yr between 16 and 10.5 Ma (Kohn et al., 2004).

Munsiari Thrust (MT)

In central Nepal, the MT (also referred to as the Ramgarh Thrust: DeCelles et al., 2001) sits just below the MCT and emplaces stratigraphically lower LHS atop

stratigraphically higher LHS (Robinson et al., 2003; Kohn et al., 2004; Pearson and DeCelles, 2005). It was first active in central Nepal at 10.5±0.5 Ma (Kohn et al., 2004). Displacement on the thrust is estimated at ~120 km in western Nepal and ~110 km in central Nepal (DeCelles et al., 2001; Robinson et al., 2003, 2006; Kohn et al., 2004).

Lesser Himalayan Duplex (LHD)

In central Nepal, the LHD is not well exposed; however, metamorphic ages that have been collected within the duplex range from ~8 Ma to 3 Ma (Catlos et al., 2001, 2002; Kohn et al., 2004). The geometry of some of the P-T paths assembled for LHS rocks suggest that there is at least one thrust between the MT and the MBT, providing petrologic support for a duplex (Catlos et al., 2001; Kohn et al., 2001). In western Nepal, the LHD accommodated ~300 km of shortening (Robinson et al., 2006), but in central Nepal, displacement is more difficult to estimate due to the strong influence of erosion on cooling; in this region, exhumation and cooling rates have been calculated at 6-7 mm/yr and >150°C/Myr, respectively (Lavé and Avouac, 2001; Kohn et al., 2004), comparable to erosion rates at both the eastern and western syntaxes (Zeitler et al., 2001; Stewart et al., 2004).

Main Boundary Thrust (MBT)

The MBT places the lowest-grade crystalline rocks of the LHS on top of the unmetamorphosed Siwalik Group syntectonic clastic rocks to the south. Initial movement on the fault has been placed as early as ~5 Ma (DeCelles et al., 1998).

Main Himalayan Thrust (MHT)

The MHT is the main decollment beneath the Himalaya. All of the other major thrusts in the Himalaya are believed to sole into the MHT (Zhao et al., 1993; Nelson et al., 1996), implying that the main faults within the GHS are abandoned and exhumed former MHT surfaces (Kohn, 2008).

2.1.3 Inverted Metamorphic Field Gradient

The most distinctive feature of Himalayan metamorphism is its inverted metamorphic field gradient, which forms a gradual progression from chlorite-grade at the base of the LHS to sillimanite-grade toward the top of the GHS. This does not necessarily represent an inverted geotherm, however. Inverted metamorphism has been associated with regions that have experienced widespread thrust faulting (Ernst, 1973; Graham and England, 1976; Spear et al., 1995). Models of inverted metamorphism in central Nepal include pre- to syn-metamorphic thrusting and then thermal relaxation (Le Fort, 1975; Hubbard and Harrison, 1989; Molnar and England, 1990; Macfarlane, 1995), syn- to post-metamorphic thrusting (Harrison et al., 1998; Huerta et al., 1998), or postmetamorphic thrusting (Inger and Harris, 1992; Hubbard, 1996). Age distributions and P-T histories of rocks from central Nepal seem consistent with a thermotectonic model in which the inverted metamorphic sequences formed during continuous syn-metamorphic to post-metamorphic thrust-sense shearing which caused heating of footwall rocks during transport of the hot hanging wall (Harrison et al., 1998; Catlos et al., 2001; Kohn et al., 2001, 2004; Kohn, 2008).

2.1.4 Magmatism

Many geochronological studies of the Himalaya have concentrated on leucogranites from Nepal and the south Tibet section of the Himalayas, and these studies show that the most important period of leucogranite production occurred ~20-24 Ma (e.g., Schärer, 1984; Harrison et al., 1999b). The most intensively studied pluton in the Himalaya is the Manaslu leucogranite. The pluton intrudes high-grade rocks at the top of the Greater Himalayan Sequence, but does not cross-cut the South Tibetan Detachment (Searle and Godin, 2003). The Manaslu leucogranite is located ~100 km east-northeast of Annapurna, making it the closest large leucogranite body to the Annapurna region. Ion microprobe monazite data indicate that the Manaslu intrusive complex was constructed during pulses of magmatism at 22.9 ± 0.6 Ma and 19.3 ± 0.3 Ma, with the second phase representing approximately one-third of the overall leucogranite exposure (Harrison et al., 1999b). At Annapurna, Hodges et al. (1996) reported ~22.5 Ma deformed kyanitegrade migmatitic segregations in Formation I above the Main Central thrust, and zircons from leucogranite dikes crosscutting Formation II calc-silicates that yield a lower intercept age of 21.5 Ma.

Leucogranites in the Rongbuk valley, on the north side of the Everest Massif and adjacent to the Qomolangma detachment, a segment of the South Tibetan detachment system, are the nearest dated igneous rocks to the Arun River valley. The leucogranites are exposed in the footwall of the detachment, dated at 16-17 Ma, and do not crosscut the fault system (Murphy and Harrison, 1999).

Overall, the ages of these high-level igneous rocks overlap inferred ages of thrusting from Annapurna, and the geochemistry of the leucogranites is broadly consistent with derivation from a metasedimentary source (Le Fort et al., 1987; Guillot and Le Fort, 1995). There is no specific structural or geochemical link between low-level migmatites exposed in Formation 1b-1c and high-level plutons. For example, partial melts within Formation 1b-1c at Annapurna are segregated on scales of centimeters to decimeters, and lack dikes or other evidence of larger-scale mobility (Hodges et al., 1996; Martin et al., 2005).

2.2 Models for the Origin and Evolution of the Himalayan Orogen

Numerous models have been proposed to explain the evolution of the Himalayan orogen, and, in conjunction, its major fault systems (i.e., STDS, MCT, MHT), since the apparent inverted metamorphism observed in the Himalayas must be indicative of whatever tectonothermal process is at work. The two most popular models for the development of the Himalayan orogen are channel flow (Beaumont et al., 2001, 2004; Jamieson et al., 2002, 2004) and critical taper (e.g., Henry et al., 1997).

2.2.1 Channel Flow Model (CFM)

The channel flow model of the Himalaya (Beaumont et al., 2001, 2004; Jamieson et al., 2002, 2004) requires a large partial melt channel that dominates the tectonics of the orogen by coupling with the erosional front (Figure 2.2). Over time and distance (10's of Myr and up to 100's of km), the channel moves toward the orogenic front, advecting heat and bringing partially molten rocks close to the surface. Flow in the channel requires an inclined thrust surface with normal shearing between the GHS and Tethyan Himalaya, and is particularly sensitive to lateral pressure gradients and viscosity. The channel flow

model is compatible with the development of an inverted metamorphic sequence, and predicts a systematic trend in P–T–t histories across the presumed channel (Jamieson et al., 2004). For GHS rocks, P-T paths are predicted to show isothermal exhumation until rocks begin cooling relatively close to the surface (ca. 10 km depth at ca. 10 Ma) at a rate of ~150 °C/Myr.

2.2.2 Critical Taper Model (CTM)

The critical taper model (Henry et al., 1997; Bollinger et al., 2006) assumes that the Himalaya is a wedge of material between the underthrust Indian plate and the overriding Asian plate. In order to maintain its steady-state geometry, the wedge will deform internally, typically through in-sequence, flat thrusting, with thrusts soling to a master decollment; however, it does not discount the possibility of out-of-sequence thrusting that may occur should the wedge become sub-critical. Thrust packages are successively abandoned as thrusting (and deformation) propagates forward in response to underplating of new material from the Indian subcontinent. In contrast to the CFM, normal faulting at the top of the structural package (i.e., the STDS) is not a primary feature, though in the CTM, shear could have involved normal faulting if the wedge became supercritical. Another important difference between the CTM and CFM is the dominance of flow in the weak mid-crust. In critical taper, the effect of a partially molten channel is small; while the model does not rule out partial melts at depth, those melts are not transported forward to the orogenic front by flow. This would be reflected in an older crystallization age than one would expect from the CFM (ca. 20 Ma rather than ca. 10

Ma), and P-T paths for GHS rocks would show initial isobaric cooling far from the surface at a rate of \sim 50 °C/Myr.

2.3 Discussion

There is a distinct lateral consistency of the first-order lithotectonic units along strike in the Himalaya (Figure 2.1). However, there are some smaller-scale dissimilarities between one transect and another, which may provide new evidence for the tectonic processes at work. Regional-scale models for the Himalayan orogen need to be able to account for local variations observed in different transects in order to provide a robust picture of Himayalan evolutionary development. When evaluating the validity of different tectonic models in this context, one may ask what observations might distinguish among competing models. Since all researchers agree that the GHS was partially molten in the middle and lower crust at one point (i.e., a weak channel), one may further ask how strongly the channel affected the development of the orogen. While both end-member models account for lower-level thrusting, higher-level normal faulting, and a partial-melt zone deformed at high temperatures and consequently exhumed, they require different thrust geometries, P-T-t histories, and kinematic relationships between the Tethyan and GHS. Attempts will be made to explore the similarities and differences between the two models in anticipation of finding a model that best represents the observed and calculated data.







Figure 2.2. Channel Flow model of Himalayan development (from Kohn, 2008; based on Beaumont et al., 2004 and Jamieson et al., 2004). Partial melt channel (gray zone, \geq 700°C) couples with an erosional front (EF) and propagates forward over time until hot, partially molten rocks are brought close to the surface before cooling begins.



Figure 2.3. Critical taper model of Himalayan development (from Kohn, 2008; based on Henry et al., 1997 and Bollinger et al., 2006). As erosion uniformly removes material from on top of the section, thrusts progressively under-plate, creating a series of in-sequence thrusts over time. A partial melt zone exists, but remains far from the orogenic front.

CHAPTER 3. METHODS OF ANALYSIS

3.1 *In situ* Th-Pb Ion Microprobe Dating of Monazite

Monazite grains were first identified in thin section using back-scattered electron (BSE) imaging on the Cameca SX-100 electron microprobe housed in the Department of Earth and Environmental Sciences at Rensselaer Polytechnic Institute, Troy, New York. The monazite grains are typically zoned and range in size from 20 to 300 µm. Monazite grains were then X-ray mapped using the electron microprobe (operating conditions below) to identify chemically distinct zones. They were then re-identified in thin section using an optical microscope, and, using a Medenbach microdrill, were subsequently drilled from the slides in ~3 mm disks, preserving the textural relationship of the monazite grain with the surrounding matrix. The glass disks containing the monazites were then mounted in 1-inch epoxy rounds with a block of 5-10 grains of polished standards (monazite 554, Harrison et al., 1999b). To aid in locating the monazite for analyses, reflected light images were taken of each glass disk containing a monazite grain, as well as the entire epoxy round (Figure 3.1). The sample rounds were cleaned in distilled water in an ultrasonic bath, dried, and gold coated. Monazite grains were analyzed using the Cameca ims 1270 ion microprobe in Department of Earth and Space Sciences, University of California-Los Angeles.

Details of analytical protocols for ²⁰⁸Pb/²³²Th dating of monazite using the Cameca ims 1270 ion microprobe have been describe previously by Harrison et al. (1995, 1999b).
Operating conditions for this study involved a primary beam current of 6-10 nA, a spot size of ~10-20 μ m, and a mass resolving power of 4500, which sufficiently separates all molecular interferences in the 204-208 mass range. Energy offsets were +10 to +15 eV for ²³²Th⁺, and -8 to -13 eV for ThO₂⁺. Total time per spot analysis was 12 minutes. ²⁰⁸Pb⁺/Th⁺ is corrected for common Pb using the relationship ²⁰⁸Pb^{*}/Th⁺ = (²⁰⁸Pb⁺/Th⁺)[1-(²⁰⁸Pb/²⁰⁴Pb)₅(²⁰⁴Pb⁺/²⁰⁸Pb⁺)], where (^{208Pb}/²⁰⁴Pb)₈ is the known ratio of the standard; the asterisk indicates the species is corrected for common Pb. The Pb/Th relative sensitivity factor required to calculate a Th-Pb age from isotopic data obtained from an unknown monazite is determined by referring the ThO₂/Th ratio determined in the sample analysis to a linear calibration curve that is constructed from several ion microprobe spot measurements of ThO₂/Th vs. Pb/Th from standard monazite 554 (Figure 3.2). This correction factor permits the determination of Pb/Th ratios of unknown grains measured under the same instrumental conditions. Reported age uncertainties reflect counting statistics and the reproducibility of the standard calibration curve.

3.2 Garnet Geochronology

Whole rock samples were processed in a jaw crusher and sieved until the entire portion passed through a 750 µm sieve. Approximately 0.5 to 1 g of garnet was handpicked under a binocular microscope, excluding grains with abundant visible inclusions. Contrary to many other studies, garnets were still retained if they contained a moderate number of inclusions since rim fragments tended to be inclusion-free and the sought-after age is that of the core. Dissolving these fractions should have no effect on calculated age as zircon is the only phase present as an inclusion that contains measurable Hf, and zircon is not dissolved with the garnet during table-top digestion procedures. Sample dissolution, chemical separations, and isotopic analysis protocols follow Vervoort et al. (2004) and Cheng et al. (2008), and are described in detail below. Each geochronological point called for ~0.25 g of garnet or whole-rock sample. Two aliquots of whole-rock powders were digested per sample. The first followed the same tabletop dissolution procedure as that of garnet to determine the zircon-free whole-rock composition. The second used high-pressure Teflon bombs to ensure total dissolution of refractory phases, including zircon.

Stage 1: Leaching and Sample Dissolution

Approximately 250 mg of garnet separate or whole-rock powder was weighed into clean 15 ml Savillex® beakers (or bombs, depending on dissolution method). The garnet samples were then subjected to a light leaching process where ~4 ml 1M HCl was added to the beakers, which were then capped, sonicated for 5 minutes, and let rest for 5 minutes. The acid was pipetted off, ~5 ml mQ H₂O was added, and the beaker was once again sonicated for 5 minutes. The water was pipetted from the sample.

The sample dissolution procedure for garnet and whole-rock powders began by adding 1-2 ml of a 10:1 mix of 29M HF: 14M HNO₃ to the beakers and drying them down on a hotplate set at ~150°C. Next, 10 ml of the 10:1 mixture was added and the beakers were capped and left on the hotplate for 24 hours. Whole-rock samples intended for total bomb dissolution were instead inserted into steel jackets and put in an oven for 3 days. Beakers (or bombs) were then uncapped and dried down on the hotplate. A boric acid/HCl mixture (1 ml H₃BO₃ + 2 ml 6M HCl per every 100 mg sample) was added to

complex fluorides, and the beakers were capped and left overnight on a hotplate at 100 °C. For bomb-dissolved whole-rock samples, a few additional steps were applied at this point: after the orthoboric acid step, the samples were dried down, ~10 ml 6M HCl was added, the bombs were capped and then left to sit on the hotplate overnight. The next day, the bombs were uncapped, dried down, and ~ 10 ml 6M HCl was added. The bombs were once again capped, jacketed, and put in the oven for ~6 hours. Once removed, the samples were transferred to Savillex® beakers. All samples were then dried down. 10 ml 6M HCl was added and the beakers were capped and left to equilibrate on the hotplate. Beakers were removed from the hotplate to cool, spiked with a mixed ¹⁷⁶Lu-¹⁸⁰Hf tracer, capped, equilibrated on a hotplate overnight, uncapped, and thoroughly dried down.

Stage 2: 1st Stage Hf and Lu (Separation of Hf and Lu from bulk matrix)

The sample was dissolved in 0.5 ml 6M HCl and sonicated for 5 minutes, then diluted with 2.5 ml mQ H₂O to make a 1M HCl solution, which was once again sonicated for 5 minutes. 20 μ L 8M HF was added to the solution to complex the Hf, and the beaker was warmed on the hotplate for ~10 minutes. Each sample was transferred to two 1.5 ml centrifuge vials and centrifuged at 3000-3500 rpm for 10-15 minutes.

Columns filled with 10 ml of Bio-Rad AG50w-X12 cation exchange resin were backwashed and conditioned with 1M HCl. Only the supernatant from the centrifuge tubes was loaded onto the resin. In the load solution, Hf and other HFSE were complexed as fluoride anions which do not interact with the resin. Consequently, the columns are washed in only 1 ml 1M HCl/0.1M HF before Hf is collected with 6 additional washes of 1 ml 1M HCl/0.1M HF. The columns were washed with 5 additional ml of 1M HCl/0.1M HF, followed by 100 ml 2.5M HCl to flush away most of the other elemental components in each sample. Lu + HREE were collected with 60 ml 2.5M HCl. To clean, columns were washed with 160 ml 6M HCl, 20 ml 6M HCl + 2 ml 8M HF, and then a full reservoir of 6M HCl. The Hf and Lu sample cuts were dried down in preparation for additional separation.

Stage 3: 2nd Stage Hf (Separation of Hf from other HFSE)

The dried sample cut of Hf from the previous separation was redissolved in 5 ml 2.5M HCl, capped, and heated to equilibrate. Columns filled with 1.0 ml of Ln-Spec resin were resettled and conditioned with 2.5M HCl. The sample was loaded onto the resin bed and washed with 10 ml 2.5M HCl, then 3 iterations of a 10 ml 6M HCl wash to remove any incorporated REE. The columns were washed with 5 ml H₂O to rinse away the HCl. To remove Ti, the columns were rinsed 4-6 times with 10 ml of 0.09M HCit/0.45M HNO₃/1wt% H₂O₂, until the orange color (created by the interaction of Ti with the citric acid) was washed from the resin, ensuring total removal of Ti. The columns were then washed with 5 ml 0.09M HCit/0.45M HNO₃ and 2 times 10 ml 6M HCL/0.06M HF. Hf was recovered in 5 ml 6M HCl/0.4M HF and dried down for the final Hf purification stage. The columns were cleaned with 10 ml 2M HF, 10 ml 6M HCl, 10 ml H₂O, and then stored in 0.14M HCl.

Stage 4: Hf Cleanup

The final stage of Hf separation and cleaning is to remove any remaining HREE that may present isobaric interferences during mass spectrometry. The columns used are

essentially extremely scaled-down versions of the 1st stage columns. To the dried Hf cut from the previous separation, 25 μ L 2.5M HCl were added, and the sample was heated on the hotplate for ~5 minutes. 30 μ L H₂O were added, and the sample was once again heated on the hotplate for ~5 minutes. 20 μ L 0.3M HF were added, and the sample was loaded into the columns conditioned with 1M HCl. Hf was collected with 5 rinses of 50 μ L 1M HCl/0.1M HF. The columns were cleaned with 120 μ L 1M HCl/0.1M HF and 5 ml 2.5M HCl. The final Hf fraction was dried down, treated with 1 drop 14M HNO₃ to remove any organic compounds, dried down, treated with 1 drop 8M HF to complex the Hf with fluoride, and dried down again.

Stage 5: 2nd Stage Lu (Separation of Lu from HREE)

The dried Lu cut from the first stage separation was redissolved in 0.5 ml 2.5M HCl. Columns filled with 0.9 ml of Ln-Spec resin were resettled and conditioned with 2.5M HCl. The solution was loaded onto the columns and washed 4 times with 7.5 ml 2.5M HCl and 2 times with 5 ml 2.5M HCl, which removes most, but not all, of the Yb from the sample, leaving enough to make an accurate fractionation correction, but not so much that the ¹⁷⁶Yb interference on ¹⁷⁶Lu is prohibitively large (Vervoort et al., 2004). Lu was collected with 5 ml 6M HCl. The columns were cleaned by filling the reservoir 3 times with 6M HCl, and were then stored in 0.14M HCl. The Lu cut was dried down, treated with 1 drop 14M HNO₃ to remove any organics, and dried down again.

Stage 6: Mass Spectrometry via MC-ICP-MS

The dried Lu and Hf fractions were dissolved in 750-1000 μ L 2% HNO3, sonicated for 10-15 minutes, and transferred to centrifuge tubes for mass spectrometry analysis.

Isotopic data were collected on a ThermoFinnigan NeptuneTM multi-collector inductively coupled plasma mass spectrometer (MC-ICP-MS) in the GeoAnalytical Laboratory at Washington State University. Protocols for Lu analysis and data reduction are described by Vervoort et al. (2004). Hafnium ratios were corrected for mass fractionation using ¹⁷⁹Hf/¹⁷⁷Hf = 0.7325 and the exponential law for mass bias correction. Isotopic compositions were then normalized using an external Hf standard, JMC475 (¹⁷⁶Hf/¹⁷⁷Hf = 0.282160). The average ¹⁷⁶Hf/¹⁷⁷Hf for JMC475 taken over the course of this study was 0.282142 \pm 0.000025 (2 σ , n=54). External uncertainties applied to measured data are 0.5% for ¹⁷⁶Lu/¹⁷⁷Hf and the quadratically combined 2-sigma in-run error and a blanket 0.005% uncertainty for ¹⁷⁶Hf/¹⁷⁷Hf. Lutetium-hafnium ages were calculated using the ¹⁷⁶Lu decay constant of 1.867x10⁻¹¹ (Scherer et al., 2001; Söderlund et al., 2004) and isochrons were produced using the program Isoplot/Ex (Ludwig, 2003). Errors are reported at 95% confidence.

3.3 Electron Microprobe Chemical Analysis

Elemental compositions and X-ray maps were collected using the Cameca SX-100 electron microprobe housed in the Department of Earth and Environmental Sciences at Rensselaer Polytechnic Institute, Troy, New York. Natural and synthetic silicates and oxides were used for calibrations, and quantitative measurements were made using an accelerating voltage of 15 kV and a current of 20 nA (major silicates) or 100 nA (rutile and titanite). A minimum beam size was used on most minerals, except plagioclase and micas (10 μ m), and rutile and titanite (2 μ m). Peak count times were 10 s (Na, Ca, Fe, Mn, Si, Al), 20 s (Mg, Ti, K), 45 s (Zr in rutile), and 75 s (Zr in titanite). Operating conditions for the X-ray maps consisted of an accelerating voltage of 15 kV, current of 200 nA, pixel time of 30 msec, and step size of 2-5 μ m/pixel. X-ray compositional maps of the elements Fe, Mg, Mn, Ca, and Al were collected on garnet, and Th, Y, U, Ce, and Ca on monazite. Garnet maps were collected via stage-mapping, monazites via beammapping.

3.4 Laser Ablation ICP-MS

Understanding the distribution of Lu in garnet is important for interpreting Lu-Hf isochrons within the context of progressive metamorphism. Therefore, trace element profiles were collected across eclogite garnets using a UP193 laser and Element XA magnetic sector ICP-MS housed at the University of Lausanne, Switzerland. Prior to data collection, line traverses were pre-ablated with a spot size of 25 μ m, a stage movement rate of 50 μ m/sec, and a power intensity of 6.2 J/cm², as metered directly at the surface. Traverse data were then collected with a spot size of 15 μ m, a stage movement rate of 10 μ m/sec and a power intensity of 6.2 J/cm². Intensities were measured in low resolution mode for a large suite of major and trace elements ranging from Mg to Hf, the most important issue being the distribution of Lu. Raw counts of ¹⁷⁵Lu were normalized to ²⁷Al, and standardized against NIST612 glass. For sample AR01-43c, concentrations of Lu were calculated from averages measured across the core versus rim.



Figure 3.1. Example reflected-light images of monazite samples. A) 1-inch epoxy round containing disks drilled from thin section and rectangular block of standards. B) Example 3 mm glass disk containing monazite from sample AS01-16a.



Figure 3.2. Th-Pb calibration curve for standard monazite 554 (see Harrison et al., 1999b). Along with the age of 554 (45 ± 1 Ma), this relationship allows the relative sensitivity factor to be calculated, permitting determination of Pb/Th ratios of unknown monazites under the same analytical conditions. Error ellipses represent 1 σ uncertainties on the measurement; error margin on calibration line represents 1 σ uncertainty on the calibration. n = 55.

CHAPTER 4. GEOCHRONOLOGY AND THERMOBAROMETRY OF THE CENTRAL HIMALAYA, ANNAPURNA TRANSECT, CENTRAL NEPAL

4.1 Introduction

Syntheses of Himalayan tectonics indicate 400-700 km of shortening have been taken up across the Himalaya over the last 20-25 Myr (e.g., Yin and Harrison, 2000; DeCelles et al., 2001; Guillot et al., 2003; Robinson et al., 2006). The implied average shortening rate, $\sim 2 \text{ cm/yr}$, corresponds well with estimates of modern rates of shortening across the Himalaya (Bilham et al., 1997; Larson et al., 1999). Yet despite decades of research, basic questions still remain about how strain was partitioned. Has the convergence rate remained constant at ~ 2 cm/yr, or has it fluctuated significantly over time? Can strain measurements and shortening rate estimates from one location be extrapolated across the orogen, or are they only valid for that region? Was movement along the major faults in different regions contemporaneous, or was movement stalled in some areas while active in others? Thermal and mechanical models typically assume a constant shortening rate in the Himalaya since ~25 Ma (e.g., Bilham et al., 1997; Henry et al., 1997; Harrison et al., 1998; Huerta et al., 1998, 1999; Beaumont et al., 2001, 2004; Jamieson et al., 2002, 2004), but significant variations in the displacement rate on million-year timescales have been proposed in central Nepal (Kohn et al., 2004), and elsewhere within the Indo-Asian collision (Dunlap et al., 1998). With the development of a new method that combines petrology and geochronology to determine thrust displacement rates in crystalline rocks (Kohn et al., 2004), strain partitioning in the past can be compared to present rates, as well as comparing rates and timing across the Himalaya. Fluctuations in the timing and rate of thrusting could have significant implications for strain partitioning across the Himalaya, and for partitioning during collisional orogenesis as a whole.

To address these issues, samples were collected from a transect through the crystalline Himalaya along the Modi Khola in central Nepal, south of the Annapurna Massif (Figure 4.1). This transect is located ~ 200 km west of a Langtang transect that has been studied extensively (Inger and Harris, 1992; Macfarlane, 1993, 1995; Fraser et al., 2000; Kohn, 2004; Kohn et al., 2004, 2005) and therefore serves as a good basis of comparison. Martin et al. (2010) reported new P-T estimates for Greater and Lesser Himalayan rocks along the Modi Khola, but did not offer them in a geochronological context. P-T estimates have also been published from other transects through the central Himalaya of Nepal, including a transect along the Kali Ghandaki river to the west of the Modi Khola (Vannay and Hodges, 1996), and along the Marsyandi (Catlos et al., 2001) and Darondi (Kohn et al., 2001) rivers to the east between the Modi Khola and Langtang. In order to verify whether thrusting and convergence at Annapurna had the same rate, timing, and thermochemical evolution along strike to Langtang (which would indicate whether results from one area of an orogen are likely to be representative of the orogen as a whole), the pressure-temperature evolution, age of peak metamorphism, timing of

movement along major Himalayan thrusts, and transport rate of the region will be established.

This work relies heavily on monazite crystallization ages. Monazite [(LREE, Y, Th)PO₄] is a common accessory mineral in metapelites (Overstreet, 1967; see also review of Spear and Pyle, 2002), and is a very popular chronometer in metamorphic rocks because of its high U and Th contents, very low initial Pb, and high retentivity of radiogenic Pb (e.g., see summaries of Parrish, 1990; Harrison et al., 2002). A critical step when employing monazite chronology is linking ages to corresponding metamorphic conditions via monazite chemistry. Many geochronologists still fail to make this link, even though numerous studies have shown the petrogenetic and tectonic value of this endeavor (e.g., Pyle and Spear, 1999, 2003; Ferry, 2000; Foster et al., 2000, 2002, 2004; Pyle et al., 2001, 2005; Spear and Pyle, 2002; Wing et al., 2003; Gibson et al., 2004; Kohn and Malloy, 2004; Dahl et al., 2005; Kohn et al., 2004, 2005; Corrie and Kohn, 2008). By investigating these links, monazite ages may be assigned to specific points along the P-T path of a metamorphic assemblage, i.e., reflecting prograde, retrograde, or hydrothermal processes.

4.2 Petrography and Petrology

4.2.1 Mineral Assemblages

Lesser Himalayan Sequence metapelites are generally characterized by the mineral assemblage $Qtz + Ms \pm Pl \pm Bt \pm Grt \pm Chl$, with accessory apatite + tourmaline + zircon \pm ilmenite \pm allanite (Table 4.1). Two samples from one location contain

chloritoid. LHS rocks that have a higher Ca-content may also contain hornblende and calcite. The foliation in rocks below the Munsiari Thrust is defined primarily by muscovite and chlorite; biotite in these assemblages is fine grained. Further upsection, within 500 m below the MT, and between the MT and MCT, biotite and muscovite define the primary foliation and the mode of prograde chlorite decreases until it disappears from the prograde assemblage altogether.

The mineral assemblage of GHS Formation I pelitic rocks is predominately $Qtz + Pl + Bt \pm Ms \pm Grt$, with accessory apatite + tourmaline + zircon ± rutile ± monazite ± staurolite ± xenotime ± epidote (Table 4.2). A few of the more calcic beds may contain hornblende, chlorite, and clinozoisite. Although several samples contain kyanite, AS01-22a is the only sample that contains sillimanite, which is locally developed as fibrous mats along quartz-plagioclase and plagioclase-plagioclase grain boundaries (Figure 4.2). This sillimanite does not appear to be a regional feature. Kaneko (1995) reported similar sillimanite textures in the Modi Khola valley.

The calc-silicates and calcic quartzites from GHS Formation II contain the assemblage $Qtz + Pl + Hbl \pm Grt \pm Ms$, with accessory calcite + clinopyroxene + apatite \pm titanite \pm clinozoisite \pm zircon \pm potassium feldspar \pm epidote. Formation II amphibolites have the assemblage $Qtz + Pl + Hbl + Cpx \pm Grt \pm Chl$, with accessory apatite \pm titanite \pm zircon \pm epidote.

4.2.2 Mineral Chemistry

X-ray maps of garnet below the Munsiari Thrust show that major element zoning is smooth and typical of garnet growth (Figures 4.3-4.5). Grossular and spessartine contents decrease from core to rim, whereas almandine and pyrope increase, as expected for prograde garnet growth (Spear et al., 1990). Additionally, these garnets tend to display S-shaped or snowball inclusion textures due to rotation during crystal growth (Figure 4.6). These spiral inclusion trails are suggestive of up to 160° of rotation relative to the matrix schistosity, which consistently indicates top-to-the-SSW shearing (Hodges et al., 1996).

Garnet from rocks within 500 m of the MCT (above and below) retains primary zoning patterns, particularly for calcium (Figure 4.7 and 4.8). Oscillatory zoning of calcium has been observed in garnets near the MCT in the Darondi and Langtang regions of central Nepal (Kohn et al., 2001; Kohn et al., 2005) and was proposed to result from heterogeneous (instead of constant) thrusting and heating along the MCT (Kohn et al., 2005).

Garnet X-ray maps from GHS Formation I samples indicate progressively more diffusional modification structurally upwards. Near the MCT, from GHS Formation 1a, garnet still retains Mn and Mg trends that are consistent with growth zonation (Figure 4.8). Further from the MCT, garnet zoning becomes increasingly relaxed due to diffusion. This is typical in high-grade metamorphic rocks (e.g., Florence and Spear, 1991), and is consistent with that reported from this transect (Kaneko, 1995; Martin et al., 2010) and in the Marsyandi and Langtang regions of central Nepal (Catlos et al., 2001; Kohn et al., 2004). In Formation 1b, garnet zoning profiles are flat (Figure 4.9), while even higher upsection in Formation 1c, garnets display homogenous zoning in the cores, but have rim zoning that is consistent with resorption and back diffusion during cooling (Florence and Spear, 1991; Figure 4.10).

Plagioclase grains show core-rim zoning with higher-Ca cores and lower-Ca rims (conversely, lower-Na cores and higher-Na rims). This is expected from garnet growth and fractional crystallization, as garnet growth depletes the matrix in Ca (Spear et al., 1990).

4.3 Thermobarometry

Peak temperature and pressure estimates were calculated via exchange reactions and thermodynamic equilibria. The garnet-biotite thermometer of Ferry and Spear (1978) with the Berman (1990) garnet solution model was used for all samples except AS01-41b, in which the garnet-chlorite thermometer of Dickenson and Hewitt (1986; modified in Laird, 1988) was used in conjunction with the Berman (1990) garnet solution model. Depending on the mineral assemblage of the sample, pressures were calculated using the barometer garnet-plagioclase-aluminosilicate-quartz (Koziol and Newton, 1988), garnetplagioclase-muscovite-biotite (Hoisch, 1990), or garnet-plagioclase-biotite-quartz (Hoisch, 1990). In two samples from the Lesser Himalayan Sequence plagioclase was not found, so only temperature was estimated. Different thermobarometric calibrations may shift the pressures and temperatures reported for each sample by as much as 25°C and 1 kbar, but the major trends are preserved.

4.3.1 Selecting Compositions for Thermobarometry

Appropriate mineral compositions were selected using standard petrologic criteria (e.g., Kohn, 2008) to provide the best estimate of peak metamorphic conditions (Tables 4.3-4.8). Mineral inclusions suitable for thermobarometry are sparse, so directly constraining prior metamorphic assemblages, and their pressures and temperatures, is difficult. Thus, garnet chemistry provides the principal record of evolving P-T conditions. The composition of garnet that most closely preserves peak metamorphic conditions is the composition nearest the rim that was not affected by retrograde reactions. In the LHS, temperatures were sufficiently low that prograde garnet compositions are commonly retained, and compositions at or near the rim were selected. These garnet compositions were combined with compositions of plagioclase (when present), biotite, muscovite, and chlorite (only one sample) near the garnet that appeared to be texturally equilibrated, rather than retrograde products.

In the GHS, however, sufficiently high temperatures were reached that diffusion, retrograde exchange reactions, and/or retrograde net transfer reactions (ReNTRs) either partially or entirely modified the original growth zoning in garnet (e.g., Kohn and Spear, 2000; Catlos et al., 2001; Kohn et al., 2001, 2004). Homogenization of major element zoning in garnet compromises retrieval of prograde pressures and temperatures. Exchange reactions entail exchange between two elements in two minerals without consuming or producing new minerals. For example, Fe-Mg exchange between biotite and garnet during cooling causes biotite to become more Mg-rich and garnet to become more Fe-rich, causing estimated temperatures to be erroneously low. Retrograde net

transfer reactions consume garnet, returning its constituent elements to the matrix; this process can enrich nearby minerals in Fe and Ca, but can also cause Mn (±Fe) to pile up on the garnet rim and back-diffuse towards the core. Garnet has a much greater Fe/Mg ratio than biotite, and dissolution enriches both biotite and garnet rims in Fe (Kohn and Spear, 2000). These retrograde modifications cause garnet zoning patterns to exhibit a increase at the rim in Mn and Fe, and a decrease in Mg. A common method to circumvent this problem is to use the composition of garnet where Mn and Fe/(Fe+Mg) form a trough. That is, the composition nearest the rim that was not affected by retrograde reactions. In some cases where diffusional homogenization has occurred, this location is near the core of the garnet. However, pairing a garnet composition with a nearby biotite composition that has experienced Fe enrichment will cause the estimated temperatures to be too high. Therefore, a correction must be made to the biotite composition to account for Fe enrichment due to garnet dissolution.

Using garnet X-ray maps to estimate the amount of dissolution of the garnet, the amount of Fe enrichment in biotite due to ReNTRs may be estimated, and the biotite compositions subsequently corrected (Kohn and Spear, 2000). In some samples, the correction to the biotite compositions was minor, and equivalent temperatures could be calculated by using compositions from biotite distal to the garnet that were not affected by ReNTRs. In other samples, correction to the biotite compositions lowered estimated temperatures 10 to 50 °C. To calculate P-T conditions in GHS samples, the garnet composition with the lowest Fe/(Fe+Mg) and Mn were chosen and combined with

corrected or distal biotite (depending on the sample), proximal muscovite, and rim compositions of proximal plagioclase.

Evidence to support the estimated P-T conditions includes agreement with the stability of observed mineral assemblages and consistency with P-T conditions of nearby samples that have different compositions.

4.3.2 Estimated P-T Conditions

Rocks from the Modi Khola valley show chemical and textural evidence for only one episode of metamorphism, and display the familiar apparent inverted metamorphism associated with the Main Central Thrust. Mineral isograds increase progressively from biotite and garnet in the LHS up to at least kyanite in the GHS. Below the Munsiari Thrust in the LHS, rocks experienced peak metamorphic conditions of about 515 °C and 7.5 kbar (Table 4.9, Figure 4.11). Within the Munsiari Thrust sheet, pressures and temperatures were about 555 °C and 10.5 kbar. Peak pressure and temperature conditions increase across the MCT (Figs. 4.12 and 4.13) where, at the base of the thrust sheet, P-T conditions are 650 °C and 12.0 kbar, similar to data reported from the hanging wall of the MCT in the Annapurna region (Kaneko, 1995) and in the Marsyandi, Darondi and Langtang regions of central Nepal (Catlos et al., 2001; Kohn et al., 2001, 2004; Kohn, 2008), though much lower than the P-T conditions reported by Martin et al. (2010) for the same transect (Figure 4.14). Temperatures increase systematically upsection away from the MCT, to 735 °C and 12.5 kbar in the middle of Formation I (Fm. 1b) to 775 °C and 12.5 kbar at the top of Formation I (Fm. 1c). These P-T conditions accord with those

of Martin et al. (2010), as do those from Formation II, which we estimate to be \sim 775 °C and 14.0 kbar.

4.4 Monazite

4.4.1 Monazite Petrogenesis

Monazite is a common accessory in metapelitic GHS rocks and is typically zoned in Thorium (Th) and Yttrium (Y). As previously mentioned, variations in these elements are often associated with reactions among the major silicates. As monazite grows, it preferentially fractionates Th, progressively depleting the matrix, and resulting in earlier vs. later grown monazite having higher vs. lower Th contents (Kohn and Malloy, 2004). Allanite also hosts Th, and though allanite is the primary LREE-accessory phase in the LHS, it was found in only one GHS rock – a quartzite with a high-calcium bulk composition. Monazite also sequesters Y into its structure, but so do other minerals including garnet, allanite, and xenotime (Spear and Pyle, 2002), potentially complicating interpretation of Y systematics. No evidence exists for prograde allanite or xenotime in any of the monazite-bearing GHS rocks. Garnet x-ray maps of Y do not indicate growth in the presence of xenotime, though a few samples contain minor retrograde (?) xenotime that is texturally associated with monazite rims. Yttrium systematics in most samples thus depend principally on the interplay between garnet and monazite (Pyle and Spear, 1999, 2003; Pyle et al., 2001). During prograde growth, Y is incorporated into garnet, so concentrations of Y in monazite should decrease with progressively higher grade. This simple trend changes during partial melting (normally via muscovite dehydrationmelting), where monazite begins to dissolve, erasing direct chemical or chronological evidence of peak metamorphic temperatures. However, this reaction liberates Y to the melt (Spear and Pyle, 2002), where, upon cooling and melt crystallization, it is incorporated in relatively high concentrations in re-growing monazite. Thus, post-anatectic rims of monazite should be chemically distinct (high-Y and possibly high-Th) from monazite that grew along the prograde path (e.g., Kohn et al., 2005). That is, Th and Y systematics in monazite allow specific zones and ages to be chemically correlated within P-T space (Kohn et al., 2005).

4.4.2 Monazite Chemistry and Th-Pb Geochronology

Tables 4.10 and 4.11 list compositions of Himalayan monazites from two GHS rocks (corresponding to spot analysis locations in Figure 4.15). ThO₂ and Y₂O₃ concentrations ranges from 3.2-8.1 wt% and 0.5-2.8 wt%, respectively. Most GHS monazite grains show substantial Y and Th zoning. As expected due to Th fractionation, grains from several samples display a rimward decrease in Th. Many grains exhibit low-Y cores and higher-Y rims, but a number of samples contain grains that have either a relatively homogenous Y distribution or a mottled Y-zoning pattern that shows no distinct Y zones, but is not homogenous. These different zones were targeted for *in situ* ion probe analysis to chemically link (wherever possible) each zone and its corresponding age.

Monazite from GHS Formation 1a contains mottled Y and Th zoning and no distinct high-Y rims that would indicate partial melting, which is in agreement with calculated sub-anatectic P-T conditions for those samples (Figure 4.16, Table 4.12). The 21-25 Ma ages from these grains are thus considered to represent prograde metamorphism. One monazite from sample AS01-15b contains evidence for an inherited component, with a low Y and Th core and anomalously old, probably mixed ages that are not interpretable within the context of the Indo-Asian collision.

Formation 1b monazite either displays distinct low-Y cores and high-Y rims, or an indistinct zoning that is nevertheless geochronologically inhomogeneous (Figure 4.17, Table 4.12). Low-Y, prograde core ages are 24-29 Ma, though one monazite from AS01-16a has an age of 33 Ma, presumably the result of early prograde nucleation and growth. High-Y rim cooling ages are 17-22 Ma. Sample AS01-17b does not show signs of multiple episode of growth, as it does not contain well-defined cores with overgrowth rims. Rather, these monazites have slightly high Th cores, rimward Th decreases, and a relatively flat distribution of Y. However, the monazites are small, high-Y, and exhibit relatively young ages (~18 Ma), and so are interpreted to have entirely grown during post-anatectic cooling.

Monazite in samples from Formation 1c generally have low-Y cores and high-Y rims, the product of prograde growth and post-anatectic cooling, respectively (Figure 4.18, Table 4.12). One grain from sample AS01-32 contains mottled zoning and slightly older ages, representing early prograde growth. Post-anatectic cooling and monazite growth occurred 19-15 Ma, as indicated by high-Y rims or small homogenous high-Y monazites with ages coincident with rim ages.

Monazite from GHS Formation III, just below the South Tibetan Detachment, exhibits either a mottled zoning pattern, or slightly higher-Th cores that decrease rimward. There is no evidence for rim growth associated with cooling after melting (Figure 4.19, Table 4.12). Higher Th zones in the monazite are 35-38 Ma and are ascribed to early prograde growth with later prograde growth represented by 30-35 Ma mottled Y and Th zones.

4.4.3 Monazite Age Distribution

Four generations of monazite have been identified in samples south of the Annapurna range along the Modi Khola in central Nepal. From oldest to youngest they are: inherited, early prograde, late prograde, and post-anatectic, and can be distinguished chemically and/or chronologically.

Age probability diagrams for monazite (Figure 4.20) show distinct peaks that correspond to specific chemistries, and thus to petrologic origins. Geochronologic analyses that clearly contained mixed zones (i.e., the ion probe spot covers two distinctly separate zones in the monazite) were not used in the age profile calculations. Particularly significant is the timing of the last growth of prograde monazite (youngest low Y and Th peak) versus the first instance of melt crystallization during cooling (high-Y overgrowth). Together, these ages bracket the timing of anatexis. Additionally, the diagrams demonstrate decreasing temperatures and ages of metamorphism structurally downward, consistent with progressive underthrusting of tectonic slices. In the structurally highest rocks of the Formation III orthogneiss, the last prograde monazites that formed have an age of 30-33 Ma. Structurally lower in the next highest unit, Formation 1c, the latest prograde sub-solidus monazite formed 27-30 Ma (3 Ma younger than Fm. III), with crystallization of *in situ* melts 19-22 Ma. Peak-T metamorphism and anatexis thus probably occurred $\sim 25 \pm 2$ Ma in Formation 1c. Structurally lower rocks have the same chemically defined generations of monazite, but again displaced to younger ages and lower temperatures. Ages from Formation 1b imply melting at 23 ± 1 Ma, with post-anatectic monazite growth from 22 to 17 Ma. This means that both packages of rocks (Formations 1b and 1c) underwent similar reactions (specifically muscovite dehydration melting and ensuing melt crystallization), but at different times. Monazite in Formation 1a experienced sub-solidus growth from 21-16 Ma.

4.5 Discussion

Due to the observed chronologic and thermobarometric differences between GHS rock units in the Annapurna region, two thrusts are identified within GHS Formation I: the Bhanuwa Thrust (BT; between Formations 1a and 1b), and the Sinuwa Thrust (ST; between Formations 1b and 1c; Figure 4.1). The Sinuwa Thrust is inferred to place high-T, pervasively migmatitic rocks of GHS Formation 1c atop slightly lower-T locally migmatitic rocks of Formation 1b. Based on monazite geochronology (Figure 4.20), initial melts may have formed in the Sinuwa thrust sheet (Formation 1c) as early as 27 Ma, but not until 23 Ma in the Bhanuwa thrust sheet (Figure 4.21). The large gap between pre- and post-anatectic monazite in the Sinuwa thrust sheet precludes pinpointing the timing of initial cooling. Possibly cooling commenced as early as 25-26

Ma, during heating of the Bhanuwa thrust sheet, implying emplacement of the ST at that time. Alternatively, the ST and BT rocks may have been buffered at anatectic conditions until cooling and melt crystallization in both sheets at ~22 Ma. From this point on, the two sheets probably experienced similar cooling histories as they were transported together in the hanging wall of the Bhanuwa Thrust. Presuming that cooling resulted primarily from thrust emplacement (Kohn et al., 2004; Kohn, 2008) rather than from erosion or vertical displacement, these data imply initial ST movement some time between 22 and 27 Ma, and initial BT movement at 22 Ma.

A fault between GHS Formations 1a and 1b was first identified by Martin et al. (2010) as a normal fault based upon a 4 kbar pressure difference between hanging wall and footwall rocks, shorter retrograde chemical profiles in garnet in footwall rocks, and the presence of kyanite in the hanging wall vs. its absence in the footwall. P-T data from this study do not support such a disparity in the pressures across the fault. Instead, our data indicate very little increase in pressure structurally upward, but ~85 °C increase in temperature (Table 4.9). This temperature increase can account for the presence of kyanite and longer retrograde diffusion profiles in the hanging wall vs. the footwall. P-T calculations and monazite geochronology do, however, support the presence of a fault in this location. Thus, we revise the interpretation of the Bhanuwa Fault from a normal fault to a thrust. Monazite data from the Bhanuwa thrust sheet and the structurally lower Formation 1a rocks in the MCT thrust sheet support the estimated P-T conditions in both packages. Temperatures in the Bhanuwa thrust sheet were high enough that partial melting would be expected via the muscovite dehydration reaction, and high-Y rims on

monazite due to post-anatectic cooling and crystallization confirm this. Underneath, in the MCT thrust sheet, peak temperatures were ~650 °C (Table 4.9), below partial melting temperatures, and monazite in these rocks does not contain evidence that they experienced melting and post-anatectic growth (Figure 4.16).

As noted by Viskupic and Hodges (2001), Harrison et al. (2002), and Kohn et al. (2005), there are some significant interpretational uncertainties for most published Greater Himalayan Sequence monazite ages. Recent work has shown that there can be three to five different generations of monazite of varying age and chemical compositions that can encompass over 20 million years of Himalayan metamorphism (Kohn et al., 2005; this study). A single age from a bulk analysis, single grain, or even grain fragment analyzed via ID-TIMS (e.g., Hodges et al., 1996; Coleman, 1998; Simpson et al., 2000; Godin et al., 2001; Johnson et al., 2001) may be a real age, but it could also be a geologically meaningless mix of inherited components, prograde and retrograde growth domains, and late-stage alteration (Harrison et al., 2002; Kohn et al., 2005).

Other monazite Th-Pb ion probe ages have been published from along the Marsyandi River valley, approximately 100 km east of the Annapurna transect (Catlos et al., 2001). However, the spot locations for the ages were presented without chemical context. That is, it is unclear whether the ages are collected from cores or rims, high-Y or low-Y domains, or even if the spot location overlapped two distinct chemical or chronological zones, thus rendering that age geologically meaningless. Probability distributions of these data generally show a range in ages that spans the entire course of Himalayan metamorphism (Figure 4.21). The age distribution from Formation 1b in the Marsyandi valley could, in principle, reflect a similar cooling history to Formation 1b in the Annapurna region (i.e., 17-20 Ma), assuming those ages are from high-Y, postanatectic rim domains and not cores. That is, monazite ages from the Marsyandi region show no overall disagreement with the data presented in this study, but are not especially diagnostic of the timing of thrust movement.

Combining thermobarometric and geochronologic data allows rates of thrusting to be calculated (e.g., Kohn et al., 2004). These calculations require that cooling in the hanging wall results from thrust juxtaposition against cold footwall rocks, rather than from erosion or displacement with a significant vertical component. This assumption is consistent with original shallow thrust orientations (DeCelles et al., 2001), allowing these faults to accommodate significant convergence at depth (30-40 km depth; Figure 4.22). Barometric estimates of footwall and hanging wall rocks are not so different that vertical exhumation due to erosion or vertical transport could have been the primary means of cooling. In addition, where the Main Himalayan Thrust is at those depths, modern erosion directly above is <2 mm/yr (Lavé and Avouac, 2001). The other assumption is that among the Sinuwa, Bhanuwa, and Main Central Thrusts, only one thrust may be active at any given time. This is problematic (see below), but allows for limits to be placed on the slip rate along fault surfaces.

Thrust rates were calculated using the method of Kohn et al. (2004). The minimum amount of cooling during fault slip was approximated based on the peak temperatures in the hanging wall and those in the footwall. Thermal models for the central Himalaya (Henry et al., 1997; Huerta et al., 1999) define temperature-depth

distributions on which peak P-T data may be plotted, and allow Δ T's of cooling to be converted into thrust distances (Figure 4.22, Table 4.13). Effectively, thermal models define lateral thermal gradients, from which Δ T's are converted to displacements. Peak metamorphic ages from monazite in different thrust sheets provide a maximum duration of thrust displacement. The thrust rate is thus derived from the displacement distance over the time period of displacement.

In reality, uncertainties in peak temperature and chronology propagate to large errors for estimates for the Sinuwa and Bhanuwa thrust rates. Thus we mostly consider our results in terms of evaluating consistency (or not) with a nominal 2 cm/yr Himalayan convergence rate. For the MCT sheet, the cooling rate is more precise, but because pressures appear slightly higher than in the underlying MT sheet, some component of vertical cooling must be considered. A maximum pressure difference of 3 kbar (10 km), and typical Himalayan thermal models (e.g., Royden, 1993; Henry et al., 1997; Harrison et al., 1998; Huerta et al., 1998, 1999; Beaumont et al., 2001, 2004; Jamieson et al., 2002, 2004), imply that up to 50 °C of MCT sheet cooling could be ascribed to vertical rather than lateral displacement. This leaves at minimum 200 (±30) °C of cooling to be explained by lateral transport.

The calculated slip rates for the Sinuwa, Bhanuwa, and Main Central Thrusts are 0.5 to 2.3 cm/yr at 23 to 27 Ma (ST), 0.8 to 3.2 cm/yr at 19 to 23 Ma (BT), and 1.4 to 2.4 cm/yr at 15 to 22 Ma (MCT; Figure 4.22, Table 4.13). These calculations show P-T-t consistency with a 2 cm/yr convergence rate component across the Himalaya from ~25 to ~15 Ma. These results support models that presume constant rates since at least 25 Ma

(Royden, 1993; Henry et al., 1997; Harrison et al., 1998; Huerta et al., 1998, 1999; Beaumont et al., 2001, 2004; Jamieson et al., 2002, 2004).

The similarity in inferred peak ages for the ST, BT, and MCT could in principle indicate simultaneous initial cooling of all three sheets at ~22 Ma (Figure 4.23 model 2). Because pressures appear to increase structurally upward slightly, such cooling is unlikely to occur in a single thrust with an inverted thermal gradient. Rather, a simpler explanation would involve transport and cooling of each rock package along the same basal thrust (the MCT), and later juxtaposition of already-cooled rocks along the ST and BT. That is, this model implicitly requires as much as 100 km of out-of-sequence thrusting. Whereas we cannot rule out this possibility on the basis of P-T-t data alone, we note that such large out-of-sequence thrusts have not yet been reported elsewhere in the Himalaya.

Geochemically similar generations of monazite occur at both Langtang and Annapurna (i.e., early prograde mottled zoning, late prograde low Y, and post-anatectic high-Y rims), but monazite ages along the Modi Khola appear older by ~5 Myr compared to Langtang. The MCT sheet at Langtang records late prograde monazite growth 16-23 Ma and post-anatectic cooling 13-16 Ma (Kohn et al., 2005), while MCT rocks at Annapurna record late prograde monazite growth 21-23 Ma and cooling 16-20 Ma. By comparison, rocks in the Langtang Thrust (LT) sheet at Langtang record late prograde monazite growth 22-23 Ma and cooling 17-19 Ma (Kohn et al., 2005). This could potentially indicate that the MCT at Annapurna is temporally equivalent to the LT at Langtang. One explanation for the observed geochronological differences between the MCT at Annapurna and Langtang is the presence of a lateral ramp along strike (Figure 4.24, t₁). That is, the thrust plane cuts up section from west to east. Alternatively, the LT and MCT shear zones could have died out laterally (Figure 4.24, t₁). Because each thrust surface represents the accumulation of strain over millions of years, and because thrust surfaces may have shifted differently in different areas, the present-day distribution of lithologic packages (GHS Fm 1a, 1b, 1c, LHS, etc.) may not uniquely elucidate thrust evolution. That is, the MCT surface is defined on lithologic, not chronologic, criteria and juxtaposition of GHS and LHS rocks may have been diachronous along strike.

Assuming that the MCT at Annapurna and the LT at Langtang were active at the same time (Figure 4.24, t_1), the sequence of thrusting could have followed a number of progressions to obtain their current geometry. One possibility is that the thrust plane consistently cut upsection between Annapurna and Langtang, connected by a lateral ramp (Figure 4.24, t_2a), so that the thrust surfaces followed a sequence of progressive under-thrusting in both regions. This is consistent with the BT and ST in Annapurna being older structures than the LT in Langtang, presuming that the MT in the Annapurna region would be of similar age to the MCT at Langtang. A similar model suggests the same progression down to the MCT, but protracted MCT transport at Annapurna coincident with movement at Langtang (Figure 2.24, t_2b). Yet another alternative suggests concurrent movement of the MCT at Annapurna and LT at Langtang along non-continual thrust planes, and the initiation of a lateral ramp cutting down-section 16-17 Ma that

would juxtapose already cooled BT and ST sheets coeval with initial movement along the MCT at Langtang (Figure 2.24, t₂c).

As stated before, such large-scale diachroneity or out-of-sequence thrusting has not been reported before in the Himalaya, however our ability to resolve geological events has only recently progressed to the point that allows us to look at strain partitioning over million year timescales and potentially distinguish previously unrecognized spatial and temporal heterogeneity. The data show a complexity that does not lead to a concrete answer, but nonetheless, indicates that more work examining differences in strain partitioning on million-year or smaller timescales might elucidate previously unrecognized variabilities.

Despite differences in absolute ages, the similarities among the chemical systematics of monazite, peak P-T conditions, and thrust rates calculated for Langtang and Annapurna do, however, imply that strain measurements in one part of an orogen can be realistically extrapolated to another within a few hundred kilometers, although the timing of movement on discrete thrust surfaces may differ. This lateral predictability may only hold for geologically similar regions of the orogenic belt. With distance may come a change in boundary conditions (like rotation at the syntaxes) that would prevent long-distance extrapolation of strain estimates. Additional measurements from other parts of the orogen in India and Bhutan would be needed to verify whether these estimates could be applied to the orogen as a whole.



Figure 4.1. Geologic map of the Annapurna region along the Modi Khola. Geology after Colchen et al. (1986), Hodges et al. (1996), Pearson and DeCelles (2005), and Martin et al. (2007). Numbers for sample locations follow the pattern AS01-X. Barbed lines represent thrust faults: MT = Munsiari Thrust, MCT = Main Central Thrust, BT = Bhanuwa Thrust, ST = Sinuwa Thrust. Single dashed lines are normal faults: STDS = South Tibetan Detachment system. Labeled squares indicate towns. Black triangles are peaks: AS = Annapurna South, H = Hiunchuli.



Figure 4.2 Photomicrograph of fibrolitic sillimanite from sample AS01-22a. Image taken under plane-polarized light.









Mn

Figure 4.4. Lesser Himalayan Sequence AS01-40e garnet. White line is traverse across garnet from rim to core. Vertical tick marks on Fe/(Fe+Mg) indicate analysis points. Scale bars 200 microns.







Figure 4.6. Garnet photomicrograph from sample AS01-41b exhibiting snowball inclusion textures due to rotation during crystal growth. Photomicrograph taken under plane-polarized light.










Figure 4.9. Greater Himalayan Sequence AS01-18d garnet. White line is traverse across garnet from left to right. Vertical tick marks on Fe/(Fe+Mg) indicate analysis points. Scale bars 500 microns.







Figure 4.11. Pressure vs. temperatire plot for rocks collected along the Modi Khola. Thermobarometers are from Ferry and Spear (1978), Berman (1990), Dickenson and Hewitt (1986; modified in Laird, 1988; AS01-41b only), Koziol and Newton (1988), and Hoisch (1990).



Figure 4.12. Temperature vs. structural distance from Main Central Thrust. Temperatures show a strong gradient across the MCT, then a gradual increase structurally upward. Heavy gray line represents T-distance trend from tansects across the MCT in the Langtang and Darondi regions of Nepal (Kohn et al., 2001; Kohn, 2008). The T profile from this study corresponds well with those collected elsewhere in Nepal.



Figure 4.13. Pressure vs. structural distance from Main Central Thrust. Pressures are typically high, show a strong gradient across the MT and MCT, then a gradual increase structurally upward. Heavy gray line represents P-distance trend from tansects across the MCT in the Langtang and Darondi regions of Nepal (Kohn et al., 2001; Kohn, 2008). The P profile from this study does not show the decrease in pressures structurally upward as seen in Darondi and Langtang. One potential reason for this is the narrowness of the GHS in the Annapurna region versus the relatively wide GHS elsewhere in Nepal.



Figure 4.14. Pressure vs. temperature plot for rocks from the Annapurna region (boxes = this study; circles = Martin et al., 2010). Data from Martin et al. (2010) corresponds well with calculated P's and T's in this study with the exception of data from GHS Formation 1a.



Figure 4.15. X-ray maps of monazite with quantitative chemical analysis locations. Analyses are from samples AS01-16a and AS01-19c. Red dots indicate spot locations for electron microprobe analyses (Tables 4.10 and 4.11). Numbers next to the red dots correspond to the analysis number in the tables. Warm colors indicate higher concentration of element, cool colors lower concentration. Scale bars are 50 µm.



Figure 4.16. X-ray maps of monazite from Greater Himalayan Sequence Formation 1a with ion probe data. Ovals and their corresponding ages represent SIMS analysis spot locations. These monazites exhibit mottled Y and Th zoning patterns, except for two monazites from AS01-15b, which have low Th cores. Geochronologic analysis of this domain indicates an inherited component, yielding geologically meaningless ages. Uncertainties are 2σ ; scale bars are 50 µm. Xno = xenotime.



Figure 4.17. X-ray maps of monazite from Greater Himalayan Sequence Formation 1b with ion probe data. Ovals and their corresponding ages represent SIMS analysis spot locations. Monazites display distinct low-Y cores and high-Y rims, or an indistinct zoning that is geochronologically inhomogeneous. Uncertainties are 2σ ; scale bars are 50 μ m.



Figure 4.17 (continued).



Figure 4.18. X-ray maps of monazite from Greater Himalayan Sequence Formation 1c with ion probe data. Ovals and their corresponding ages represent SIMS analysis spot locations. Monazites display distinct low-Y cores and high-Y rims, the product of prograde growth and post-anatectic cooling, respectively. Uncertainties are 2σ; scale bars are 50 μm.



Figure 4.19. X-ray maps of monazite from Greater Himalayan Sequence Formation III with ion probe data. Ovals and their corresponding ages represent SIMS analysis spot locations. Monazites display mottled zoning patterns or slightly high-Th cores that decrease rimward. There is no evidence for rim growth associated with cooling after melting. Uncertainties are 2σ ; scale bars are 50 µm. Xno = xenotime.



Figure 4.20. Probability distributions of monazite ages, distinguished according to monazite chemistry. Red lines illustrate timing of peak metamorphism. Arrow in GHS Fm 1a box shows a published Ar/Ar muscovite age from the same structural level (Vannay and Hodges, 1996).



Figure 4.21. Probability distributions of monazite ages. Data from Catlos et al. (2001) Greater Himalayan Sequence Formation I samples. Further division into Formation 1a, 1b, and 1c for comparison to this study is based upon petrology (e.g., mineral assemblages and degree of migmatization).



Figure 4.22. A) Schematic illustration of the displacements of different thrust packages in the context of the modern Main Himalayan Thrust and thermal models for the Himalaya (Henry et al., 1997). The location and timing of each spot is constrained by the thermobarometric calculations, the thermal structure, and cooling ages determined by monazite or muscovite geochronology. B) Pressure-temperature diagram with predicted P-T paths based on thermal model in (A) and average peak P-T conditions in GHS and LHS. Critical taper models predict nearly isobaric heating and cooling P-T paths for GHS rocks and 'hairpin' P-T paths for LHS rocks (Kohn, 2008).



Figure 4.23. Schematic diagram showing two end-member models of thrust development. Model 1 exhibits in-sequence thrusting as the thrust plane cuts down section. Model 2 describes out-of-sequence thrusting initial cooling of all three thrust packages at ~22 Ma. This involves transport and cooling of each thrust package along the same basal thrust surface (the MCT) and later juxtaposition of the BT and ST. MCT = Main Central Thrust; MT = Munsiari Thrust; BT = Bhanuwa Thrust, ST = Sinuwa Thrust.



Figure 4.24. Schematic diagram of potential thrust development. t₁ represents an initial thrust structure at 19-20 Ma in which the two thrusts either die out along strike, or have a lateral ramp connecting the Annapurna and Langtang regions. Potential progressions include continued under-thrusting preserving the lateral ramp (t₂a), prolonged MCT transport at Annupurna allowing the MCT at Langtang to catch up (t₂b), or the development of a lateral ramp accomodating synchronous movement of cooled BT and ST thrust sheets at Annapurna and a hot MCT sheet at Langtang. BT = Bhanuwa Thrust; MCT = Main Central Thrust; MT = Munsiari Thrust; ST = Sinuwa Thrust.

I able 4.1	Mineral Assemblage:	s For Kocks from	the L	'esser	HIM	alayan 2	equen	ce, AI	nnapurna	Kegioi	1.				
Sample	Rock Unit	Rock Type	ΡI	Bt	Grt	M_{S}	Hbl	Chl	Kv/Sil	Rut	Mnz	Ilm]	ltn Tu	ir Othe	
AS01-47	LLHS Kuncha	qtzite/phyllite	X	Х		X		X						aln	
AS01-46	LLHS Kuncha	phyllite	X	X	X	Х		X						aln, zr	u
AS01-45a	LLHS Kuncha	quartzite	X	X	X	X		Х				Х		aln, zr	u
AS01-45c	LLHS Kuncha	schist	Х	X		X							×		
AS01-45d	LLHS Kuncha	schist	X	Х				Х							
AS01-44a	LLHS Kuncha	schist	X	Х	X	Х		Х					X	zrn	
AS01-44b	LLHS Kuncha	quartzite	X	Х	X	X		Х							
AS01-43a	LLHS Kuncha	schist		X	X	X						Х	X	aln	
AS01-43b	LLHS Kuncha	schist	×	X		X		Х				Х	×		
AS01-42	LLHS Ulleri	gneiss	X	Х		X									
AS01-42b	LLHS Ulleri	gneiss	Х	Х		X					Хh			aln	
AS01-41a	LLHS Kuncha	schist			X	Х		X				Х		cld	
AS01-41b	LLHS Kuncha	schist			X	X		Х				X		cld	
AS01-40a	LLHS Kuncha	schist		Х		X		Х				X			
AS01-40b	LLHS Kuncha	amphibolite	X	X			X	Х				Х		ZTN	
AS01-40c	LLHS Kuncha	amphibolite	×	X			X	Х				Х		ZTN	
AS01-40d	LLHS Kuncha	quartzite				X									
AS01-40e	LLHS Kuncha	schist	X	Х	X	X		Х			Xh	X			
AS01-40f	LLHS Kuncha	schist	X	Х	X	X		Х				X	X		
AS01-39	ULHS Dhading	marble				X								cal, ep	
AS01-38a	LLHS Kuncha	schist	×	X	×							X	×	aln	
AS01-38b	LLHS Kuncha	schist	Х	Х	Х	X						Х	×		
AS01-37a	ULHS Dhading	quartzite	X			Х							×	cal	
AS01-37b	ULHS Dhading	phyllite				Х								cal	
AS01-37c	ULHS Dhading	schist	×	X		Х						Х	×	aln, zr	u
AS01-14a	ULHS Benighat	schist		Х	Х	X							×	aln, zr	u
AS01-36a	ULHS Malekhu	schist	×	Х	×	X						X	×	aln, xr	o, zrn
AS01-36b	ULHS Malekhu	schist	Х	Х	Х	Х						Х			
Note: sam _l	ples are ordered from	n lowest structura	l leve	l to h	ighest	; all san	nples c	contain	n quartz a	nd apa	itite				
Xh: monaz	rite is not prograde, b	out rather a produ	lict of]	lydro	therm	al altera	ation								
LLHS: Lo	wer Lesser Himalaya	an Sequence; UL	HS: U	pper	Lesse	r Himal	layan S	sequer	lce						

			1	t	int.	1s	lbl	hl	y/Sil	ut	Inz	ш	tn	ur	other
#	Unit	Rock Type	Р	В	9	2	Ξ	0	K	К	2	П	Τ	Τ	0
35	Fm Ia	schist	Х	Х		Х					Х				zrn
15a	Fm Ia	schist	Х	Х	Х		Х			Х					ep
15b	Fm Ia	schist	Х	Х					ky	Х	Х			Х	xno, zrn
15c	Fm Ia	quartzite	Х	Х	Х	Х				Х					
33a	Fm Ia	quartzite	Х	Х	Х	Х									czo
33b	Fm Ia	schist	Х	Х	Х	Х				Х	Х				zrn
16a	Fm Ib	gneiss	Х	Х	Х	Х				Х	Х				par, xno
16b	Fm Ib	gneiss	Х	Х	Х	Х				Х	Х				par, zrn
16c	Fm Ib	gneiss	Х	Х	Х	Х			ky	Х	Х			Х	st
17a	Fm Ib	quartzite	Х	Х	Х							Х			aln, zrn
17b	Fm Ib	schist	Х	Х	Х	Х					Х	Х			
18a	Fm Ib	schist	Х	Х	Х	Х			ky		Х			Х	
18b	Fm Ib	schist	Х	Х	Х	Х				Х	Х				zrn
18d	Fm Ib	gneiss	Х	Х	Х				ky	Х	Х				xno, zrn
19a	Fm Ib	calc-silicate		Х	Х			Х						Х	cpx, czo
19b	Fm Ib	quartzite	Х	Х		Х					Х			Х	ksp, zrn
19c	Fm Ib	gneiss	Х	Х	Х	Х				Х	Х				st, zrn
20a	Fm Ic	gneiss	Х	Х	Х				ky	Х	Х				st, zrn
20b	Fm Ic	leucosome	Х	Х	Х				ky	Х		Х			
20c	Fm Ic	gneiss	Х	Х	Х	Х			ky	Х					
21	Fm Ic	gneiss	Х	Х	Х	Х									
22a	Fm Ic	schist	Х	Х	Х	Х			ky		Х				zrn
22b	Fm Ic	schist	Х	Х					ky+sil*						
22c	Fm Ic	schist	Х	Х					ky		Х			Х	
32	Fm Ic	gneiss	Х	Х	Х				ky	Х	Х			Х	zrn
23	Fm Ic	quartzite	Х	Х	Х	Х					Х				
24a	Fm Ic	gneiss	Х	Х	Х	Х									ep, zrn
24b	Fm Ic	quartzite	Х	Х	Х	Х									zrn
24c	Fm Ic	quartzite	Х	Х	Х	Х									zrn
25a	Fm II	leucogranite	X	Х	X	Х							••		ksp, zrn
25b	Fm II	amphibolite	Х		Х		Х					Х	Х		cpx, zrn
25c	Fm II	quartzite	X		•••		X						••		
25	Fm II	amphibolite	X		X		X						Х		cpx
25e	Fm II	amphibolite	X		Х		Х	Х					••		cpx, ep, zrn
25f	Fm II	amphibolite	X					Х					X		cpx
31a	FmII	quartzite	X	X	X	Х	X					Х	X		ep, ksp, zrn
316	Fm II	calc-silicate	X	X	Х		X						X		cal, cpx, ksp, zrn
31c	Fm II	calc-silicate	X	X			X						X		cal, cpx, zrn
30a	Fm II	calc-silicate	X	Х			X						X		cal, cpx, ep, ksp, zrn
30b	Fm II	calc-silicate	X	37			X						X		cal, cpx, ep, ksp, zrn
29	Fm II	calc-silicate	X	X			X						X		cal, cpx, czo
29-1	Fm II	calc-silicate	X	X	Х		Х						Х		cal, cpx, czo
28	Fm II	quartzite	X	X						Х	X			Х	zrn
27	Fm III	orthogneiss	Х	X		Х					Х				xno, zrn
26a	Fm II	marble		X									X		cal, cpx, zrn
26b	Fm II	marble		Х									X		cal, cpx, zrn

Table 4.2. Mineral Assemblages for Rocks from the Greater Himalayan Sequence, Annapurna Region.

Note: samples are ordered from lowest structural level to highest; all samples contain quartz and apatite Sample numbers follow the format AS01-xx

*: sillimanite is locally developed on plagioclase rims, and not a regional feature

	40e 40e	3	1 1	222 224	1 2.962 2.976	6 0.003 0.003	6 2.029 2.046	6 0.166 0.161	8 0.322 0.320	3 0.217 0.228	8 2.321 2.263	3 0.001 0.001		8 100 00 100 80		3 0./6/ 0./61	8 0.106 0.108	5 0.055 0.054	4 0.072 0.077	1 0.933 0.934				le	e 4	le 4 17.7	le 17.7
	Je 40e	2	1	34 135	57 2.96	0.00	024 2.02	49 0.16	376 0.35	252 0.25	263 2.24	0.00 0.00		100 4		/44 0./4	24 0.11	0.05	83 0.08	38 0.93				e 40	e 40	e 40 7.7 1	e 40 7.7 1
	41b 4(4	-	143 13	2.994 2.9	0.008 0.0	2.023 2.0	0.165 0.1	0.438 0.3	0.100 0.2	2.256 2.2	0.002 0.0	ı	99 01 90		0./07 0./	0.148 0.1	0.056 0.0	0.034 0.0	0.932 0.9				40	31	$ \begin{array}{c} 40 \\ 31 \\ 3.2 \\ 1 \end{array} $	31 31 32 1
	41b	4	1	142	2.996	0.006	2.016	0.174	0.447	0.108	2.242	ı	·	90 55		CC/.0	0.150	0.059	0.036	0.928				40e	40e 29	40e 29 18	40e 29 18
	41b	4	-	141	2.965	0.002	2.044	0.176	0.417	0.108	2.298	0.002	,	08 47		0./66	0.139	0.059	0.036	0.929				44a	44a 32	44a 32 18.7	44a 32 18.7
	43a	1	-	147	2.937	0.004	2.012	0.112	0.462	0.164	2.357	0.008	0.001	00 24		0./02	0.149	0.036	0.053	0.955				44a	44a 31	44a 31 18.2	44a 31 18.2
	43a	1	-	7	2.939	0.004	2.068	0.097	0.553	0.141	2.222	ı	0.001	100.65		0./3/	0.183	0.032	0.047	0.958				4a	4a 16	4a 16 18.3	4a 16 18.3
4	43a	1	1	9	2.940	0.004	2.058	0.097	0.532	0.130	2.267	·	ı	100.84		0./49	0.176	0.032	0.043	0.959				4	4	4 L	4 C.
	44a	7	1	86	2.987	0.003	2.020	0.108	0.489	0.191	2.199	0.003	ı	00 43		0.730	0.164	0.036	0.064	0.953				45a	45a 36	45a 36 17	45a 36 17
	44a	1	1	6	2.977	0.004	2.000	0.111	0.464	0.157	2.306	ı	ı	100.68		96/.0	0.153	0.037	0.052	0.954	sus.			45a	45a 32	45a 32 17.5	45a 32 17.5
	45a	1	-	9	2.949	0.017	2.079	0.126	0.575	0.408	1.836	0.012	0.002	59 00		0.624	0.195	0.043	0.139	0.936	12 oxyge			45a	45a 30	45a 30 14.6	45a 30 14.6
	45a	1	-	5	2.974	0.025	2.073	0.120	0.530	0.399	1.842	0.005	0.001	100 76		0.63/	0.183	0.041	0.138	0.939	basis of			5	3.0	5 3 17.4	5 3 17.4
	45a	4	-	С	2.982	0.007	2.041	0.114	0.718	0.451	1.676	0.001	0.001	90 48		000.0	0.243	0.038	0.152	0.936	hydrous			4	40	9 17 9 1	9 17 9 1
4	46	1	-	7	9 3.025	0 0.002	4 1.967	0 0.118	1 0.471	6 0.136	0 2.261	5 0.018	2 0.001	1 99.86		4 U./2	0 0.158	2 0.040	4 0.045	6 0.950	to an an			46	46 169	46 169 15.	46 169 15.
	46	ŝ	7	2	12 2.949	0.020	4 1.99	4 0.130	5 0.49	1 0.160	1 2.280	0.00(0.00	18 00 77		8 0./4	1 0.16	1 0.042	0 0.05	8 0.940	malized			46	46 167	46 167 16.6	46 167 16.6
t	46	Э	1	2	2.99	00.00	1.99	0.12	0.45	0.15	2.28	00.00	0.00	101 2		C/.0	0.15	0.04	0.05) 0.94	s are nor	بامدم	Jeb1		si	si o	si o
Garne	Sample	Garnet	Traverse	Analysis	Si	Τi	Al	Mg	Са	Mn	Fe	Na	K W/+0/	Total	10101	Alm	Grs	Prp	Sps Fe/	(Fe+Mg)	Cation	Dlamo	1 Idgiv	Sample	<u>Sample</u> Analys	Sample Analys X _{An} (%)	Sample Analys X _{An} (%)

Table 4.3 Mineral Compositions from the Lesser Himalayan Duplex

52

	4a	lh.	18	.128	.010	.769	.975	.007	.016	.070		.005		7.96	.671	40e	14	36	5.254	0.034	.675	.184	007	0.027	602	0.023	0.157		8.97	0.638
	4		1	1 5.	9 0.	2 5.	5 2.	0.0	0.	8 6.	4	3 0.		5 8	8	0		1 –	54 5	0 0	576 5	26 3	05 0	135 0	82 5	0 14 0	0 00		.63 8	27 0
	44a	M_{S}	5	3.07	0.01	2.80	0.05	00.00	ı	0.07	0.10	0.86		95.1	0.58	406	Ch.	20	5.2	0.0	5.6	3.3	0.0	0.0	5.5	0.0	0.0		88.	0.6
	44a	M_{S}	4	3.124	0.012	2.683	0.099	'	0.000	0.105	0.105	0.890		95.45	0.516	40e	Ms	38	3.092	0.014	2.731	0.092	•	0.001	0.104	0.179	0.814		94.52	0.531
	44a	Bt	22	2.729	0.101	1.657	0.766	0.004	0.005	1.608	0.014	0.932		95.56	0.677	40e	Ms	27	3.119	0.013	2.679	0.111		0.002	0.115	0.125	0.861		95.18	0.508
	44a	Bt	14	2.699	0.057	1.732	0.784	0.006	0.008	1.606	0.023	0.949		94.44	0.672	40e	Bt	22	2.711	0.096	1.669	0.891	0.006	0.003	1.523	0.030	0.889		94.44	0.631
	44a	Bt	13	2.719	0.113	1.655	0.762	0.007	0.006	1.613	0.014	0.915		95.68	0.679	40e	Bt	21	2.711	0.104	1.691	0.881	0.002	0.006	1.492	0.016	0.890		95.60	0.629
	44a	Bt	12	2.708	0.111	1.652	0.789	0.004	0.007	1.615	0.018	0.920		95.06	0.672	40e	₿ţ	6	2.718	0.083	1.680	0.906	0.004	0.007	1.503	0.018	0.894		95.71	0.624
	45a	Chl	28	6.305	0.114	4.794	2.823	0.027	0.033	4.646	0.002	0.882		88.30	0.622	41b	Ctd	35	2.015	0.002	3.924	0.286	0.002	0.029	1.760	0.003	0.001		93.13	0.860
	45a	M_{S}	35	3.077	0.055	2.731	0.061	·	,	0.096	0.062	0.903		94.58	0.614	41b	Ctd	7	2.029	0.001	3.953	0.256	0.001	0.023	1.731	·	0.001		92.26	0.871
	45a	Bt	36	2.778	0.131	1.558	0.949	0.006	0.007	1.414	0.021	0.917		93.99	0.598	41b	Chl	17	5.232	0.012	5.906	3.680	0.003	0.018	4.950				84.55	0.574
	45a	Bt	29	2.795	0.138	1.580	0.915	0.001	0.007	1.382	0.024	0.892		94.36	0.601	41b	Chl	Ξ	5.146	0.013	5.960	3.725	0.004	0.024	4.983	0.003	0.007		84.57	0.572
	45a	Bt	25	2.779	0.122	1.605	0.935	0.001	0.008	1.371	0.015	0.938		93.38	0.594	41b	Chl	8	5.251	0.012	5.924	3.589	0.012	0.016	4.901	0.112	0.032		85.50	0.577
	45a	Bt	24	2.750	0.119	1.639	0.926	0.006	0.008	1.384	0.032	0.925		93.45	0.599	41b	Ms	40	3.071	0.008	2.811	0.059	0.001		0.077	0.229	0.750		93.97	0.567
	46	Chl	62	5.195	0.016	5.803	3.194	0.002	0.017	5.657		0.007		86.32	0.639	41b	Ms	39	3.067	0.013	2.814	0.055	0.003		0.081	0.223	0.738		94.24	0.596
	46	Chl	61	5.503	0.052	5.591	2.965	0.001	0.016	5.346	0.012	0.340		87.69	0.643	43a	Ms	39	3.068	0.012	2.796	0.060	0.002	0.002	0.109	0.112	0.836		95.15	0.644
	46	Ms	64	0.828	2.765	0.011	0.001	3.068	0.156	0.079	0.120	0.002		94.78	0.602	43a	Bt	53	2.729	0.112	1.694	0.662	0.001	0.004	1.637	0.010	0.938		94.75	0.712
rals	46	Ms	47	3.142	0.016	2.668	0.084		1	0.107	0.097	0.886		94.11	0.561	43a	Bt	52	2.722	0.110	1.658	0.686	0.003	0.004	1.682	0.016	0.933		94.79	0.710
ied e Miner	46	Bt	159	2.710	0.075	.694	.828 (0.003	.009	.589 (.008 (.912		3.49	.657 (43a	Bt	51	2.695	111.0	999.	.679 (000.	007	.720	0.010	.956 (3.82	.717 (
Continu r Silicat	46	Bt	46	.793 2	.107 (.654 1	.782 (.003 (.004	.493 ì	.012 () 098.		4.57 5	.656 (t3a ,	Bt	38	.722 2	.113 (.647]	.691 (.005 (.005 (.681 1	.006 (.953 (3.16 5) 601.
e 4.3 (r Majoi	le ^z	ral	sis 1	6	.0	1.	.0	0	0	1.	0	0		9	1g) 0.	-le 4	ra	sis	5	0	1	0	0	0	1	0	0		6	1g) 0.
Tabl Othe	Samp	Mine	Analy	Si	Ti	Al	Mg	Са	Mn	Fe	Na	K	Wt%	Total Fe/	(Fe+N	Samp	Mine	Analy	Si	Ti	Al	Mg	Са	Mn	Fe	Na	К	Wt%	Total Fe/	(Fe+N

Table 4.3 Continued Other Maior Silicate Min

Garnet									
Sample	38b	38b	38b	38b	36a	36a	36a	36b	36b
Garnet	1	1	2	2	2	2	2	1	1
Traverse	1	1	1	1	1	1	1	1	1
Analysis	11	12	53	57	325	328	336	7	ŝ
Si	2.925	2.942	2.950	2.965	2.952	2.958	2.907	2.989	2.967
Ti	0.006	0.006	0.002	0.004	0.003	0.005	0.002	0.002	0.002
Al	2.078	2.078	2.070	2.054	2.049	2.065	2.029	2.089	2.050
Mg	0.199	0.189	0.182	0.198	0.202	0.198	0.204	0.276	0.288
Ca	0.595	0.602	0.623	0.586	0.562	0.570	0.568	0.404	0.422
Mn	0.040	0.039	0.026	0.027	0.269	0.266	0.266	0.015	0.022
Fe	2.185	2.154	2.158	2.169	1.985	1.943	2.099	2.183	2.250
Na	0.006	0.008	0.003	0.001	ı		0.001	0.011	0.010
K	ı	ı	·		ı	·	ı	0.003	0.001
Wt% Total	99.74	98.73	100.11	99.95	101.96	101.67	98.93	101.63	100.04
Alm	0.724	0.722	0.722	0.728	0.658	0.653	0.669	0.759	0.754
Grs	0.197	0.202	0.208	0.197	0.186	0.191	0.181	0.140	0.142
Prp	0.066	0.063	0.061	0.066	0.067	0.067	0.065	0.096	0.097
Sps	0.013	0.013	0.009	0.009	0.089	0.089	0.085	0.005	0.007
Fe/(Fe+Mg)	0.917	0.919	0.922	0.916	0.908	0.908	0.911	0.888	0.886
Cations are nor	malized to a	n anhydrou	is basis of 1	2 oxygens.					
Plagioclase									
Sample	38b	38b	38b	36a	36a	3(бb	36b	36b
Analysis	65	99	67	48	50	1	7	13	14
${ m X}_{ m An}(\%)$	17.0	17.3	3 14.	6 15.	9	16.2	14.5	14.1	13.4
Wt% Total	102.63	102.8	3 101.6	8 102.3	8 10]	.80	99.95	100.00	99.19

Table 4.4 Mineral Compositions of Rocks from the Lesser Himalayan Sequence, Munsiari Thrust.

54

	6b	$_{\rm Is}$	4	074)23	766	104	<u> </u>	000	381	157	786	.23	440	
	3(V	7	3.(0.0	5.7	1 0.1	0.0	0.0	0.0	0.]	0.0	93	0.4	
	36b	$M_{\rm S}$	0	3.083	0.018	2.743	0.114	0.001	0.00	0.090	0.126	0.833	93.62	0.442	
	36b	Bt	10	2.741	0.115	1.660	1.143	0.004	0.005	1.199	0.015	0.880	93.23	0.512	
	36b	Bt	S	2.752	0.121	1.637	1.156	0.000	0.004	1.181	0.020	0.895	93.39	0.505	
	36b	Bt	1	2.749	0.115	1.653	1.135	0.001	0.006	1.189	0.016	0.906	93.39	0.512	
	36a	$\mathbf{M}_{\mathbf{S}}$	34	3.153	0.024	2.597	0.150	ı	0.002	0.116	0.063	0.902	95.81	0.434	
	36a	$\mathbf{M}_{\mathbf{S}}$	43	2.738	0.116	1.645	1.088	0.004	0.011	1.271	0.015	0.889	97.55	0.539	
	36a	Bt	42	2.734	0.103	1.645	1.118	0.001	0.006	1.282	0.011	0.890	97.16	0.534	
	36a	Bt	41	2.747	0.095	1.653	1.094	0.002	0.009	1.278	0.013	0.896	97.26	0.539	
	36a	Bt	40	2.757	0.104	1.667	1.066	ı	0.008	1.256	0.008	0.888	97.61	0.541	
	38b	$M_{\rm S}$	79	3.133	0.021	2.646	0.140	0.001	0.002	0.109	0.093	0.853	96.50	0.438 NS	3
	38b	$M_{\rm S}$	64	3.130	0.023	2.652	0.131	0.001	ı	0.103	0.102	0.864	96.54	0.439 0xvoet	
	38b	Bt	57	2.759	0.115	1.629	1.044	0.006	0.006	1.320	0.015	0.853	97.38	0.558 rous 11	
nerals	38b	Bt	56	2.737	0.105	1.607	1.075	0.006	ı	1.380	0.013	0.875	97.14	0.562 n anhvd	י שנייוש וו
cate Mi	38b	Bt	55	2.748	0.113	1.687	1.005	0.002	0.002	1.280	0.014	0.908	97.41	0.560 red to a	
ajor Sili	38b	Bt	53	2.748	0.110	1.682	1.020	0.002	0.002	1.278	0.014	0.906	96.79	0.556 0.rmaliz	
Other Ma	Sample	Mineral	Analysis	Si	Ti	Al	Mg	Ca	Mn	Fe	Na	K Wt%	Total Fe/	(Fe+Mg) Cations r	

Table 4.4 Continued Other Maior Silicate Mi

Garnet													
Sample	15a	15a	15a	15a	15c	15c	15c	33a	33a	33a	33b	33b	33b
Garnet	1	2	2	2	1	1	2	1	1	1	1	2	3
Traverse	1	-	1	2	1	1	1	1	1	1	1	1	1
Analysis	4	12	13	13	7	8	116	ς	4	9	8	10	8
Si	3.011	3.009	2.998	3.010	2.958	2.973	2.968	2.946	2.934	2.942	3.015	3.023	3.022
Ti	,	I	0.002	0.001	·	ı	0.001	0.003	0.004	0.003	ı	0.001	0.001
\mathbf{AI}	1.985	1.975	1.983	1.976	2.026	2.005	2.041	2.040	2.057	2.042	1.945	1.933	1.923
Mg	0.567	0.528	0.530	0.533	0.529	0.535	0.537	0.239	0.231	0.244	0.591	0.595	0.593
Ca	0.573	0.779	0.780	0.748	0.231	0.240	0.238	1.218	1.228	1.222	0.176	0.160	0.182
Mn	0.175	0.144	0.151	0.162	0.098	0.098	0.086	0.230	0.237	0.236	0.067	0.086	0.076
Fe	1.684	1.566	1.567	1.569	2.186	2.173	2.139	1.352	1.339	1.342	2.216	2.209	2.214
Na	0.005	0.001	0.000	0.005	,	ı	ı	0.004	0.008	0.006	0.005	0.004	0.010
K		0.003		0.001	,	ı	ı	0.001	0.002	0.003	0.001	ı	0.002
Wt% Total	101.34	101.25	101.64	101.71	101.09	100.65	100.17	100.74	99.91	100.30	100.76	101.34	101.08
Alm	0.561	0.519	0.518	0.521	0.718	0.713	0.713	0.445	0.441	0.441	0.727	0.724	0.722
Grs	0.191	0.258	0.258	0.248	0.076	0.079	0.079	0.401	0.405	0.402	0.058	0.052	0.059
Prp	0.189	0.175	0.175	0.177	0.174	0.176	0.179	0.079	0.076	0.080	0.194	0.195	0.193
Sps	0.058	0.048	0.050	0.054	0.032	0.032	0.029	0.076	0.078	0.078	0.022	0.028	0.025
Fe/(Fe+Mg)	0.748	0.748	0.747	0.746	0.805	0.802	0.799	0.850	0.853	0.846	0.789	0.788	0.789
Cations are no	rmalized to	an anhydrc	ous basis of	12 oxygens									
Plagioclase													

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5 Miner
Table 4.

Sample15a15a15a15a15a15c15c15c15c33a33a33b33b33b33bAnalysis19119419819925264445485052266267269 $X_{An}(\%)$ 27.329.027.426.613.914.214.814.626.625.725.09.79.410.0Wt%Total99.7690.1699.1698.5298.69100.50100.72100.97100.8399.95

Other Major Si	licate Min	erals									
Sample	15a	15a	15a	15a	15a	15a	15a	15c	15c	15c	15c
Mineral	Bt	Bť	Bť	Bt	Hbl	Hbl	Hbl	Bt	Bť	Bt	Bt
Analysis	235	236	237	242	295	296	297	20	21	23	37
Si	2.790	2.777	2.808	2.805	6.328	6.433	6.364	2.763	2.789	2.803	2.783
Ti	0.070	0.066	0.081	0.049	0.082	0.061	0.072	0.090	0.095	0.094	0.110
Al	1.524	1.505	1.470	1.530	2.694	2.612	2.624	1.636	1.678	1.574	1.612
Mg	1.801	1.827	1.789	1.790	2.581	2.633	2.638	1.321	1.253	1.338	1.304
Ca	ı	0.001	0.003	0.003	1.784	1.704	1.785	ı	0.006	0.003	ı
Mn	0.003	0.003	0.006	0.006	0.021	0.019	0.025	0.003	0.000	ı	ı
Fe	0.766	0.804	0.791	0.768	1.412	1.416	1.416	1.060	0.993	1.051	1.032
Na	0.070	0.055	0.045	0.056	0.547	0.563	0.563	0.058	0.063	0.041	0.047
K	0.783	0.786	0.816	0.802	0.137	0.081	0.096	0.853	0.866	0.864	0.876
Wt% Total	93.03	92.91	93.13	92.73	97.20	96.60	97.03	94.79	93.06	93.11	94.14
Fe/(Fe+Mg)	0.298	0.305	0.306	0.300	0.354	0.350	0.349	0.445	0.442	0.440	0.442
Sample	15c	15c	33a	33a	33a	33a	33b	33b	33b	33b	
Mineral	$M_{\rm S}$	M_{S}	Bt	Bť	Bt	$M_{\rm S}$	Bt	Bt	Bt	M_{S}	
Analysis	40	42	42	43	46	45	246	253	255	285	
Si	3.091	3.093	2.797	2.791	2.767	3.162	2.759	2.760	2.752	3.110	
Ti	0.044	0.037	0.152	0.143	0.139	0.035	0.097	0.086	0.094	0.046	
Al	2.690	2.714	1.566	1.567	1.589	2.532	1.600	1.618	1.570	2.666	
Mg	0.113	0.098	1.225	1.243	1.250	0.203	1.410	1.438	1.441	0.128	
Ca	0.002		0.001	0.009	0.000	0.001	·	ı	0.002	0.002	
Mn	0.000		0.014	0.015	0.017	0.001	0.005	0.001	0.004	0.000	
Fe	0.079	0.081	1.053	1.052	1.067	0.117	1.026	1.000	1.054	0.077	
Na	0.232	0.231	0.021	0.034	0.023	0.048	0.055	0.036	0.047	0.232	
K	0.769	0.746	0.901	0.889	0.918	0.924	0.839	0.855	0.855	0.730	
Wt% Total	94.32	95.02	94.85	93.26	94.31	93.26	94.05	93.88	93.14	93.39	
Fe/(Fe+Mg)	0.412	0.450	0.462	0.458	0.460	0.366	0.421	0.410	0.422	0.375	
Cations normal	ized to an	anhydrou	s 11 (biotite	e and musc	ovite) or 23	3 (hornblen	de) oxygen	S			

Table 4.5 Continued.

Garnet																	
Sample	16a	1 6a	16a	16b	16b	17b	17b	17b	18b	18b	18b	18d	18d	18d	19c	19c	19c
Garnet	2	2	2	2	2	1	1	1	1	1	1	1	1	1	1	1	1
Traverse	1		-	1	-1	1	-	-	-		1	-	-		-	-	-
Analysis	ς	4	5	7	6	16	18	22	14	16	56	22	23	32	51	54	57
Si	3.009	2.985	2.992	2.971	2.968	2.958	2.942	2.928	2.954	2.956	2.943	2.913	2.934	2.910	2.920	2.921	2.918
Ti	0.000	0.001	0.000	0.002	0.002		0.000	0.000	0.001	0.000	0.001	·	ı		0.000	0.001	
Al	2.003	2.009	2.000	2.053	2.061	2.069	2.049	2.048	2.047	2.041	2.047	2.072	2.059	2.080	2.082	2.056	2.060
Mg	0.651	0.650	0.650	0.691	0.694	0.607	0.611	0.622	0.492	0.502	0.493	0.686	0.684	0.683	0.690	0.692	0.695
Са	0.187	0.184	0.181	0.183	0.183	0.329	0.341	0.332	0.434	0.410	0.437	0.122	0.123	0.130	0.168	0.170	0.158
Mn	0.165	0.173	0.178	0.055	0.060	0.131	0.131	0.128	0.149	0.158	0.165	0.095	0.094	0.092	0.096	0.098	0.095
Fe	1.964	1.997	2.000	2.045	2.030	1.909	1.952	1.984	1.941	1.953	1.945	2.162	2.140	2.154	2.082	2.111	2.122
Na	0.018	0.018	0.011	0.001	0.003	0.009	0.010	0.010	0.006	0.005	0.004	0.001	0.004	0.003	0.003	0.001	0.006
K	0.003	0.003	0.002	0.002	0.001		0.001			0.001	0.001		0.001	0.001	0.001	0.001	
Wt%																	
Total	100.23	102.25	101.39	99.48	99.52	101.62	100.17	100.00	100.57	100.46	100.16	99.22	98.54	99.11	98.32	98.55	98.74
Alm	0.662	0.665	0.665	0.688	0.684	0.641	0.643	0.647	0.644	0.646	0.640	0.706	0.704	0.704	0.686	0.687	0.691
Grs	0.063	0.061	0.060	0.061	0.062	0.111	0.112	0.108	0.144	0.136	0.144	0.040	0.041	0.043	0.055	0.055	0.051
Prp	0.219	0.216	0.216	0.232	0.234	0.204	0.201	0.203	0.163	0.166	0.162	0.224	0.225	0.223	0.227	0.225	0.226
$\operatorname{Sps}_{\operatorname{Fo}'}$	0.056	0.058	0.059	0.019	0.020	0.044	0.043	0.042	0.049	0.052	0.054	0.031	0.031	0.030	0.032	0.032	0.031
(Fe+Mg)	0.751	0.754	0.755	0.747	0.745	0.759	0.762	0.761	0.798	0.795	0.798	0.759	0.758	0.759	0.751	0.753	0.753
Cations ar	e normal	ized to ¿	an anhydr	rous basi:	s of 12 ox	vygens.											
Plagioclas	e																
Sample	16a	16a	16a	16b	16b	16b	17b	17b	17b	18b	18b	18d	18d	18d	19c	19c	19c
Analysis	ς	9	7	56	131	136	15	17	18	30	35	99	68	72	84	92	93
${ m X}_{ m An}(\%)$ Wt%	12.2	14.1	12.3	12.3	12.5	13.3	20.3	19.3	20.7	30.4	28.3	14.0	15.4	13.6	17.2	15.9	16.9
Total	100.55	99.65	100.83	100.94	101.03	100.69	99.67	99.19	100.85	100.60	100.95	98.90	99.50	99.39	100.73	100.96	101.01

Table 4.6 Mineral Compositions of Rocks from the Greater Himalayan Sequence Formation 1b.

58

100.85

99.65

Other Major	Silicate N	linerals														
Sample	16a	16a	16a	16a	16a	16b	16b	16b	16b	16b	16b	17b	17b	17b	17b	17b
Mineral	Bt	Bt	Bt	M_{S}	Par	Bt	Bt	Bt	Bt	M_{S}	Par	Bt	Bt	Bt	M_{S}	M_{S}
Analysis	11	15	20	17	19	39*	47*	101^{*}	107*	123	42	ŝ	13	57	9	56
Si	2.794	2.814	2.799	3.107	2.994	2.789	2.783	2.768	2.776	2.903	2.942	2.735	2.755	2.775	3.063	3.102
Ti	0.093	0.090	0.088	0.031	0.011	0.078	0.096	0.056	0.084	0.081	0.013	0.107	0.116	0.117	0.024	0.041
Al	1.564	1.561	1.583	2.659	2.937	1.645	1.637	1.633	1.656	2.299	3.021	1.662	1.684	1.625	2.771	2.698
Mg	1.386	1.393	1.385	0.094	0.011	1.421	1.402	1.456	1.456	0.580	0.016	1.398	1.355	1.372	0.115	0.114
Ca	0.001		0.002	0.001	0.020	0.001	0.002	0.004	0.001	0.016	0.017	0.007	0.002	0.001	0.001	
Mn	0.005	0.002	0.004		0.000	0.006	0.003	0.008	0.004	0.004	0.000	0.004	0.005	0.003	0.002	·
Fe	1.027	0.996	1.003	0.146	0.050	0.944	0.952	1.029	0.979	0.514	0.039	0.959	0.934	0.950	0.074	0.068
Na	0.045	0.042	0.047	0.331	0.894	0.046	0.041	0.026	0.039	0.402	0.810	0.033	0.027	0.036	0.124	0.113
К	0.879	0.874	0.868	0.659	0.108	0.853	0.866	0.761	0.847	0.539	0.161	0.876	0.848	0.868	0.834	0.857
Wt% Total	94.83	94.76	94.89	95.04	95.41	94.19	93.37	93.55	93.92	94.01	94.95	93.47	93.32	93.81	93.16	93.22
Fe/(Fe+Mg)	0.426	0.417	0.420	0.607	0.816	0.399	0.404	0.414	0.402	0.470	0.712	0.407	0.408	0.409	0.393	0.373
Original Bt																
Fe/(Fe+Mg)						0.417	0.422	0.430	0.419							
Sample	18b	18b	18b	18b	18d	18d	18d	18d	18d	19c	19c	19c	19c	19c	19c	
Mineral	Bt	Bt	Bt	Ms	Bt	Bt	Bt	Ms	Ms	Bt	Bt	Bt	Bt	Ms	Ms	
Analysis	21*	23*	26^{*}	22	42*	44*	45*	46	69	39*	40*	82*	83*	LL LL	81	
Si	2.775	2.759	2.770	3.076	2.767	2.766	2.758	3.088	3.114	2.726	2.727	2.761	2.774	3.132	3.102	
Ti	0.145	0.150	0.158	0.059	0.175	0.153	0.151	0.041	0.025	0.124	0.119	0.121	0.115	0.062	0.061	
Al	1.615	1.625	1.619	2.705	1.586	1.607	1.631	2.744	2.770	1.669	1.683	1.678	1.637	2.674	2.679	
Mg	1.245	1.243	1.238	0.120	1.310	1.320	1.317	0.083	0.076	1.350	1.333	1.297	1.327	0.084	0.104	
Ca	0.008	0.005	0.001	0.002	ı	,	0.000	,	'	0.001	0.001	0.002	0.002	0.003	0.002	
Mn	0.002	0.004	0.004	0.000	0.002	,	0.001	,	0.002	0.000	0.005	0.007	0.002	ı	0.000	
Fe	1.045	1.053	1.031	0.076	0.993	0.998	0.984	0.062	0.057	1.001	0.994	0.982	1.013	0.056	0.064	
Na	0.032	0.034	0.024	0.120	0.055	0.056	0.065	0.255	0.161	0.043	0.050	0.044	0.039	0.161	0.168	
K	0.857	0.860	0.869	0.828	0.823	0.824	0.809	0.706	0.705	0.867	0.864	0.828	0.817	0.759	0.805	
Wt% Total	94.07	94.71	94.30	94.51	94.06	93.26	93.92	94.19	93.83	94.90	95.13	95.48	95.46	94.98	94.19	
Fe/(Fe+Mg)	0.456	0.459	0.454	0.390	0.431	0.431	0.428	0.425	0.425	0.426	0.427	0.431	0.433	0.399	0.382	
Original Bt																
Fe/(Fe+Mg)	0.460	0.462	0.458		0.436	0.436	0.433			0.430	0.431	0.437	0.435			
Cations norm	alized to	an anhydi	rous 11 or	xygens												
* Biotite com	position (corrected	for ReNI	'R's												

Table 4.6. Continued

0c 20c
2 2
1 1
31 35
6 2.956 2.954
0 0.001 -
7 2.045 2.040
0 0.663 0.661
8 0.213 0.216
4 0.112 0.111
5 2.028 2.040
0.005 0.005
0.001 -
4 99.97 99.90
5 0.672 0.674
3 0.071 0.071
5 0.220 0.218
7 0.037 0.037
5 0.754 0.755
oxygens.
20c 32
73 44
23.6 18.9
100.74 99.05

Table 4.7 Mineral Compositions of Rocks from the Greater Himalayan Sequence Formation 1c.

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Sample	20a	20a	20a	20a	20c	20c	20c	20c	20c	32	32	32
Mineral	Bt	Bt	Bt	Bt	Bt	Bt	Bt	M_{S}	Ms	Bt	Bt	Bt
Analysis	25*	27*	28*	29*	54*	*09	61^{*}	59	62	38*	41^{*}	51*
Si	2.641	2.802	2.803	2.779	2.731	2.730	2.746	3.105	3.085	2.715	2.745	2.738
Ti	0.060	0.090	0.082	0.074	0.162	0.169	0.138	0.071	0.066	0.105	0.128	0.119
Al	1.728	1.591	1.593	1.579	1.668	1.656	1.652	2.647	2.681	1.709	1.686	1.664
Mg	1.800	1.607	1.631	1.689	1.304	1.278	1.337	0.129	0.127	1.505	1.448	1.499
Ca	0.007	0.001	0.008	0.017	0.002	0.015	0.002	0.001	0.001	0.001	ı	0.000
Mn	0.004	0.001	0.004	0.004	0.000	0.005	0.004	ı	ı	0.001	0.004	0.004
Fe	0.885	0.797	0.819	0.856	1.001	0.958	0.998	0.071	0.070	0.851	0.824	0.841
Na	0.047	0.078	0.073	0.059	0.052	0.061	0.041	0.133	0.141	0.069	0.060	0.059
K	0.589	0.787	0.745	0.714	0.842	0.819	0.857	0.817	0.820	0.789	0.835	0.854
Wt% Total	93.27	95.33	93.80	93.40	95.13	93.69	93.53	94.58	94.77	93.28	93.38	94.40
Fe/(Fe+Mg)	0.330	0.332	0.334	0.336	0.434	0.428	0.428	0.356	0.353	0.361	0.363	0.359
Original												
Fe/(Fe+Mg)	0.337	0.340	0.343	0.344	0.457	0.451	0.450			0.369	0.371	0.368
Sample	23	23	23	23	24a	24a	24a	24a	24a	24a		
Mineral	Bt	Bt	Bt	Ms	Bt	Bt	Bt	Bt	Ms	Ms		
Analysis	10^{*}	12*	13*	0	ж Ж	\$ *	*6	12*	7	11		
Si	2.711	2.707	2.706	3.130	2.717	2.720	2.757	2.720	3.070	3.084		
Ti	0.193	0.187	0.181	0.053	0.155	0.156	0.137	0.143	0.039	0.034		
\mathbf{AI}	1.643	1.660	1.664	2.628	1.650	1.700	1.654	1.700	2.724	2.723		
Mg	1.162	1.149	1.135	0.126	1.250	1.205	1.234	1.214	0.118	0.115		
Ca	0.000	0.003	0.003	0.002	I	0.001	0.004	0.002	0.001	0.004		
Mn	0.004	0.008	0.005	0.002	0.006	0.007	0.008	0.009	I	0.002		
Fe	1.107	1.107	1.103	0.107	1.050	1.011	1.024	1.025	0.081	0.076		
Na	0.016	0.035	0.027	0.049	0.018	0.021	0.017	0.028	0.061	0.060		
K	0.933	0.930	0.911	0.864	0.931	0.930	0.916	0.926	0.930	0.907		
Wt% Total	93.23	93.24	93.90	94.03	94.05	94.84	94.18	93.29	94.18	93.66		
Fe/(Fe+Mg)	0.488	0.491	0.493	0.459	0.457	0.456	0.453	0.458	0.406	0.397		
Original												
Fe/(Fe+Mo)	0.507	0510	0512		0.16	0.16	0 7 5 0	0 163				

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Gainet			
Sample	25a	25a	25a
Garnet	2	2	2
Traverse	1	1	1
Analysis	14	16	18
Si	2.970	2.973	2.967
Ti	-	-	0.002
Al	2.013	2.013	2.016
Mg	0.285	0.280	0.278
Ca	0.305	0.333	0.362
Mn	0.327	0.319	0.325
Fe	2.119	2.098	2.068
Na	0.009	0.010	0.006
Κ	0.000	0.000	0.002
Wt% Total	99.39	99.24	99.10
Alm	0.698	0.693	0.682
Grs	0.100	0.110	0.119
Prp	0.094	0.092	0.092
Sps	0.108	0.105	0.107
Fe/(Fe+Mg)	0.881	0.882	0.881

Table 4.8 Mineral Compositions of Rocks from GHS Formation III. Garnet

Cations are normalized to an anhydrous basis of 12 oxygens.

Plagioclase

Sample	25a	25a	25a	25a
Analysis	318	319	321	322
$X_{An}(\%)$	16.6	16.2	16.8	17.5
Wt% Total	100.22	101.60	100.30	101.38

Sample	25a	25a						
Mineral	Bt	Bt	Bt	Bt	Ms	Ms	Ksp	Ksp
Analysis	304*	305*	310*	311*	309	313	324	327
Si	2.721	2.711	2.730	2.709	3.117	3.126	2.973	2.973
Ti	0.146	0.194	0.175	0.197	0.051	0.052	-	-
Al	1.596	1.545	1.559	1.582	2.616	2.602	1.041	1.040
Mg	0.855	0.857	0.861	0.844	0.100	0.109	-	-
Ca	0.001	-	0.000	0.004	0.000	0.000	0.002	0.003
Mn	0.015	0.017	0.012	0.013	0.001	0.001	-	-
Fe	1.507	1.507	1.503	1.475	0.136	0.133	0.000	-
Na	0.002	0.015	0.006	0.004	0.053	0.054	0.094	0.089
Κ	0.965	0.961	0.969	0.963	0.953	0.944	0.886	0.893
Wt% Total	94.40	95.98	94.43	94.96	93.23	93.17	100.16	100.30
Fe/(Fe+Mg)	0.638	0.637	0.636	0.636	0.576	0.549		
Original Bt								
Fe/(Fe+Mg)	0.646	0.645	0.644	0.644				

Other Major Silicate Minerals

Cations normalized to an anhydrous 11 (biotite and muscovite) or 8 (K-Feldspar) oxygens

* Biotite composition corrected for ReNTR's

Table 4.9 Thermobarometric Results from Rocks Along the Modi Khola.

Sample	T (°C)	P (kbar)
GHS Fm II (MCT)		
AS01-25a	$775 \pm 20*$	$14.0 \pm 1.0*$
Average	775	14.0
GHS Fm Ic (MCT)		
AS01-24a	$775 \pm 20*$	$13.0 \pm 0.5^*$
AS01-23	$780 \pm 20^{*}$	$13.0 \pm 0.5^*$
AS01-32	$765 \pm 20*$	$12.0 \pm 0.5^*$
AS01-20c	$765 \pm 30*$	$12.5 \pm 0.5^*$
AS01-20a	$775 \pm 20*$	$11.5 \pm 0.5*$
Average	775	12.5
GHS Fm Ib (MCT)		
AS01-19c	$745 \pm 20*$	$11.5 \pm 0.5*$
AS01-18d	$725 \pm 20*$	$11.5 \pm 1.0*$
AS01-18b	$740 \pm 20*$	$12.0 \pm 0.5*$
AS01-17b	735 ± 20	13.5 ± 0.5
AS01-16b	$725 \pm 20*$	$13.0 \pm 0.5*$
AS01-16a	740 ± 30	12.5 ± 1.0
Average	735	12.5
GHS Fm Ia (MCT)		
AS01-33b	650 ± 20	11.5 ± 1.0
AS01-33a	645 ± 20	12.0 ± 0.5
AS01-15c	670 ± 30	11.5 ± 0.5
AS01-15a	625 ± 40	12.0 ± 1.0
Average	650	12.0
LHS (MT)		
AS01-36b	575 ± 20	11.5 ± 0.5
AS01-36a	545 ± 20	10.0 ± 0.5
AS01-38b	550 ± 30	10.5 ± 1.0
Average	555	10.5
LHS (LHD)		
AS01-40e	515 ± 30	7.0 ± 1.0
AS01-41b	530 ± 50^{a}	
AS01-43a	$530\pm50^{\mathrm{a}}$	
AS01-44a	500 ± 20	7.0 ± 0.5
AS01-45a	505 ± 20	8.5 ± 1.0
AS01-46	510 ± 30	8.0 ± 0.5
Average	515	7.5

Note: All temperatures and associated errors rounded to nearest 5°C, pressures and associated errors to 0.5 kbar. All temperatures calculated via garnet-biotite thermometry except AS01-41b, which was calculated via garnet-chlorite thermometry. Pressures calculated from garnet-plagioclase-muscovite-biotite, garnet-plagioclase-kyanite-quartz, and garnet-plagioclase-biotite-quartz barometry.

a: T calculated at an assumed pressure of 8 ± 1 kbar.

*Biotite corrected for ReNTRs

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# zuu	1	1	1	1	3	3	3	3	4	4	4	4	4	5	5	5	5
Analysis	44	45	46	47	51	52	54	55	56	58	59	60	62	63	64	65	66
SiO_2	0.38	0.38	0.25	0.32	0.29	0.40	0.27	0.43	0.24	0.29	0.20	0.33	0.44	1.35	0.28	0.30	0.40
P_2O_5	30.09	30.28	30.34	30.41	29.69	30.06	29.75	30.27	29.99	30.04	30.21	30.08	30.00	28.45	30.04	29.89	29.73
SO_3	0.71	0.77	0.79	0.44	0.85	0.65	0.90	0.29	0.84	0.73	0.98	0.74	0.59	0.47	0.54	0.80	0.70
CaO	2.01	1.94	1.51	1.63	1.60	1.94	1.62	1.30	1.52	1.71	1.58	1.94	2.08	1.59	1.30	1.79	2.09
Y_2O_3	2.59	2.57	2.11	2.53	2.11	2.82	2.10	2.57	1.52	2.02	1.81	2.04	2.12	0.65	1.56	1.87	2.25
La_2O_3	12.23	12.32	15.00	13.21	14.51	11.81	14.43	12.20	15.35	13.31	14.64	12.94	14.03	13.81	14.48	13.64	12.01
Ce_2O_3	25.84	26.26	28.50	27.00	27.73	25.24	28.31	26.78	29.13	27.38	28.28	26.82	25.85	27.59	28.68	27.33	25.67
Pr_2O_3	2.84	2.91	3.02	2.96	2.97	3.05	2.99	3.08	2.92	2.94	3.09	2.91	2.58	2.96	3.05	2.91	2.98
Nd_2O_3	11.28	11.59	10.97	11.13	11.10	11.90	11.13	11.84	11.10	11.41	11.24	11.13	10.19	11.02	11.66	11.59	11.23
$\mathrm{Sm}_2\mathrm{O}_3$	1.88	2.00	1.69	1.94	1.86	2.11	1.72	2.18	1.76	1.86	1.83	2.03	1.63	1.79	1.96	2.01	2.03
Gd_2O_3	1.68	1.73	1.42	1.63	1.41	1.73	1.48	1.97	1.44	1.63	1.72	1.79	1.70	1.33	1.51	1.65	1.80
Tb_2O_3	0.12	0.11	0.09	0.16	0.16	0.16	0.14	0.05	0.04	0.13	0.17	0.16	0.13	0.08	0.16	0.14	0.12
Dy_2O_3	0.75	0.72	0.60	0.72	0.61	0.88	0.61	0.85	0.53	0.65	0.71	0.69	0.71	0.35	0.55	0.65	0.74
ThO_2	7.11	6.51	4.11	6.11	4.26	7.18	4.22	5.31	3.68	5.20	3.22	6.55	8.12	5.59	4.18	5.41	7.54
UO_2	1.03	0.86	0.57	0.82	0.55	1.01	0.52	0.71	0.49	0.80	0.63	0.94	0.80	0.69	0.60	0.77	1.00
PbO	0.02	0.02	0.02	0.02	0.02	0.01	0.01	0.01	0.01	0.02	0.01	0.01	0.02	0.01	0.01	0.01	0.02
Total	100.57	100.96	100.98	101.03	99.66	100.95	100.20	99.85	100.56	100.11	100.31	101.10	100.99	97.72	100.56	100.76	100.32
Si	0.0144	0.0142	0.0097	0.0123	0.0111	0.0153	0.0102	0.0167	0.0091	0.0110	0.0075	0.0126	0.0168	0.0533	0.0107	0.0113	0.0153
Р	0.9723	0.9725	0.9750	0.9815	0.9682	0.9708	0.9662	0.9870	0.9715	0.9756	0.9740	0.9709	0.9709	0.9526	0.9784	0.9686	0.9683
S	0.0203	0.0218	0.0225	0.0125	0.0246	0.0185	0.0258	0.0083	0.0241	0.0210	0.0279	0.0210	0.0168	0.0138	0.0156	0.0231	0.0202
Са	0.0822	0.0787	0.0615	0.0664	0.0661	0.0792	0.0667	0.0537	0.0624	0.0702	0.0643	0.0794	0.0852	0.0672	0.0534	0.0736	0.0862
Y	0.0527	0.0520	0.0426	0.0513	0.0432	0.0571	0.0428	0.0526	0.0311	0.0412	0.0366	0.0414	0.0431	0.0137	0.0319	0.0380	0.0461
La	0.1722	0.1724	0.2100	0.1858	0.2062	0.1661	0.2042	0.1734	0.2166	0.1884	0.2056	0.1819	0.1978	0.2014	0.2055	0.1926	0.1703
Ce	0.3610	0.3647	0.3961	0.3769	0.3910	0.3525	0.3976	0.3777	0.4081	0.3845	0.3944	0.3743	0.3618	0.3995	0.4040	0.3830	0.3615
Pr	0.0395	0.0402	0.0418	0.0412	0.0417	0.0424	0.0418	0.0432	0.0407	0.0411	0.0429	0.0404	0.0360	0.0427	0.0428	0.0406	0.0418
PN	0.1538	0.1571	0.1487	0.1515	0.1527	0.1621	0.1525	0.1629	0.1518	0.1563	0.1529	0.1515	0.1391	0.1556	0.1603	0.1585	0.1542
Sm	0.0248	0.0262	0.0221	0.0255	0.0247	0.0277	0.0227	0.0290	0.0231	0.0245	0.0240	0.0266	0.0215	0.0244	0.0259	0.0265	0.0269
Gd	0.0212	0.0218	0.0179	0.0206	0.0180	0.0219	0.0188	0.0251	0.0182	0.0208	0.0218	0.0226	0.0215	0.0175	0.0193	0.0209	0.0230
Тb	0.0015	0.0013	0.0011	0.0020	0.0021	0.0020	0.0018	0.0007	0.0005	0.0016	0.0021	0.0020	0.0017	0.0011	0.0020	0.0018	0.0016
Dy	0.0093	0.0088	0.0074	0.0089	0.0075	0.0108	0.0075	0.0106	0.0066	0.0080	0.0087	0.0085	0.0088	0.0044	0.0068	0.0080	0.0092
Th	0.0617	0.0562	0.0355	0.0530	0.0373	0.0624	0.0368	0.0465	0.0321	0.0453	0.0279	0.0568	0.0706	0.0503	0.0366	0.0472	0.0660
N	0.0087	0.0073	0.0048	0.0070	0.0047	0.0086	0.0045	0.0061	0.0042	0.0069	0.0053	0.0080	0.0068	0.0061	0.0052	0.0065	0.0086
Pb	0.0002	0.0002	0.0002	0.0002	0.0002	0.0001	0.0001	0.0001	0.0001	0.0002	0.0001	0.0001	0.0002	0.0001	0.0001	0.0001	0.0002

Table 4.10 Electron Microprobe Chemical Analysis of Monazite from Sample AS01-16a.

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# zum	1	1	1	1	1	1	1	1	2	2	2	2	2	2	2	2	3
Analysis	1	ю	4	5	9	7	6	10	12	13	14	15	16	17	18	19	21
SiO_2	0.25	0.42	0.25	0.24	0.39	0.37	0.26	0.46	0.24	0.35	0.26	0.26	0.52	0.50	0.25	0.26	0.26
P_2O_5	30.80	30.00	30.28	30.67	30.15	29.95	30.58	30.04	30.58	30.12	30.58	30.72	30.23	29.07	30.77	30.75	30.58
SO_3	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.01	0.02	0.01	0.04	0.01	0.00	0.01
CaO	1.31	1.10	1.25	1.30	1.18	1.19	1.22	1.17	1.30	1.01	1.29	1.31	1.53	1.19	1.34	1.25	1.30
Y_2O_3	2.69	0.47	2.39	2.23	0.55	0.45	2.69	0.45	2.84	0.48	2.24	2.27	0.61	0.50	2.81	2.57	2.85
La_2O_3	13.52	14.14	13.43	13.37	14.43	14.94	13.35	14.12	12.92	14.71	13.32	13.55	13.61	14.08	13.21	13.30	12.76
Ce_2O_3	27.50	29.76	27.53	27.81	29.49	29.62	27.38	29.36	27.18	29.54	27.49	27.84	27.82	28.95	27.05	27.12	26.94
Pr_2O_3	3.00	3.27	3.02	3.03	3.21	3.18	3.09	3.18	2.89	3.23	3.01	2.95	3.08	3.24	2.99	3.03	2.86
Nd_2O_3	11.64	11.90	11.54	11.49	11.98	11.64	11.36	12.12	11.42	12.06	11.69	11.19	11.49	11.92	11.30	11.51	11.48
Sm_2O_3	2.05	1.98	2.06	1.93	2.09	1.92	2.06	2.08	2.04	1.89	2.03	1.98	1.87	1.74	1.95	1.90	1.95
Gd_2O_3	1.95	1.17	1.74	1.71	1.26	1.21	1.83	1.36	1.99	1.20	1.80	1.75	1.30	1.24	1.83	1.72	1.88
Tb_2O_3	0.17	0.02	0.12	0.12	0.10	0.05	0.11	0.11	0.20	0.01	0.16	0.17	0.04	0.00	0.14	0.13	0.13
$\rm Dy_2O_3$	0.78	0.19	0.72	0.71	0.20	0.13	0.88	0.17	0.88	0.18	0.74	0.72	0.25	0.21	0.86	0.79	0.85
ThO_2	5.36	5.75	5.35	5.64	5.99	5.48	5.01	6.27	5.12	5.09	5.64	5.78	7.95	5.77	5.41	5.12	5.21
UO_2	1.29	0.71	0.97	0.96	0.67	0.65	1.21	0.67	1.47	0.62	1.03	1.02	0.99	0.65	1.38	1.33	1.52
PbO	0.02	0.01	0.01	0.01	0.02	0.01	0.01	0.00	0.02	0.01	0.02	0.01	0.02	0.01	0.01	0.01	0.01
Total	102.36	100.91	100.69	101.22	101.72	100.79	101.07	101.56	101.08	100.50	101.30	101.52	101.30	99.11	101.30	100.79	100.59
Si	0.0096	0.0164	0.0097	0.0093	0.0150	0.0144	0.0101	0.0178	0.0093	0.0135	0.0099	0.0098	0.0199	0.0199	0.0095	0.0099	0.0099
Р	0.9905	0.9873	0.9905	0.9951	0.9860	0.9870	0.9933	0.9843	0.9930	0.9922	0.9932	0.9942	0.9879	0.9777	0.9951	0.9980	0.9955
S	0.0003	0.0003	0.0003	0.0003	0.0004	0.0003	0.0004	0.0004	0.0004	0.0006	0.0004	0.0005	0.0002	0.0013	0.0004	0.0001	0.0003
Ca	0.0534	0.0460	0.0520	0.0536	0.0489	0.0494	0.0500	0.0484	0.0533	0.0421	0.0531	0.0537	0.0631	0.0506	0.0548	0.0513	0.0537
Y	0.0543	0.0097	0.0491	0.0456	0.0112	0.0093	0.0550	0.0092	0.0581	0.0099	0.0456	0.0461	0.0126	0.0105	0.0572	0.0525	0.0584
La	0.1894	0.2028	0.1914	0.1890	0.2056	0.2146	0.1889	0.2016	0.1828	0.2111	0.1885	0.1910	0.1937	0.2064	0.1861	0.1881	0.1810
Ce	0.3825	0.4236	0.3895	0.3902	0.4170	0.4222	0.3846	0.4160	0.3817	0.4207	0.3862	0.3897	0.3933	0.4212	0.3784	0.3806	0.3793
Pr	0.0415	0.0463	0.0426	0.0423	0.0452	0.0451	0.0432	0.0449	0.0403	0.0458	0.0421	0.0411	0.0433	0.0469	0.0416	0.0423	0.0400
PN	0.1579	0.1651	0.1593	0.1573	0.1652	0.1618	0.1556	0.1675	0.1564	0.1676	0.1602	0.1527	0.1585	0.1691	0.1541	0.1575	0.1576
Sm	0.0268	0.0265	0.0275	0.0255	0.0279	0.0257	0.0272	0.0277	0.0269	0.0253	0.0268	0.0260	0.0248	0.0239	0.0257	0.0250	0.0259
Gd	0.0246	0.0150	0.0223	0.0217	0.0162	0.0157	0.0233	0.0174	0.0253	0.0155	0.0229	0.0222	0.0166	0.0164	0.0232	0.0218	0.0240
Тb	0.0021	0.0003	0.0016	0.0015	0.0013	0.0006	0.0014	0.0014	0.0025	0.0001	0.0020	0.0021	0.0005	0.0000	0.0017	0.0016	0.0016
Dy	0.0096	0.0024	0.0090	0.0087	0.0024	0.0016	0.0109	0.0021	0.0108	0.0023	0.0091	0.0089	0.0031	0.0026	0.0106	0.0098	0.0105
Th	0.0464	0.0509	0.0471	0.0492	0.0526	0.0486	0.0438	0.0552	0.0447	0.0451	0.0492	0.0503	0.0699	0.0522	0.0470	0.0447	0.0456
N	0.0109	0.0062	0.0083	0.0082	0.0058	0.0056	0.0104	0.0058	0.0125	0.0053	0.0088	0.0087	0.0085	0.0058	0.0117	0.0113	0.0130
ЪЬ	0 000 0	0.0001	0.0001	0.0001	0,000	0.0001	0.0001	0,000,0	0 000 0	0.0001	0,000	0.0001	0,000	0.0001	0.0001	0.0001	0.0001

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# zuu	3	3	3	3	4	4	4	4	4	5	5	5	5	5	5	5	5
Analysis	23	24	25	26	27	28	29	31	33	34	35	36	37	39	40	42	43
SiO_2	0.40	0.48	0.22	0.25	0.26	0.33	0.43	0.26	0.25	0.25	0.42	0.26	0.26	0.38	0.51	0.26	0.26
P_2O_5	30.00	29.98	30.49	30.22	30.87	30.38	30.37	30.60	30.63	30.55	30.29	30.67	30.37	30.39	29.94	30.62	30.65
SO_3	0.00	0.01	0.01	0.01	0.01	0.01	0.02	0.01	0.00	0.02	0.01	0.02	0.00	0.02	0.02	0.01	0.01
CaO	1.37	1.25	1.28	1.31	1.32	1.12	1.24	1.30	1.32	1.25	1.18	1.31	1.25	1.24	1.32	1.34	1.19
Y_2O_3	0.67	0.57	2.02	1.87	2.62	0.46	0.49	2.21	1.97	2.09	0.53	2.29	1.29	2.70	0.48	2.23	1.10
La_2O_3	13.57	14.18	13.54	13.28	12.85	14.16	13.90	13.44	13.24	13.58	14.58	13.68	13.24	12.78	14.56	13.35	13.49
Ce_2O_3	28.56	28.62	28.20	27.91	27.46	29.58	28.78	27.67	27.59	27.69	29.05	27.72	28.60	26.83	28.52	27.55	28.46
Pr_2O_3	2.97	3.08	2.97	2.94	2.86	3.18	3.08	2.95	3.02	3.03	3.14	2.87	3.11	2.95	2.91	3.03	3.20
Nd_2O_3	11.63	11.79	11.47	11.63	11.37	11.92	11.84	11.34	11.44	11.47	11.58	10.90	12.30	11.57	11.35	11.46	11.99
Sm_2O_3	2.05	1.92	1.95	2.08	1.98	1.99	2.02	2.00	2.10	1.97	1.78	1.90	2.19	2.02	1.68	1.98	2.13
Gd_2O_3	1.40	1.16	1.64	1.85	1.76	1.20	1.25	1.83	1.82	1.63	1.22	1.67	1.67	1.89	1.10	1.71	1.62
Tb_2O_3	0.05	0.02	0.12	0.12	0.11	0.04		0.18	0.22	0.12		0.12	0.12	0.11	0.02	0.14	0.06
Dy_2O_3	0.25	0.19	0.63	0.68	0.78	0.20	0.23	0.68	0.69	0.66	0.19	0.71	0.52	0.87	0.19	0.74	0.43
ThO_2	6.82	6.66	5.51	5.73	5.65	5.53	6.28	5.73	5.87	5.48	6.14	5.93	5.48	5.18	7.22	5.70	5.35
UO_2	0.95	0.80	0.97	0.93	1.16	0.66	0.86	1.02	0.96	0.94	0.75	0.90	0.89	1.12	0.75	1.01	0.84
PbO	0.01	0.02	0.01	0.01	0.02	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01
Total	100.69	100.71	101.03	100.83	101.07	100.77	100.77	101.24	101.14	100.75	100.87	100.96	101.30	100.05	100.59	101.14	100.78
Si	0.0156	0.0186	0.0085	0.0096	0.0098	0.0128	0.0166	0.0100	0.0098	0.0097	0.0162	0.0098	0.0100	0.0145	0.0199	0.0101	0.0099
Р	0.9883	0.9870	0.9940	0.9903	0.9984	0.9958	0.9944	0.9941	0.9957	0.9956	0.9924	0.9962	0.9919	0.9935	0.9867	0.9942	0.9999
S	0.0001	0.0004	0.0002	0.0003	0.0003	0.0004	0.0005	0.0002	0.0001	0.0006	0.0002	0.0005	0.0001	0.0006	0.0005	0.0003	0.0004
Са	0.0571	0.0519	0.0526	0.0543	0.0541	0.0465	0.0514	0.0534	0.0543	0.0516	0.0491	0.0537	0.0516	0.0513	0.0549	0.0552	0.0490
Υ	0.0140	0.0118	0.0415	0.0385	0.0533	0.0095	0.0101	0.0451	0.0403	0.0428	0.0109	0.0467	0.0265	0.0554	0.0100	0.0456	0.0225
La	0.1948	0.2034	0.1923	0.1896	0.1811	0.2022	0.1983	0.1902	0.1875	0.1929	0.2081	0.1936	0.1883	0.1821	0.2090	0.1889	0.1917
Ce	0.4069	0.4075	0.3976	0.3956	0.3842	0.4193	0.4074	0.3888	0.3878	0.3903	0.4117	0.3894	0.4039	0.3794	0.4064	0.3869	0.4014
\mathbf{Pr}	0.0421	0.0436	0.0416	0.0414	0.0398	0.0448	0.0435	0.0412	0.0422	0.0425	0.0443	0.0401	0.0438	0.0416	0.0412	0.0423	0.0449
PN	0.1617	0.1638	0.1577	0.1608	0.1551	0.1648	0.1635	0.1555	0.1569	0.1578	0.1600	0.1494	0.1695	0.1596	0.1578	0.1570	0.1650
Sm	0.0274	0.0258	0.0259	0.0278	0.0260	0.0266	0.0269	0.0264	0.0278	0.0261	0.0237	0.0252	0.0291	0.0269	0.0225	0.0262	0.0282
Gd	0.0180	0.0149	0.0210	0.0238	0.0222	0.0154	0.0160	0.0232	0.0232	0.0208	0.0156	0.0212	0.0214	0.0241	0.0142	0.0217	0.0207
Tb	0.0007	0.0002	0.0015	0.0016	0.0014	0.0006	ı	0.0022	0.0027	0.0016	,	0.0015	0.0015	0.0013	0.0002	0.0017	0.0007
Dy	0.0031	0.0024	0.0078	0.0085	0.0096	0.0025	0.0028	0.0084	0.0086	0.0082	0.0023	0.0088	0.0064	0.0108	0.0024	0.0091	0.0053
Th	0.0604	0.0589	0.0483	0.0505	0.0491	0.0487	0.0552	0.0501	0.0513	0.0481	0.0541	0.0518	0.0481	0.0455	0.0639	0.0497	0.0469
N	0.0082	0.0069	0.0083	0.0080	0.0099	0.0057	0.0074	0.0087	0.0082	0.0081	0.0065	0.0077	0.0076	0.0096	0.0065	0.0087	0.0072
Ч	0 0001	0 0002	0.0001	0 0001	0 0002	0 0001	0 0001	0 0001	0.0001	0.0001	0.0001	0.0001	0.0001	0 0002	0.0001	0.0001	0.0001

Table 4.11 Continued.

Table 4.12 Ion Microprobe Age Analyses of Monazite

	Age±2σ		Samp	Age±2σ	
Sample	(Ma)	Comment	le	(Ma)	Comment
AS01-35			AS01-19)c	
1-1	21.6 ± 2.6	low-Y	3-1	19.8 ± 1.0	intermediate-Y
AS01-15b	,		3-2	20.1 ± 0.9	low-Y
1-1	25.5 ± 2.1	mottled Y, Th	3-3	18.1 ± 0.9	high-Y rim
4-1	113.7 ± 5.5	mixed	3-4	20.9 ± 0.9	intermediate-Y
4-2	47.0 ± 2.1	mixed	3-5	19.3 ± 0.9	intermediate-Y
5-1	24.3 ± 1.3	mottled Y, Th	5-1	27.5 ± 1.1	low-Y
5-2	25.3 ± 1.5	mottled Y, Th	5-2	18.5 ± 0.8	high-Y rim
AS01-33b)	·	5-3	22.9 ± 1.0	mixed date
1-1	21.7 ± 1.0	mottled Y, Th	5-4	20.8 ± 1.0	intermediate-Y
1-2	23.1 ± 1.3	mottled Y, Th	6-1*	16.9 ± 0.7	high-Y
4-1	22.1 ± 1.0	mottled Y, Th	6-2*	20.2 ± 0.4	low-Y core
4-2	22.6 ± 1.3	mottled Y, Th	7-1*	20.8 ± 0.7	intermediate-Y rim
5-1	21.3 ± 0.9	mottled Y. Th	7-2*	28.6 ± 0.8	low-Y core
5-2	22.0 ± 1.0	mottled Y. Th	7-3*	19.8 ± 0.9	high-Y rim
AS01-16a	ı	,	AS01-20)a	6
1-2	25.9 ± 1.2	low-Y. Th core	1-1	21.7 ± 1.3	intermediate-Y
1-3	24.9 ± 1.3	intermediate-Y	1-2	26.8 ± 2.0	high-Y
3-1	25.7 ± 1.3	intermediate-Y	3-1	93.6 ± 4.7	inherited
3-2	18.7 ± 1.4	intermediate-Y	5-1	74.7 ± 3.3	inherited
3-3	25.3 ± 1.9	intermediate-Y	5-2	22.7 ± 1.0	mixed date
4-1	332 + 20	low-Y Th core	AS01-22	22.7 = 1.0	initiou unto
4-2	25.2 = 2.0 25.7 + 1.5	intermediate-Y	1-1	1789 + 70	inherited
4-3	27.7 ± 1.5	low-Y Th	3-1	20.9 ± 1.0	high-Y
AS01-16P	27.2 - 1.0)*	1010 1, 111	AS01-32	20.9 = 1.0	ingii 1
1-1	20.3 ± 0.4	mottled V. Th	1_1	293 ± 14	low-V core
1-2	20.3 ± 0.1 25.2 + 0.4	mottled Y. Th	3-1	25.0 ± 1.1	mixed date
2_1	25.2 ± 0.4 24.4 ± 0.4	mottled V. Th	3_2	29.0 ± 1.2 29.1 + 1.4	low-V core
2.1	27.1 ± 0.1 27.0 ± 0.4	mottled V Th	<u> </u>	29.1 ± 1.1 28.2 ± 0.4	mottled V Th
3-1	27.0 ± 0.4 25.4 ± 0.3	mottled Y Th	4-1*	28.2 ± 0.4 28.7 + 0.4	mottled Y Th
3-2	19.8 ± 1.2	mottled Y. Th	4-3*	28.7 ± 0.1 28.2 + 0.4	mottled Y Th
4-1	19.0 ± 1.2 26.0 ± 0.4	mottled V Th	5_1*	20.2 ± 0.1 31.9 ± 1.0	high-V
4-2	20.0 ± 0.4 25.9 ± 0.4	mottled V. Th	5-1 5_2*	31.9 ± 1.0 31.5 ± 0.8	high-V
4-2	25.7 ± 0.4 26.1 ± 0.3	mottled V Th	5-2 6-1*	31.3 ± 0.3 29.7 + 0.4	mottled V Th
4501-174	20.1 ± 0.3	motiled 1, 11	4501-23	27.7 ± 0.4	mottled 1, 11
2_1	170 ± 0.0	High-V Th	1_1	180 ± 10	high_V rim
2-1 3-1	17.9 ± 0.9 17.9 ± 1.8	High-V Th	1-1 1_2	10.9 ± 1.0 23.1 ± 1.4	low-V core
4501-180	17.7 ± 1.0	111 <u>5</u> 11-1, 111	1-2	20.1 ± 1.4 20.5 ± 1.0	low-Y
1_1	256+13	mixed date	1-3	20.5 ± 1.0 20.5 ± 1.2	high-V rim
1-1	25.0 ± 1.5 25.1 ± 1.5	mixed date	1- 4 2_1	20.3 ± 1.2 20.0 ± 1.0	low-V
1-2	25.1 ± 1.5 27.7 ± 1.7	low-V core	2^{-1} 2_2	20.0 ± 1.0 21.4 ± 1.0	low-Y
1-5 4_1	27.7 ± 1.7 29.5 + 1.3	low-V core	2-2 5-1	21.4 ± 1.0 20.3 ± 0.9	low-1
4-1	29.3 ± 1.3 21.3 ± 1.2	high V rim	5 2	20.3 ± 0.9 20.3 ± 1.1	low V
4-2	21.3 ± 1.2 197 + 18	high-V rim	J-2 4501_27	20.3 ± 1.1	10w-1
4501-102	17.7 ± 1.0	ingii-1 iiii	1_1	32.7 ± 1.4	mottled V. Th
1 1	10.0 ± 0.7	high V	1-1	32.7 ± 1.4 30.2 ± 1.5	mottled V. Th
1-1	19.0 ± 0.7 18.7 ± 0.0	high V	3 1	30.2 ± 1.3 36.7 ± 1.7	mottled V. Th
1-2	10.7 ± 0.9 28 4 ± 1 2	low V coro	J-1 4 1	30.7 ± 1.7 33.0 ± 1.5	mottled V. Th
1-3	26.4 ± 1.2 24.0 ± 1.0	now-1 core	4-1 5 1*	33.0 ± 1.3 27.4 ± 1.2	mottled V. Th
1-4 2 1	24.9 ± 1.0 27 4 ± 1 4	low V core	5-1. 5-2*	$3/.4 \pm 1.3$ $3/.2 \pm 1.2$	mottled V Th
2-1	$\angle 1.4 \pm 1.4$	intermediate V	5-2° 6 1*	34.2 ± 1.3	intermediate V
2-2	10.7 ± 0.8 10.6 ± 1.1	intermediate V	0-1" 6) *	$3/.0 \pm 3.3$	intermediate V
2-3	17.0 ± 1.1	mtermediate- Y	6.2*	31.3 ± 3.0	intermediate V
			6.1*	37.2 ± 4.0	low V. Th rim
			0-4	33.0 ± 4.0	10 W 1, 111 1111

Note: All analyses contained 95% radiogenic 208Pb or higher *Data from Kohn, 2005

Thrust	Peak T	t	Cooling T	t	Distance	Rate
	(°C)	(Ma)	(°C)	(Ma)	(km)	(cm/yr)
Sinuwa	775±20	27±1	735±20	23±1	55	1.4±2.8
Bhanuwa	735±20	23±1	650±25	19±2	80	2.0±1.2

 Table 4.13 Estimates of Thrust Temperatures, Times, Minimum

 Displacements and Displacement Rates.

CHAPTER 5. GEOCHRONOLOGY AND THERMOBAROMETRY OF THE CENTRAL HIMALAYA, ARUN RIVER VALLEY, EASTERN NEPAL

5.1 Introduction

Although eclogites have been found in only a few localities in the Himalaya, they are key in understanding the orogen's metamorphic and tectonic evolution (Figure 5.1). In the northwest Himalaya, ultra-high pressure eclogites (UHP) have been documented in the Kaghan Valley of Pakistan (Pognante and Spencer, 1991; Tonarini et al., 1993) and the Tso Morari dome in the Ladakh region of India (de Sigoyer et al., 1997, 2000). In the central Himalaya, eclogites have been documented in the Ama Drime range in the Kharta region of Tibet (Lombardo and Rolfo, 2000; Groppo et al., 2007; Cottle et al., 2009), and the Makalu-Everest region of the Arun River valley in eastern Nepal (Parkinson and Kohn, 2002). Both NW Himalayan eclogite localities occur within the Greater Himalayan Sequence immediately south of the Indus Suture Zone (O'Brien et al., 2001), and are considered to indicate subduction of the leading edge of the Indian continental crust during the initial stages of Indo-Asian collision. These eclogites reached conditions of 725±25°C and 28-30 kbar (Mukherjee and Sachan, 2001; O'Brien et al., 2001). Accessory zircon (Kaneko et al., 2003; Parrish et al., 2006) and garnet (Tonarini et al., 1993), whose chemistry or texture link them to UHP metamorphism, have been directly dated at ~46 Ma in the Kaghan Valley, and a garnet isochron yields ~54 Ma for Tso Morari (de Sigoyer et al., 2000).

In contrast, eclogites in the central Himalaya do not appear to have attained UHP conditions (Lombardo and Rolfo, 2000; Groppo et al., 2007) and direct correlations with eclogites in the NW Himalaya are unclear. The Ama Drime eclogites occur north of the Arun eclogite, up-drainage along the Arun/Phung Chu River within the Ama Drime Massif. Both eclogite occurrences have been strongly overprinted by subsequent metamorphism, leaving sparse evidence for eclogitization. Groppo et al. (2007) proposed four phases of metamorphism for the Ama Drime eclogites on the basis of microstructural observations, pseudosection analysis, and conventional thermobarometry: an initial (M1) eclogitic metamorphism at >15 kbar and >580°C; a peak-T (M2) granulite-facies event at 8-10 kbar and >750°C, and another (M3) at ~4 kbar and 750°C; and a final cooling stage (M4) at ~700°C and 4 kbar. Attempts to date the Ama Drime eclogites have yielded U-Pb SHRIMP ages of 13-14 Ma from low-U rims of zircon, the growth of which was assumed to result from reactions occurring during the "M3" low-P granulitic event (Groppo et al., 2007). Cottle et al. (2009) present monazite and xenotime ages of 13.2 ± 1.4 Ma that they interpret to constrain the timing of peak (M2) granulite metamorphism. Other published ages include zircon ²⁰⁶Pb/²³⁸U ages of 17.6±0.3 Ma, interpreted to represent granulite metamorphism in the northern Ama Drime Range (Li et al., 2003). In sum, all these ages are thought to reflect the granulite-facies overprint, not the age of eclogitization.

In this chapter, we use Lu-Hf dating techniques on garnet, an eclogite-facies mineral, to provide the first direct measurement of the timing of eclogitization in the central Himalaya, as garnet growth would have been initiated during the prograde transition to eclogite facies. The timing of this event is then interpreted in the context of models of Indo-Asian collision and Himalayan tectonics.

5.2 Petrology

Samples from both the LHS and GHS were collected in eastern Nepal along the west side of the Arun River (Figure 5.2). The metabasites occur as decimeter- to meter-scale boudinaged sills and small bodies within the surrounding felsic schist and gneiss, similar to those documented by Cottle et al. (2009) along strike in southern Tibet. Amphibolite samples were collected from LHS (meta)basaltic sills that are structurally continuous with the surrounding schists and gneisses. The granulitized eclogites from the GHS Barun Gneiss were collected from the core of a ~3 meter diameter boudinaged sill. All field documentation, that is photographs, field notes, and GPS points, were lost during the Hurricane Katrina disaster.

Samples of the GHS Barun Gneiss (AR01-43c and AR01-43e) from this study are similar to the granulitized eclogites described previously from the Ama Drime massif (Lombardo and Rolfo, 2000; Groppo et al., 2007; Cottle et al., 2009). These internally undeformed, medium-grained rocks contain garnet, clinopyroxene, plagioclase, amphibole, quartz, trace orthopyroxene (in AR01-43c only), biotite, scapolite, calcite, ilmenite, titanite, zircon, and apatite (Figure 5.3-5.4). The high-calcium bulk composition of these rocks stabilizes titanite as the primary titanium-bearing phase rather than rutile, which is typically assumed to be the higher-P phase. Direct evidence for eclogitization was all but eliminated during subsequent metamorphism. Calcium-rich (~30% grossular) cores of large garnet are the only direct record of the prograde eclogitic assemblage, and were overgrown by thin rims during the later thermal peak of metamorphism (Figure 5.5). True omphacite is not preserved in the matrix or as inclusions in garnet, although sparse inclusions of relatively sodic clinopyroxene (up to 15% jadeite) are preserved in garnet (Figure 5.6). However, these pyroxene inclusions may be prograde, rather than peak-P eclogitic pyroxene (Groppo et al., 2007). In the matrix, presumed omphacite has been entirely replaced by a plagioclase + Na-poor clinopyroxene symplectite that has been documented in other retrogressed eclogites as the result of the transition from eclogite to granulite facies (Lombardo and Rolfo, 2000; Groppo et al., 2007; Liu et al., 2007; Cottle et al., 2009). Similar to Cottle et al. (2009), no pseudomorphs after phengite, i.e., symplectites of high-Fe biotite and plagioclase as described by Groppo et al. (2007), were observed in any samples.

The thermal peak metamorphic assemblage is characterized by relatively coarsegrained diopsidic clinopyroxene (Figure 5.7), plagioclase, garnet, quartz, and trace orthopyroxene. A final thermal event is recorded in the retrograded eclogite by the almost total replacement of orthopyroxene by brown amphibole, which also grew in the clinopyroxene-plagioclase symplectite.

The LHS mafic amphibolites (AR01-2b, -11b, -35b, and -41) generally have a mineralogy consisting of quartz + amphibole + garnet + biotite + ilmenite \pm clinopyroxene \pm plagioclase \pm rutile or titanite (Figs. 5.8-5.11). Regionally, different amphibolites exhibit a range of mineral assemblages (Goscombe and Hand, 2000), probably reflecting basaltic to hydrothermally altered volcanic protoliths. There is no petrographic or textural evidence that these rocks ever experienced eclogite-facies conditions. The samples contain no evidence for omphacite (no pseudomorphs,

symplectites, or inclusions in garnet), and garnets exhibit simple growth zoning of major elements (Figure 5.12).

5.3 Thermobarometry

5.3.1 Major Element Thermobarometry

While no direct barometric measurements can be made for eclogite-facies conditions, an estimate for the minimum pressures experienced by the rocks can be obtained by reintegrating sodic plagioclase into pyroxene (e.g., Will and Schmädicke, 2001). This integration assumes: 1) at eclogite facies, all sodium was hosted in clinopyroxene; 2) the albite and anorthite components of plagioclase texturally associated with matrix pyroxene (either in the symplectites or surrounding large diopsidic clinopyroxenes; Figure 5.7) were hosted in omphacite as the jadeite and Ca-Tschermaks components, respectively. This does not account for sodium in matrix hornblende or calcium originally hosted by garnet, and provides only a minimum estimate of the jadeite content. Assuming an average composition for plagioclase and diopside as measured by electron microprobe, the number of moles of each species is calculated via their respective modal abundances and molar volume. In sample AR01-43e, plagioclase has an average composition of $An_{19}Ab_{79}$ and modal abundances are 23 and 16% for plagioclase and clinopyroxene (estimated via \sim 300 point counts per sample). This reintegrates to an omphacitic pyroxene with a jadeite content of 39.1 mol%, which corresponds to a minimum pressure of 15 kbar at 670°C (Figure 5.6; Holland, 1980), with temperature estimated via Zr-in-titanite thermometry (see below). The same calculation for AR01-43c (An₃₉Ab₅₈; 21% plagioclase, 12.5% clinopyroxene) returns a jadeite component of 30.9

mol%, corresponding to 14.1 kbar at 670°C. Calculations based on small regions of symplectite yield greater scatter, but comparable results.

Pressure-temperature conditions for the post-eclogite metamorphic events were calculated using the garnet-clinopyroxene (Pattison and Newton, 1989) and garnet-hornblende (Graham and Powell, 1984) thermometers and the garnet-plagioclase-clinopyroxene-quartz (Eckert et al., 1991) and garnet-plagioclase-hornblende-quartz (Kohn and Spear, 1990) barometers (Figure 5.13a). Compositions for each calculation were selected based on spatial and textural relationships among the individual components; i.e., symplectitic and coronal plagioclase was chosen for garnet-plagioclase-clinopyroxene-quartz barometry and matrix plagioclase associated with hornblende for garnet-plagioclase-hornblende-quartz barometry (Tables 5.1-5.4). Intersection of the garnet-clinopyroxene and the garnet-plagioclase-clinopyroxene-quartz equilibria indicate pressure-temperature conditions of about 780°C and 12 kbar for the granulite-facies overprint. A later, third metamorphic event, represented by the growth of amphibole in the matrix, occurred around 675°C and 6 kbar.

5.3.2 Trace Element Thermobarometry

Zirconium-in-titanite thermometry (Hayden et al., 2008) was used to calculate temperatures for retrograded eclogites AR01-43c and -43e. Back-scattered electron images of titanite show no intracrystalline chemical zoning, although different grains have different compositions. Zircon is evident in all samples, either as observable matrix grains, or in mineral separates, or by relatively high Hf contents of bombed whole rock powders. Rutile is not present in samples where Zr-in-titanite thermometry was applied, and the activity of TiO₂ was estimated to be 0.75 based on equilibrium among garnet, plagioclase, titanite, and quartz. Variations in $a(TiO_2)$ of ±0.1 change apparent temperatures by 7°C. Temperatures were calculated for matrix grains at pressures previously obtained via conventional thermobarometry for both the granulite and amphibolite metamorphic events, about 12 and 6 kbar, respectively (Table 5.5). At 12 kbar, core and rim temperatures of matrix titanite closely bracketed the ~780 °C calculated via garnet-clinopyroxene thermometry (Figure 5.13b). Yet, at 6 kbar, temperatures were ~710°C, 35°C higher than garnet-hornblende temperatures. It is possible titanite growth ceased prior to final growth and equilibration of hornblende.

Zirconium concentrations were measured from titanite inclusions in large matrix pyroxenes from sample AR01-43e (Table 5.5; Figure 5.7b). The resultant Zr-in-titanite temperatures (Hayden et al., 2008) were 35-130°C lower than matrix temperatures calculated for the same pressure (e.g., 650-745°C vs. ~780°C at 12 kbar), although apparent temperatures generally increase towards the pyroxene rim. A minimum pressure estimate for the lowest temperature titanite inclusion from a pyroxene core (~670°C), which presumably pre-dates the granulite event, is 15 kbar (Figure 5.13b; Holland, 1980), well within the eclogite-facies. This is consistent with initial formation of titanite in the eclogite-facies, with continued growth or recrystallization as the system began to exhume and heat from the eclogite-facies toward the peak-T granulite event.

The Zr-in-rutile thermometer (Watson et al., 2006) was also used on amphibolite samples AR01-35b and -41 and the temperatures calculated were approximately 660°C, similar to temperatures calculated for the amphibolite facies overprint in the eclogites. The Zr-in-rutile thermometer is only weakly pressure dependent (Tomkins et al., 2007),

and as these two samples do not contain plagioclase, an independent evaluation of pressure was not possible.

5.4 Lu-Hf Dating of Garnet

Garnet-whole rock isochrons yield ages ranging from 13.9 to 15.1 Ma for amphibolite samples, and 20.7 ± 0.4 Ma for a retrograded eclogite (Table 5.6, Figure 5.14). Analysis of bomb-digested whole-rock revealed significantly lower Lu-Hf ratios and higher Hf concentrations compared to tabletop digestions due to dissolution in the former of zircon. Such data were omitted from calculated isochrons if the ¹⁷⁶Hf/¹⁷⁷Hf of the bomb-digested whole-rock differed from the ¹⁷⁶Hf/¹⁷⁷Hf of the tabletop-digested whole-rock by more than 1.5×10^{-5} , presuming the zircon in these samples contained an inherited component not in equilibrium with garnet growth. If used in the regression, these data cause the calculated age to be spuriously old (e.g., Scherer et al., 2000). Models of Lu uptake in garnet (Lapen et al., 2003; Kohn, 2009) imply that garnet Lu-Hf ages should reflect early growth of garnet. Garnet X-ray maps from AR01-43c indicate that garnet cores represent $\sim 75\%$ of total garnet volume, but rims are enriched in Lu by a factor of ~ 1.4 compared to cores (Figure 5.5). Thus, 60-70% of the bulk garnet Lu-Hf age reflects eclogite-facies garnet growth, while 30-40% reflects growth in granulite-facies. Given the chemical simplicity of the amphibolite-facies LHS garnets in the underlying MCT footwall, their Lu-Hf ages are interpreted simply to reflect later amphibolite-facies metamorphism.

5.5 Discussion: P-T-t evolution and tectonic implications

Three main metamorphic stages are recognized in the retrograded eclogite samples from eastern Nepal. Pressure-temperature conditions of primary eclogite facies metamorphism are loosely constrained at 670°C and >15 kbar (≥50 km depth) by albitejadeite-quartz barometry and Zr-in-titanite thermometry. These are similar to conditions inferred by Lombardo and Rolfo (2000) and Groppo et al. (2007) for rocks along strike in Tibet. The eclogite was then strongly overprinted by a peak-temperature, granulite-facies metamorphism at \sim 780°C and 12 kbar, recorded by garnet-plagioclase-clinopyroxene equilibria and Zr-in-titanite thermometry. This was followed by a later, lower pressure and temperature event at ~675°C and 6 kbar, represented by the widespread growth of amphibole in the matrix and Zr-in-rutile thermometry. These conditions produce a clockwise P-T path characterized by heating during initial exhumation (Figure 5.13b). The general shape of the P-T path agrees well with Groppo et al. (2007), but specific P-T results differ in that no evidence was found for the low pressures (e.g., 4 kbar) they reported. Our estimated final pressures correspond well with conventional thermobarometric results of Goscombe and Hand (2000) in the region for metapelites and amphibolites.

The 20.7 ± 0.4 Ma age for the GHS eclogite is the first attempt to measure directly the age of eclogitization in the central Himalaya, although several attempts have been made to date the granulite metamorphism in the region. Cottle et al. (2009) ascribed ≤ 13 Ma monazite dates to peak granulite metamorphism and prograde anatexis. However, partial melting should consume, not produce, monazite (Pyle and Spear, 2003; Kohn et al., 2004, 2005; Kelsey et al., 2008), and these ages more likely represent postpeak cooling and crystallization, presumably related to exhumation, thrust transport, or both. Groppo et al. (2007) also interpreted zircon rim ages of 13-14 Ma to represent late-stage granulite-facies metamorphism, but provide little direct evidence for this interpretation and concede that the zircon ages could instead be linked to the amphibolitefacies event. Overall, we interpret the 13-14 Ma monazite and zircon ages to reflect postgranulite facies overprinting associated with cooling during MCT transport. This interpretation agrees with the four 14-15 Ma prograde ages for garnet amphibolites from the LHS in the MCT footwall, and is consistent with relatively late initiation of the MCT along strike in Bhutan and central Nepal (Daniel et al., 2003; Kohn et al., 2004). Considering the weighted proportions of Lu in garnet cores (eclogite-facies: 60-70%) vs. rims (granulite-facies: 30-40%), and assuming that these components are mixed in these proportions in the garnet fractions, it implies an age of eclogitization of 23-26 Ma, and a latest possible granulite-facies age of 13-15 Ma.

The P-T-t evolution of the HP Arun and Ama Drime eclogites contrasts with the UHP eclogites from the northwest Himalaya in three fundamental ways: the Arun rocks were metamorphosed at much lower pressures; they show significant heating during initial exhumation; and they were metamorphosed 20-30 Myr later (Groppo et al., 2007; Cottle et al., 2009; this study). These differences potentially signal different tectonic processes. With respect to P-T paths, the nearly complete granulite-facies overprinting of the Arun eclogite may require either an extra heat source or a combination of slow erosion plus strongly diminished thrust transport in the mid-crust after loading. Rapid erosion alone should exhume rocks isothermally, whereas rapid thrusting on lower level thrusts should cool rocks isobarically. In principle, heating might also occur upon

expulsion of relatively cool lower crustal rocks into an inverted geotherm. The occurrence of rapid exhumation (~30 km in ~7 Myr) with heating points to additional heat sources or an inverted geotherm, and we evaluate possible competing models within that context.

It is important to note that all models of Himalayan orogenesis predict eclogites in the deep-crust. Thus, they key question is not the genesis of ~ 25 Ma eclogites, but rather why they were exhumed and exposed at Arun and apparently not elsewhere. One possible model is that their appearance simply reflects deepening or longer transport of the MCT in the Arun area (Figure 5.15, Model 1a and 1b). This explanation is supported by relatively low pressures of eclogitization – as little as 50 km depth, not \geq 90 km as evident in the northwest Himalaya. Peak metamorphic pressures near the base of the MCT in central Nepal are already documented at 12 kbar (Kohn, 2008), and slight changes to thrust geometry and dynamics could allow sampling of eclogite-facies, rather than amphibolite-facies GHS rocks (Henry et al., 1997). The disparity in eclogite ages then reflects different exhumation processes – slab breakoff at ~50 Ma and "normal" Himalayan thrust transport at ~20 Ma. This model, however, does not adequately explain the granulite-facies overprint at Arun. Within the context of fold-thrust belts, exhumation between 25 and 14 Ma implies the eclogites were transferred to the hanging wall, but thrust transport in the hanging wall should cause cooling, not heating.

A second possible model is that (aborted) subduction, slab breakoff, and buoyant ascent of Indian crust occurred diachronously: ~50 Ma in the western Himalaya, ~25 Ma in the central Himalaya, and presumably even younger in the eastern Himalaya. By this model (Figure 5.15, Model 2a), mechanical coupling between the oceanic and Indian

continental lithospheres subducted India beneath Tibet to eclogite-facies conditions. Decoupling and removal of the down-going oceanic plate would cause asthenospheric upwelling and buoyant extrusion of a sliver of Indian crust out of the mantle lithosphere, causing rapid exhumation of the GHS and its eclogites (Kohn and Parkinson, 2002). Increased heat flow due to mantle upwelling could have increased temperatures to the granulite facies during exhumation. Differences along strike in exhumation rate or thermal structure might have allowed eclogites in the northwest Himalaya to exhume isothermally. Subsequent thrusting and transport would have cooled the Arun eclogites as they were incorporated in the Himalayan fold-thrust belt. While this model does explain our data, there is little supporting evidence elsewhere in the Himalaya for such profound along-strike diachroneity in fundamental tectonic processes.

A third, "hybrid" model recognizes the possibility of repeated loss of mantle lithosphere, possibly by different mechanisms (Chemenda et al., 2000). For example, in the northwest Himalaya, UHP eclogites may have been exhumed early by breakoff of the oceanic slab (O'Brien, 2001; Mahéo et al., 2002; Treloar et al., 2003; Parrish et al., 2006), whereas the Arun HP eclogites may have been exhumed later by delamination of Indian lithosphere (Figure 5.15, Model 2b). Like slab breakoff, delamination is expected to produce mantle upwelling, leading to a granulite-facies overprint. Again, we know of no other supporting evidence elsewhere along strike for this process, but the possibly unique occurrence of such young eclogites at Arun may require a geographically restricted process.

Regarding more specific thermal-mechanical models of thrust-belt development in the Himalaya, both channel flow (e.g., Beaumont et al., 2001) and critical taper (Robinson et al., 2006) have been proposed to explain the metamorphism of the GHS (Jamieson et al., 2004; Kohn, 2008), and could in principle accommodate the presence of eclogites. However, neither model explains some key features of the Arun rocks. Channel flow models predict a focused erosional front for the last c. 25 Myr (Beaumont et al., 2001; Jamieson et al., 2004). Yet the Arun rocks were exhumed by ~30 km between 14 and 25 Ma, apparently far from any erosional front. This admittedly local observation casts doubt on how channel flow models account for erosion distributions and hence their predictive power for the Arun rocks. Moreover, the upper Arun River Valley is already differentiated from other areas along strike in the Himalaya by local E-W extensional structures that formed \leq 13 Ma (Jessup et al., 2008; Cottle et al., 2009). As described previously, critical taper has trouble explaining P-T paths that combine exhumation with heating. Major changes must be invoked for the mechanical behavior of the wedge (to accommodate such profound thinning) and thrust rates (to accommodate heating) between 25 and 14 Ma.

One way of reconciling both channel flow and critical taper models with the Arun eclogite P-T-t path postulates that as parts of the lower crust heat and weaken, they may be expelled to the mid-crust, e.g. as modeled by Beaumont et al. (2006, model LHO-2; Figure 5.15, Model 3). An inverted geotherm is required to produce heating during expulsion, but this is predicted both in steady-state and time-dependent models (e.g., Royden, 1993; Henry et al., 1997; Beaumont et al., 2001). Less clear are the effects of expulsion on thermal state and later structures, particularly given the young age of exhumation (14-25 Ma). In this regard, the P-T-t history of the Arun and Ama Drime

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eclogites may signal a change in the physical state of the Himalayan metamorphic wedge at 16-25 Ma, ultimately giving rise to the MCT by 15-16 Ma.



STDS = South Tibetan Detachment System; MCT = Main Central Thrust; MBT = Main Boundary and locations of Himalayan eclogites (modified from Yin, 2006). Black box shows area of Fig. 2. Figure 5.1. Generalized geologic map of the Himalaya showing major rock types and structures, Thrust.



Figure 5.2. Geologic map of the Arun River Valley (after Goscombe and Hand, 2000; Goscombe et al., 2006). Samples were collected along the west side of the Arun Gorge on trails (represented by the thick black line). Numbers for sample locations follow the pattern AR01-X. Inset: Generalized regional map showing relationship between the orthogneiss in the Ama Drime Massif (ADM) and the rocks exposed in the Arun River Valley, Nepal (after Cottle et al., 2009). Thick barbed lines are the High Himalayan Thrust (HHT; a thrust within the GHS), Main Central Thrust (MCT), and Munsiari Thrust (MT). Single dashed lines are the Ama Drime Detachment (ADD) and the Nyonno Ri Detachment (NRD); double-dashed line is the South Tibetan Detachment system (STDS). Labeled circles indicate towns.



Figure 5.3. Photomicrograph of retrograded eclogite AR01-43c. Mineral assemblage consists of: hornblende + clinopyroxene + garnet + plagioclase + quartz + biotite + trace orthopyroxene + ilmenite + titanite + zircon + apatite + scapolite + calcite. The rock is speckled dark green and white, with 2-3 mm garnets scattered heavily throughout the matrix.



Figure 5.4 Photomicrograph of retrograded eclogite AR01-43e. Mineral assemblage consists of: clinopyroxene + hornblende + garnet + plagioclase + quartz + biotite + ilmenite + titanite + zircon + apatite + scapolite + calcite. The rock is speckled dark green and white, and contains garnets that are 2-3 mm in diameter. Large diopsidic clinopyroxenes in matrix are up to a centimeter in length.



Figure 5.5. Ca X-ray map of garnet from sample AR01-43c with sketch showing core-rim relationship. Garnets exhibit broad cores with a relatively high Ca content (XGrs ≥ 0.28) and lower-Ca rims (XGrs ≤ 0.23). Lutetium concentrations are in ppm and show a slight enrichment of Lu in the overgrowth rims relative to the cores. Also note that matrix plagioclase is zoned from more sodic cores to more calcic rims. Scale bar is 500 mm. Grt = garnet, Hbl = hornblende, Pl = plagioclase, Scp = scapolite.



Figure 5.6. (Wollastonite + enstatite + ferrosilite) – jadeite – aegirine pyroxene compositional diagram (Morimoto et al., 1988). Matrix clinopyroxene is Na-poor (diopsidic). Clinopyroxene inclusions in garnet contain up to 15 mol% jadeite and plot close to the omphacite field. Re-integrated pyroxenes plot well within the omphacite field. n = 57.



Figure 5.7. X-ray maps of a large crystal of diopsidic matrix pyroxene from sample AR01-43e. (A) Sodium x-ray map. Pyroxene surrounded by corona of sodic (average = Ab66) plagioclase. (B) Titanium x-ray map illustrating abundant titanite in the matrix and as inclusions in pyroxene. Scale bar is 2 mm. Grt = garnet, Hbl = hornblende, Pl = plagioclase, Cpx = clinopyroxene, Qtz = quartz, Ttn = titanite, Ilm = ilmenite.



Figure 5.8. Photomicrograph of amphibolite AR01-2b. Mineral assemblage consists of: hornblende + clinopyroxene + garnet + quartz + plagioclase + ilmenite + biotite + titanite + apatite. The rock is dark green and contains abundant garnets that are 2-4 mm in diameter.



Figure 5.9. Photomicrograph of amphibolite AR01-11b. Mineral assemblage consists of: hornblende + plagioclase + garnet + ilmenite + quartz + rutile + apatite. The rock is dark green, exhibits a weak foiliation and contains irregularly shaped garnets that are \sim 5 mm in diameter surrounded by coronas of plagioclase and quartz.



Figure 5.10. Photomicrograph of amphibolite AR01-35b. Mineral assemblage consists of: hornblende + biotite + garnet + quartz + ilmenite + rutile + apatite. The rock is dark green and contains small mm-scale segregations of quartz scattered throughout the body. Garnet sizes range from 1 mm up to 1 cm in diameter, and are scattered heavily throughout the matrix. Rutile is present as inclusions in garnet.



Figure 5.11. Photomicrograph of amphibolite AR01-41. Mineral assemblage consists of: gedrite + garnet + biotite + brown hornblende + ilmenite + rutile + quartz + apatite. The rock is greenish-brown and contains centimeter-sized garnets within a matrix primarily composed of orthoamphibole (gedrite). Thin (mm- to cm-scale) leucosomes, presumably originating from the host felsic rock, intrude into the mafic sill. Ilmenite is present only as inclusions within garnet cores, while the rutile is present in garnet rims and abundant in the matrix



Figure 5.12. Major element x-ray maps of garnet from amphibolite sample AR01-41. Compositional trends are consistent with growth zoning. Scale bars are 500 mm.



Figure 5.13. Pressure-temperature plots for retrograded eclogite AR01-43e. (A) P-T plot showing equilibrium conditions for the granulite and amphibolite metamorphic events calculated for sample AR01-43e. Intersection of the garnet-clinopyroxene (Pattison and Newton, 1989) and garnet-plagioclase-clinopyroxene-quartz (Eckert et al., 1991) equilibria indicates ~780°C and ~12 kbar for the granulite event. A later amphibolitefacies overprint is recorded at ~675°C, ~6 kbar by garnet-hornblende (Graham and Powell, 1984) and garnet-plagioclase-hornblende-quartz (Kohn and Spear, 1990) equilibria. (B) Pressure-temperature plot showing Zr-in-titanite equilibrium lines plotted with the barometers Jd-Ab-Qtz (Holland, 1980, calculated from the re-integration of plagioclase into clinopyroxene), Grt-Pl-Cpx-Qtz and Grt-Pl-Hbl-Qtz from (A). Small black arrows indicate minimum pressure estimate for Jd-Ab-Qtz equilibria; yellow boxes represent average P-T conditions from (A); small white boxes denote points along the P-T path determined by combining Zr thermometry with conventional barometry; large dark gray arrow indicates general P-T trajectory; and light gray arrow indicates P-T trajectory of Groppo et al. (2007) for the Ama Drime.



Figure 5.14. Lu-Hf isochron plots for eclogite and amphibolites. An age of ~ 21 Ma is calculated for the GHS Arun eclogite, and LHS amphibolite ages range from 14 to 15 Ma. All errors are at 95% confidence. G = garnet analysis; WR = whole-rock analysis.



Figure 5.15. Schematic illustrations of potential tectonic models for the P-T-t evolution of Arun eclogites showing overall thrust geometries, implied thermal structure, and migmatite zone ($T \ge 700$ °C, gray shading). MCT = Main Central Thrust. Gray ellipse = GHS rocks from Arun; Gray squares = GHS rocks elsewhere. Model 1: Steady-state, critical taper model for the Himalaya (based on Henry et al., 1997 and Bollinger et al., 2006). At Arun, the MCT may have transported material further (model 1a) or cut deeper (model 1b) than elsewhere in the orogen. This model does not explain the granulite-facies overprint at Arun.

Model 2: Model of slab breakoff (model 2a; after Kohn and Parkinson, 2002) or delamination (model 2b; after Chemenda et al., 2000). Detachment and removal of the down-going plate or mantle lithosphere led to asthenospheric upwelling and buoyant extrusion of a sliver of Indian crust, causing rapid exhumation of the GHS and the eclogites. Both models explain granulite-facies overprint. Dark gray shaded area = Indian continental crust.

Model 3: Channel Flow model (based on model LHO-2 from Beaumont et al., 2006) in which weak areas in the lower plate detach, transfer to the upper-plate, and exhume. This model may explain granulite-facies overprint. Stippled area = strong lower plate; W = weak portion of lower plate.

Analysis									
#	<u>43e.23</u>	<u>43e.101*</u>	<u>43e.121</u>	<u>43e.116*</u>	<u>43e.105*</u>	<u>43e.140*</u>	<u>43e.4</u>	<u>43e.109</u>	<u>43e.225</u>
Textural									
setting	rim	rim	rim	mantle	mantle	mantle	core	core	core
SiO ₂	37.87	37.59	37.36	37.61	37.34	37.43	38.26	38.07	38.17
TiO ₂	0.06	0.04	0.05	0.06	0.05	0.08	0.05	0.06	0.06
Al_2O_3	21.82	22.02	22.10	21.83	21.67	22.07	21.62	22.34	21.73
FeO	28.38	28.51	27.90	26.73	26.52	27.46	27.21	26.75	25.63
MnO	0.74	1.06	1.36	0.82	0.82	1.15	1.97	0.42	0.61
MgO	3.92	4.02	3.82	4.30	4.24	3.32	2.85	3.45	3.45
CaO	8.35	8.28	8.57	9.63	9.85	9.89	10.18	10.95	10.51
Total	101.13	101.51	101.16	100.99	100.48	101.40	102.14	102.03	100.17
Si	2.960	2.933	2.925	2.936	2.932	2.927	2.974	2.941	2.989
Ti	0.003	0.003	0.003	0.004	0.003	0.005	0.003	0.003	0.003
Al	2.009	2.025	2.039	2.009	2.005	2.034	1.980	2.034	2.006
Fe	1.855	1.861	1.827	1.745	1.741	1.795	1.769	1.728	1.679
Mn	0.049	0.070	0.090	0.054	0.054	0.076	0.130	0.027	0.041
Mg	0.456	0.468	0.446	0.501	0.496	0.386	0.330	0.398	0.403
Ca	0.699	0.692	0.719	0.806	0.828	0.829	0.847	0.906	0.882
Alm	60.6	60.2	59.3	56.2	55.8	58.2	57.5	56.5	55.9
Grs	22.9	22.4	23.3	25.9	26.6	26.8	27.6	29.6	29.3
Рур	14.9	15.1	14.5	16.1	15.9	12.5	10.7	13.0	13.4
Sps	1.6	2.3	2.9	1.7	1.7	2.5	4.2	0.9	1.4

 Table 5.1 Representative Electron Microprobe Analyses of Garnet

All Fe reported as 2+ Cations calculated on the basis of 12 oxygens * Composition used for thermobarometric analysis

1 aute 2.2	Nepresent	Lauve Ele	CULUII IVI	ici opi or	e Allal	VSES UL 1	riagiocia	sc							
Analysis #	43e.143*	43e.203*	43e.227	43e.38	43e.39	43e.15	43e.207*	43e.208*	43e.209	43e.210	<u>43e.8</u>	43c.13	43c.14	43c.62	43c.63
Textural setting	sympl.	sympl.	sympl.	mtx, core	mtx, core	mtx, rim	mtx, rim	mtx, rim	cpx corona	cpx corona	vein in grt	sympl.	sympl.	sympl.	cpx corona
SiO_2	61.61	61.64	62.71	56.06	55.39	49.22	52.24	52.26	63.77	64.03	46.84	56.89	56.50	57.38	58.26
Al_2O_3	24.11	23.88	23.47	28.08	28.57	33.25	30.27	30.04	23.02	22.75	34.45	27.37	27.37	26.75	25.45
FeO	0.07	0.01	0.32	0.04	0.04	0.24	0.13	0.09	0.01	0.02	0.29	0.22	0.18	0.19	0.08
CaO	4.99	5.07	4.50	9.33	9.93	14.91	12.54	12.09	3.68	3.63	16.65	8.82	8.73	8.06	7.03
Na_2O	8.81	8.89	9.34	6.20	5.98	2.88	4.36	4.62	9.82	9.56	1.93	6.81	6.89	7.19	8.06
K_2O	0.26	0.40	0.39	0.56	0.61	0.12	0.24	0.21	0.32	0.19	0.08	0.30	0.29	0.58	0.46
Total	99.84	06.66	100.73	100.27	100.52	100.62	99.78	99.30	100.60	100.17	100.24	100.40	96.66	100.15	99.34
Si	2.739	2.742	2.767	2.517	2.487	2.234	2.376	2.385	2.805	2.821	2.146	2.549	2.543	2.576	2.630
Al	1.263	1.252	1.221	1.486	1.512	1.779	1.622	1.616	1.193	1.181	1.861	1.445	1.452	1.415	1.354
Fe	0.003	0.001	0.012	0.001	0.002	0.009	0.005	0.003	0.000	0.001	0.011	0.008	0.007	0.007	0.003
Са	0.238	0.242	0.213	0.449	0.478	0.725	0.611	0.592	0.173	0.171	0.818	0.423	0.421	0.387	0.340
Na	0.759	0.767	0.799	0.540	0.521	0.254	0.384	0.409	0.837	0.816	0.171	0.591	0.602	0.626	0.706
K	0.015	0.023	0.022	0.032	0.035	0.007	0.014	0.012	0.018	0.011	0.005	0.017	0.017	0.033	0.026
An	23.5	23.4	20.6	44.0	46.2	73.6	60.5	58.4	16.8	17.2	82.3	41.0	40.5	37.0	31.7
Ab	75.1	74.4	77.3	52.9	50.4	25.7	38.1	40.4	81.4	81.8	17.2	57.3	57.9	59.8	65.8
Or	1.4	2.2	2.1	3.1	3.4	0.7	1.4	1.2	1.7	1.1	0.5	1.6	1.6	3.2	2.5
All Fe rep	orted as 2	+													

Table 5.2 Representative Electron Microprobe Analyses of Plagioclase

Cations calculated on the basis of 8 oxygens * Composition used for thermobarometric analysis 106
Table 5.3	Represe	entative	electror	n micropr	obe analy	ses of clir	nopyroxe	ne						
Analysis #	43e.18	43e.21	43e.27	43e.135*	43e.138*	43e.139*	43e.154	43e.157	43c.108	43c.119	43c.120	43c.202	43c.203	43c.204
Textural setting	matrix	matrix	matrix	sympl.	sympl.	sympl.	sympl.	sympl.	matrix	matrix	matrix	incl. in grt	incl. in grt	incl. in grt
SiO_2	51.95	52.45	52.65	51.52	52.19	52.73	51.57	51.81	52.06	51.48	52.20	52.27	51.69	52.99
TiO_2	0.20	0.10	0.07	0.10	0.13	0.14	0.22	0.19	0.16	0.16	0.10	0.07	0.13	0.12
Al_2O_3	1.66	1.12	0.69	3.69	2.91	4.17	3.90	3.65	1.58	2.72	1.08	7.01	6.92	4.95
FeO	10.92	10.45	8.88	13.27	12.77	12.03	12.36	11.86	9.10	8.97	8.31	8.84	8.32	8.50
MnO	0.11	0.13	0.10	0.03	0.10	0.09	0.05	0.05	0.06	0.08	0.08	0.03	0.04	0.01
MgO	12.13	12.03	13.15	11.64	11.10	10.33	11.58	11.18	12.96	12.61	13.20	11.52	12.51	13.40
CaO	21.96	22.86	23.27	19.07	20.08	19.70	19.89	20.64	23.78	22.99	24.40	17.97	18.23	17.66
Na_2O	0.31	0.23	0.22	0.60	0.61	0.71	0.67	0.69	0.32	0.49	0.23	2.22	1.96	2.01
Total	99.25	99.36	99.02	99.92	99.89	96.90	100.23	100.06	100.03	99.50	99.58	99.93	99.79	99.65
Si	1.969	1.987	1.987	1.944	1.974	1.994	1.935	1.949	1.943	1.929	1.955	1.923	1.898	1.947
Al^{IV}	0.031	0.013	0.013	0.056	0.026	0.006	0.065	0.051	0.057	0.071	0.045	0.077	0.102	0.053
AI^{VI}	0.043	0.037	0.018	0.108	0.104	0.180	0.108	0.111	0.013	0.049	0.002	0.226	0.198	0.162
Ti	0.006	0.003	0.002	0.003	0.004	0.004	0.006	0.005	0.005	0.005	0.003	0.002	0.004	0.003
Fe^{3+}	0.000	0.000	0.007	0.000	0.000	0.000	0.000	0.000	0.057	0.048	0.054	0.005	0.035	0.027
Fe^{2+}	0.346	0.331	0.274	0.419	0.404	0.380	0.388	0.373	0.227	0.233	0.206	0.267	0.220	0.234
Mg	0.685	0.679	0.739	0.654	0.626	0.582	0.647	0.627	0.721	0.704	0.736	0.631	0.684	0.734
Mn	0.004	0.004	0.003	0.001	0.003	0.003	0.002	0.002	0.002	0.002	0.003	0.001	0.001	0.000
Са	0.892	0.928	0.941	0.771	0.814	0.798	0.800	0.832	0.951	0.923	0.979	0.708	0.717	0.695
Na	0.023	0.017	0.016	0.044	0.045	0.052	0.048	0.050	0.023	0.035	0.017	0.158	0.139	0.143
Aeg	0.0	0.0	0.7	0.0	0.0	0.0	0.0	0.0	2.3	3.5	1.7	0.5	3.5	2.7
Jd	2.3	1.7	0.9	4.4	4.5	5.2	4.8	5.0	0.0	0.0	0.0	15.4	10.4	11.6
Wo	44.6	46.4	47.0	35.6	39.2	39.4	36.5	38.8	47.6	46.2	48.9	35.4	35.9	34.8
En	34.2	34.0	37.0	32.7	31.3	29.1	32.4	31.3	36.1	35.2	36.8	31.6	34.2	36.7
$\mathbf{F}_{\mathbf{S}}$	17.3	16.6	13.7	20.9	20.2	19.0	19.4	18.7	11.4	11.7	10.3	12.4	11.0	11.7
Fe3+ recal	culated a	lccording	3 to stoic	chiometry										
Cations ca	lculated	on the bi	asis of 6	oxygens										
* Compos	ition used	d for the	rmobaroi	metric ana	lysis									

Analysis #	<u>43e.148</u>	<u>43e.149</u>	<u>43e.211*</u>	<u>43e.303*</u>	<u>43c.21</u>	<u>43c.24</u>	<u>43c.26</u>	<u>43c.61</u>
Textural setting	sympl.	sympl.	matrix	matrix	matrix	sympl.	sympl.	matrix
SiO ₂	40.55	40.67	42.24	42.04	42.44	41.16	41.32	41.39
TiO ₂	2.03	1.91	1.95	2.04	1.66	1.75	1.97	1.79
Al_2O_3	12.68	12.85	12.68	12.32	12.88	13.28	13.33	13.63
FeO	17.75	17.44	16.03	16.95	14.71	15.56	15.21	14.68
MnO	0.13	0.15	0.12	0.10	0.09	0.09	0.06	0.02
MgO	9.10	9.17	9.45	9.84	10.87	10.19	10.13	10.47
CaO	11.04	11.04	11.25	11.21	11.22	11.03	11.06	11.36
Na ₂ O	1.35	1.34	1.35	1.32	1.21	1.37	1.45	1.60
K_2O	2.84	2.81	3.10	2.92	3.23	3.44	3.42	3.46
Total	97.47	97.37	98.18	98.75	98.31	97.87	97.95	98.39
Si	6.213	6.225	6.356	6.317	6.339	6.224	6.232	6.202
Ti	0.234	0.220	0.221	0.231	0.186	0.199	0.223	0.202
Al	2.290	2.318	2.250	2.182	2.268	2.366	2.369	2.407
Fe	2.275	2.233	2.017	2.130	1.837	1.968	1.918	1.839
Mn	0.016	0.019	0.015	0.013	0.011	0.012	0.008	0.002
Mg	2.079	2.093	2.119	2.204	2.420	2.297	2.277	2.339
Ca	1.813	1.810	1.814	1.805	1.795	1.788	1.788	1.823
Na	0.402	0.398	0.394	0.385	0.351	0.401	0.424	0.466
Κ	0.555	0.548	0.596	0.560	0.616	0.664	0.657	0.661

Table 5.4 Representative Electron Microprobe Analyses of Hornblende

All Fe reported as 2+

Cations calculated on the basis of 23 oxygens

* Composition used for thermobarometric analysis

	Zr-in-Rt			Zr-in-Ttn*			
Sample	<u>35b</u>	<u>41</u>	Sample	<u>43c</u>	<u>43e</u>	<u>43e</u> incl_in	
	matrix	matrix		matrix	matrix	cpx	
# of analyses	5	8	# of analyses	7	9	9	
ppm Zr ($\pm 1\sigma$)	367±25	355±20	ppm Zr (±1\sigma)	168±63	162±56	39±23	
T°C ($\pm 1\sigma$)	659±6	656±5	$T \circ C^{a}(\pm 1\sigma)$	708±21	707±18		
			$T \circ C^{b} (\pm 1\sigma)$	776±22	775±20	697±34	
			$T \circ C^{c} (\pm 1\sigma)$			718±35	

Table 5.5 Zr-in-Rutile and Zr-in-Titanite Results

*Activity of TiO₂ assumed to be 0.75 ^a Temperature calculated at 6 kbar ^b Temperature calculated at 12 kbar

^c Temperature calculated at 14 kbar

Table 5.6 Lu-Hf Isotopic Results for Garnet and Whole-rock Separates

Sample	Fraction	Lu	Hf	¹⁷⁶ Lu/ ¹⁷⁷ Hf	¹⁷⁶ Hf/ ¹⁷⁷ Hf	Initial ¹⁷⁶ Hf/ ¹⁷⁷ Hf ^a	Isochron
		(ppm)	(ppm)				Age (Ma)
AR01-2b	grt1	1.60	0.162	1.400	0.283291±13	0.282890±15	15.1±0.7
	grt2	1.60	0.142	1.593	0.283327±13		
	WR	0.160	0.779	0.0291	$0.282894{\pm}06$		
	*WR'	0.158	1.97	0.0114	0.282869 ± 06		
AR01-11b	grt1	1.87	0.115	2.312	0.283086±12	0.282480±11	14.1±0.4
	grt2	1.95	0.114	2.422	$0.283119{\pm}14$		
	WR	0.269	0.591	0.0647	$0.282494{\pm}05$		
	WR'	0.284	4.16	0.0097	$0.282486{\pm}05$		
AR01-35b	grt1	1.54	0.155	1.412	0.282635±14	0.282240 ± 48	13.9±2.5
	grt2	1.77	0.167	1.504	0.282635 ± 23		
	grt3	1.62	0.182	1.264	$0.282542{\pm}24$		
	WR	0.234	0.417	0.0795	$0.282263{\pm}10$		
	WR'	0.254	3.38	0.0107	$0.282252{\pm}06$		
AR01-41	grt1	2.33	0.147	2.244	0.282865±17	0.282270±86	14.5±2.7
	grt2	2.23	0.204	1.550	0.282716±11		
	grt3	3.54	0.234	2.149	$0.282847{\pm}11$		
	WR	0.085	0.506	0.0240	0.282271 ± 05		
	*WR'	0.104	3.43	0.0043	0.282069 ± 05		
AR01-43c	grt1	1.57	0.094	2.378	0.283787±18	$0.282880{\pm}11$	20.7±0.4
	grt2	1.57	0.104	2.140	$0.283708{\pm}15$		
	WR	0.158	0.738	0.0303	$0.282878{\pm}03$		
	WR'	0.177	1.95	0.0129	0.282889 ± 05		

WR, whole rock by savillex digestion; WR', whole rock by bomb digestion. reported errors are 2σ

* data not used in regression a calculated from isochron

CHAPTER 6: SUMMARY OF CONCLUSIONS

Several themes emerge from the data reported in this dissertation. The following outlines the contributions and prospective future research related to the questions posed in Chapter 1.

6.1 Annapurna

• Rocks in the Modi Khola valley display the apparent inverted metamorphic gradient associated with the Main Central Thrust. Temperatures increase structurally upsection, with a steep gradient at the Munsiari and Main Central Thrusts. Pressures also increase upsection, including all samples from the GHS, unlike at Langtang, 200 km to the east. With increasing structural level, peak P-T conditions were: 515 °C and 7.5 kbar (LHS below the Munsiari Thrust), 555 °C and 10.5 kbar (within the Munsiari Thrust sheet), 650 °C and 12.0 kbar (base of the Main Central Thrust Sheet), 735 °C and 12.5 kbar (Formation 1b), and 775 °C and 12.5 kbar (Formation 1c).

• Four generations of monazite have been identified in samples in central Nepal. From oldest to youngest they are: inherited, early prograde, late prograde, and post-anatectic, and can be distinguished chemically and chronologically.

• Peak metamorphic and post-anatectic monazite ages both progressively increase up-section. Monazite from GHS Formation 1a exhibits prograde

• metamorphic ages of 21-25 Ma. Formation 1b monazite contains prograde cores that are 24-29 Ma and high-Y rims that register cooling and melt crystallization ages of 17-22 Ma. In samples from Formation 1c, post-anatectic cooling and monazite growth occurred 19-15 Ma. In monazite from GHS Formation III, just below the South Tibetan Detachment, early prograde monazites are 35-38 Ma with later prograde growth at 30-35 Ma.

• Based on P-T gradients (increasing pressures up-section), we revise the original interpretation of the Bhanuwa Fault from a normal fault to a thrust, and identify a new thrust called the Sinuwa Fault structurally above the Bhanuwa thrust.

• Thrust rate calculations are generally consistent with a 2 cm/yr convergence rate component across the Himalaya from ~25 to ~15 Ma, supporting models that presume constant rates since at least 25 Ma, though the similarity in inferred peak ages for the ST, BT, and MCT could in principle indicate simultaneous initial cooling of all three sheets at ~22 Ma. A simultaneous cooling model could indicate as much as 100 km of out-of-sequence thrusting.

• Geochemically similar generations of monazite occur at both Langtang and Annapurna, but monazite ages along the Modi Khola appear older by ~5 Myr compared to Langtang, potentially indicating that the MCT at Annapurna is equivalent to the Langtang Thrust at Langtang (i.e., the thrust cuts up section from west to east). • The similarity between the chemical systematics of monazite, peak P-T conditions, and thrust rates calculated for Langtang and Annapurna implies that strain measurements in one part of an orogen can be realistically extrapolated to another within a few hundred kilometers.

6.2 Arun

• Garnet geochronology was used to provide the first direct measurement of the timing of eclogitization in the central Himalaya. Lu-Hf dates from garnet separates in one relict eclogite from the Arun River valley in eastern Nepal indicate an age of 20.7±0.4 Ma, significantly younger than ultra-high pressure eclogites from the western Himalaya, reflecting either different origins or substantial time lags in tectonics along strike.

• Four proximal garnet amphibolites from structurally lower horizons are 14-15 Ma, similar to post-eclogitization ages published for rocks along strike in southern Tibet.

• P-T calculations indicate three stages of metamorphism for the eclogite: i) eclogite-facies metamorphism at ~670 °C and \geq 15 kbar at 23-16 Ma; ii) a peak-T granulite event at ~780 °C and 12 kbar; iii) late-stage amphibolite-facies metamorphism at ~675°C and 6 kbar at ~14 Ma.

• Three models are considered to explain the observed P-T-t evolution: 1) deepening or longer transport of the MCT in the Arun area; 2) subduction, slab breakoff, and buoyant ascent of Indian crust occurred diachronously; and 3)

repeated loss of mantle lithosphere. These models either have difficulty explaining the P-T data, or require significant along-strike diachroneity and/or geographically restricted tectonic behavior at Arun.

• The channel flow and critical taper thermal-mechanical models of thrustbelt development in the Himalaya could in principle accommodate the presence of eclogites. However, neither model explains some key features of the Arun rocks. Channel flow models predict a focused erosional front for the last c. 25 Myr, yet the Arun rocks were exhumed by ~30 km between 14 and 25 Ma, apparently far from any erosional front. Critical taper has difficulty explaining P-T paths that combine exhumation with heating.

• A model that reconciles channel flow and critical taper with the eclogite P-T-t path suggests that as parts of the lower crust heat and weaken, they may be expelled to the mid-crust. An inverted geotherm is required to produce heating during expulsion, but this is predicted both in steady-state and time-dependent models.

• The P-T-t history of the central Himalayan eclogites may signal a change in the physical state of the Himalayan metamorphic wedge at 16-25 Ma.

6.3 Future Research

• Additional measurements of the chemical systematics of monazite, peak P-T conditions, and thrust rates from other parts of the orogen, e.g. in India or Bhutan, are needed to verify whether P-T-t results and strain estimates from Annapurna and Langtang could be applied to the orogen as a whole.

• Geochronologic and thermobarometric information reported here directly impacts our understanding of the stages of mountain building, especially with regard to the classic assumption of in-sequence thrusting and convergence of the Greater Himalayan sequence. Additional measurements of peak metamorphic and cooling ages from elsewhere in Nepal as well as other parts of the orogen are needed to assess the viability of assuming in-sequence thrusting as a part of tectonic models of Himalayan development.

• Eclogites may well be present at the surface in many locations in the central and eastern Himalaya, but have been overlooked due to the classic mistake of selective collection of rocks most appropriate for thermobarometric and geochronologic analaysis (e.g., metapelites). Thorough sampling and petrologic analysis is necessary to identify these rocks, which may have significant implications for Himalayan tectonic and thermal-mechanical models.

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